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Analysis of P- and S-Wave VSP Data from the Salton Sea Geothermal Field

T.M. Daley
(M.S. Thesis)

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# Analysis of $P-$ and $S$-Wave VSP Data from the Salton Sea Geothermal Field 

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## M.S. Thesis

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## CHAPTER 1

## INTRODUCTION

In March of 1986 the Salton Sea Scientific Drilling Project (SSSDP) completed drilling California State Well 2-14 to a depth of 10,500 feet. This deep well was drilled by the Department Of Energy (DOE) as part of the Geothermal Technology Development program with scientific studies in the well funded by the Basic Energy Sciences office of DOE, the National Science Foundation and the U.S. Geological Survey. The well is located near the Salton Sea Geothermal Field in southem Califomia. As part of the SSSDP, a vertical seismic profile (VSP) was conducted by the Lawrence Berkeley Laboratory's Earth Science Division. The processing and analysis of the Salton Sea VSP data is the basis of this thesis.

The objective of acquiring the VSP was to study the seismic wave propagation properties of the area near the well. There are several goals within this objective. The first goal is to use standard VSP analysis to obtain velocity models and search for reflective horizons within, and possibly beneath, the Salton Sea well. Another goal is the seismic detection of fracture zones surrounding the well including any indications of geothermal reservoirs. Also designed into the VSP is use of the shear wave sources to detect and analyze in-situ seismic anisotropy; if any, in rock surrounding the SSSDP. Finally, the SSSDP VSP data are used to test new particle motion analysis and display techniques for detecting fracturing and estimating the orientation of the fractures.

The usefulness of the VSP surveys has been well documented in the last decade; at least four books on VSP have been published and a number of new articles appear in the literature every year. It is assumed that the reader is somewhat familiar with VSP data, including the presence of upgoing and downgoing waves, tube waves, and the identification of reflector depth. For reference on these points any of several good review articles are available including Oristaglio, 1985 and Kennett, 1980. The application of VSP to anisotropy studies and fracture detection is more recent and is gaining in use. A 1987 meeting of the SEG included such titles as "Seismic Detection of Subsurface Fractures" ( Becker and Perelberg, 1987) and "Estimate of Shear Wave Anisotropy Using Multicomponent

Seismic Data" (Corrigan, et.al., 1987). Both of these papers, among others published in recent months, are looking at the use of VSP to study fracture-induced anisotropy. The VSP performed at the SSSDP provides a fairly complete data set taken in an area whose seismic propagation properties and fracture properties are not well understood. This paper is an attempt to apply current analysis techniques to understand an unusual geologic setting, rather than a proof of the techniques in an area whose seismic properties are known.

In order to understand any geophysical data, geologic information is necessary. For this reason, the thesis will begin with a summary of the geology of the Salton Trough region and the Salton Sea Geothermal Field (SSGF). The information available from the SSSDP will also be summarized, although much has not yet been published. After the geologic summary, the design of the VSP will be discussed, including acquisition equipment and procedures. The data processing procedures and software used will be discussed as a separate section. Processing procedures will also be described at various times in the thesis where more specialized procedures are used. Data analysis makes up the bulk of the thesis and it is divided into a number of sections detailing the basic VSP interpretation, the anisotropy analysis and the fracture detection and orientation analysis. A combined interpretation of the results, with probable geologic causes for observed events, is presented as a separate section from the data analysis. Finally, a summary of results for each of the goals stated above will be given. The reader should note that a large volume of data were collected and various display methods were used (from the standard wiggle-trace to three-component hodographs). Much of these data are left in the appendices with important or representative figures given in the body of the thesis. Also given in the appendices are listings of FORTRAN programs developed in conjunction with the thesis work.

## CHAPTER 2

## GEOLOGY

## 1. REGIONAL SETTING

The regional setting of the Salton Sea Geothermal Field (SSGF) is one of complex tectonic interactions. Figure 1 shows the location of the Salton Trough, a long, narrow depression extending northwest from the Gulf of California in Northern Mexico to the Coachella Valley region in Southern Califomia. The Salton Trough is a transitional zone on the boundary between the North American plate and the Pacific plate. The transition is from the transform faulting of the San Andreas fault system to the rifting associated with the East Pacific Rise. Both of these tectonic forces are currently active within the Salton Trough, but it is small spreading centers with their associated upwellings of mantle material which appear to be providing the geothermal heat source. The strike-slip faulting of the San Andreas system interacts with the rifting to create a complicated geologic province.

As an actively growing rift valley, the Salton Trough is subject to high rates of subsidence and sedimentation. The process of sedimentation has been ongoing since the Salton Trough first began developing in late Miocene or early Pliocene time (Crowell and Sylvester, 1979). Gravity and seismic refraction studies indicate the Trough contains over 6 km of sediments (Frith 1978, Randal 1974). This sedimentation has been dominated by the Colorado River, which enters in the center of the trough near Yuma, Califormia The Colorado River forms a delta which developed westward in midPleistocene time (Younker, et al, 1982). The delta has altemated flow northward and southward, separating the Imperial Valley to the north and the Mexicali Valley to the south. Southern flow, which is the current direction, discharges into the Gulf of California while northern flow discharges into the Imperial Valley and can only escape by evaporation. The current Salton Sea is one of many historic inland seas within the northern end of the Salton Trough.


Figure 1. Location of Salton Trough. Surrounding crystalline mountain ranges are stippled. Dotted lines denote approximate boundaries of East and West Mesas of Imperial Valley. From Sharp, 1982.

## 2. SALTON SEA GEOTHERMAL FIELD

The Salton Sea geothermal field is situated south of the Salton Sea in one of the young pullapart zones within the Salton Trough. The Salton Sea was formed between 1905 and 1907 when the Colorado River overflowed controlling levees and flowed north into the Salton Basin. The Salton Sea maintains its level because irrigation runoff from the Imperial Valley approximately matches the evaporation losses. Figure 2 shows the location of the SSGF as well as other geothermal fields within the Salton Trough. As an active rifting zone, the trough is an area of high regional heat flow. More than a dozen geothermal anomalies have been found in the Salton Trough, with the SSGF being one of the hottest. The SSGF is also one of the few geothermal areas of the trough showing surface manifestations of heat flow such as hot springs and fumaroles.

Drilling in the area of the SSGF began in 1927 when three wells were drilled to a maximum depth of 1473 feet (Lande and Elders, ed., 1979). These wells produced carbon dioxide commercially from shallow sands as part of the Imperial Carbon Dioxide field until production ended in 1954. The first geothermal production was from the Sinclair \#1 well drilled as an oil prospect to 4725 feet in 1957. By 1964, ten more geothermal weils were drilled in the immediate vicinity, beginning with the Sportsman \#1 (see Figure 3). All these well produced from hot water reservoirs with steam and brine mixed at the wellhead. The brine is highly mineralized with some wells having $300,000 \mathrm{ppm}$ total dissolved solids.

The sediments of the SSGF are considered a complete sequence dating from early Pliocene. While most of the sediments are detritus from the Colorado river, some erosional runoff is seen from the Chocolate Mountains on the east and the Peninsular ranges to the west. The deltaic deposits are interbedded sands, clays, silts and pebble conglomerates (Elders, ed 1979). These deposits are intermixed with lake sediments. The sedimentary section at the SSGF has been extensively altered by the hydrothermal process associated with the geothermal resource. Thermal springs at the surface provide evidence that, in certain locations, hot brines have penetrated the entire sedimentary column (Younker, et al, 1982). Temperatures up to 365 degrees centigrade have been recorded in the
geothermal wells of the SSGF.
The lateral structure of the SSGF is controlled by en echelon faulting with associated earthquake swarms (see Figure 3 for major fault locations). Sudies of well information have defined three general depth zones, shown in Figure 4, which characterize the geochermal field (Randal, 1974, Younker, et. al., 1982). At shallow depths there is a low permeability cap rock of variable thickness ( 750 to 2000 feet). The cap rock is composed of evaporite deposits with anhydrite and carbonate grains dominating. This surface zone serves as an impermeable layer to fluids and as an insulator against heat flow. Below the cap rock is the reservoir rock section. Younker et al., describe the transition from cap rock to reservoir rock as a "boundary between marine sediments (reservoir rock) deposited in the Gulf of Califormia and lacustrine sediments (cap rock) deposited in the Salton Trough after it was isolated from the southem portion of the basin in the mid-Pleistocene". The porosity and permeability of the reservoir rock is generally assumed to include secondary and fracture induced components (Younker et al. 1982, Muffler and White 1969). The porosity measured in five geothermal wells within the field was seen to decrease with depth from $25 \%$ to $5 \%$ over the 2000 to $6000^{\prime}$ range (Younker et. al. 1982).

Below the reservoir rock is a zone of major hydrothermal alteration, the altered reservoir rock. Within this zone the porosity and permeability decrease with depth. The transition from reservoir rock to altered reservoir rock is marked by a gradual loss of calcite and a corresponding increase of pore-filling epidote as an alteration product which can evolve into large crystals. These transitions mark the general structure of the SSGF. The structure is fairly well defined by well information within the center of the field. Near the edges of the field, the transitions between cap rock, reservoir rock and altered reservoir rock become more vague. Since the SSSDP is outside of previous drilling on the fringe of the field, a close correlation with these transitions was not expected. Figure 3 shows the location of the SSSDP with respect to other wells in the field.

Geophysical studies have been performed in the SSGF dating back to an early magnetic survey by Kelley and Soske in 1936. The results of these geophysical studies are well summarized by

Younker et al, 1982, who state the following:
(1) The field is associated with a local gravity high, probably resulting from intrusion of dike material, or of hydrothermal alteration of sediment, or both (Biehler et al., 1964).
(2) The field is associated with a magnetic anomaly that probably reflects the presence of igneous material near the surface (Griscom and Muffler, 1971; Kelley and Soske, 1936).
(3) A resistivity anomaly probably reflects the boundary of the saline brine. The conductance of the sedimentary sequence is greatest in the area of the drilled field, but a broad area of high conductance extends along the axis of the valley (Meidav et al., 1976; Kasameyer, 1976).
(4) Seismic refraction data reveal the presence of high-velocity material within 1 km of the surface (Frith, 1978).
(5) Numerous earthquakes indicate that the area is tectonically active (Schnapp and Fuis, 1977).
(6) Interpretations of resistivity surveys, seismic refraction data, earthquake locations, and ground magnetic surveys are suggestive of the existence of several steeply dipping faults within the field (Muffler and White, 1968: Babcock, 1971; Meidav and Furgerson, 1972; Towse, 1975; Gilpin and Lee, 1978; Frith, 1978).


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Figure 2. Geothermal fields of the Salton Trough. Arrow indicates approximate location of Salton Sea Scientific Drilling Project. Adapted from Elders and Cohen, 1983.


Figure 3. Location of wells and major faults within the Salton Sea Geothermal Field. Adapted from Younker, et al, 1980.


Figure 4. Stratigraphic zones of the Salton Sea Geothermal Field along an East-West cross section. The orientation of strata in the reservoir rock is shown by dashed lines. From Younker, et al, 1980.

## 3. SSSDP CORES AND GEOPHYSICAL WELL LOGS

The SSSDP VSP was performed after the well had been drilled to total depth. The preliminary results of mud logging, core analysis and well logs are therefore available for consultation. Figure 5 shows the stratigraphic column as interpreted from mud logging and core analysis. As expected, the well penetrated 10,500 feet of sediments consisting of sandstones, claystones, siltstones and some conglomerates. Igneous intrusives were first identified at 9500 feet. Figure 5 also shows where cores were taken and what hydrothermal alteration was noticed. Indications of alteration were actually observed in all the core sections taken, with the shallowest core (1553 feet) being described as "indurated mudstone". Major epidote alteration begins around 4500 feet. Zones of high porosity and permeability were easily identified by loss of drilling fluid circulation while drilling. These zones are identified in Table 1 (from Paillet ed, 1986).

The geophysical well logging included dual induction, natural gamma, spectral gamma, compensated neutron formation density, caliper, sonic, full waveform sonic and temperature. The quality of the logs obtained was generally poor with a number of logs being severly compromised by various problems including temperatures up to 300 degrees C , numerous zones of fluid loss, hole damage from cement plugs and mud coagulation and repeated trips in and out of the well. The sonic logs, which would normally be used in conjunction with the VSP velocity information, had serious problems with repeatability and cycle skipping and are therefore not incorporated with the VSP data. Recent studies comparing velocities derived from the full waveform sonic log with VSP derived velocities for $1500^{\prime}$ of the well do show good agreement (Cheng, C.H., 1987 personal communication). Further analysis and processing of the sonic logs may lead to useful results if a detailed velocity profile is required. For reflection seismology, the velocity information provided by the VSP should be sufficient. Use of density logs with velocity information can provide measures of the elastic constants of the material surrounding the well.


Figure 5. General Lithology and summary of the drilling phase of the SSSDP well. From Elders 1986.

TABLE 1 - LOST CIRCULATION ZONES (FROM PAILLET, 1986)
[ft, feet; bbl/h, berrels per hour; bbl, barrelyl

| Date | Depth (ft) | Fluid-lose (bbl/h) | Remarks |
| :---: | :---: | :---: | :---: |
| 11/11-12/85 | 3,107-3,200 | -15 | Only zone of fluid loss above $6,100 \mathrm{ft}$ |
| 12/22/85 | 6,119-6,133 | $-30 \times 0-100$ | Mineralized zone vith epidote fracture fill |
| 12/28-30-85 | 6,227 |  | Conducted first flow test |
| 1/5-15/86 | 6,637-6.889 | Total | 6,637 ft: added additives co drilling mud |
|  |  |  | 6.771 ft: ate cement plug |
|  |  |  | 6,850 ft: set second cement plug |
|  |  |  | 6,889 ft: injected cement under preseure. regained circulation |
| 1/16/86 | 7,030-7,090 | -85 | Added addicives Lo drilling mud |
| 1/27-29/86 | 8,095-8,160 | Total | Well flowing at depth of $8,126 \mathrm{ft}$; added combination of drilling-aud additives and cement: regeined circulacion |
| 2/1-3/86 | 8,580-8,800 | Total | Well flowing at depth of $8,580 \mathrm{ft}$; plugeed with combinacion of drillingaud additivee and cement |
| 2/4-5/86 | 8,800-8,920 | $-100$ | Looing fluid, yet drilling vith returne; vell flowed while tripping |
| 2/5/86 | 8,948-9,020 | Total | Plugged vith combination of drillingmud additives and cement |
| 2/7/86 | 9,098 | Tocal | Plugged with combination of drillingmud additives and cement |
| 2/10/86 | 9,254 | Total | Well flowing. gained 400 bbl |
| 2/11/86 | 9.450 | Toral | Plugged with combination of drilliagmud additives and cement |
| 2/19-21/86 | 9.450 |  | Cemented hole up to depth of $6,000 \mathrm{ft}$ in four atagea |
| 3/9/86 | $\begin{aligned} & 10,450- \\ & 10.460 \end{aligned}$ | -200 | Piugged with combination of drillingmud additives and cunenc by private concractor |
| 3/9/86 | 10.675 |  | Drilled without returns to total depth of 10.564 it |
| 3/17/86 | 10,564 |  | Conducted cecond flow rest |

## CHAPTER 3

## SURVEY DESIGN AND DATA ACQUISITION

The design of the Salton Sea VSP was based on the objective of studying the seismic wave propagation properties at the well site. The limiting factors were, as usual, time and money. A well is not under control when a VSP is being conducted and any open hole is in danger of being lost until circulation of drilling fluid resumes. The VSP survey can be stopped to allow circulation to be restored, but this is time consuming and therefore expensive. It was decided to limit the VSP to about 40 hours of data acquisition. This time limitation allowed two source locations, a near offset at $300^{\prime}$ and a far offset at $2300^{\prime}$. The VSP was also limited by high temperatures and unstable open hole conditions. These factors led to the decision to restrict the survey to the cased section of the well above 5750 feet To obtain complete seismic information, both a P-wave source and a shearwave source were used. The P-wave and shear-wave sources were Vibroseis (a trademark of CONOCO Inc.).

The goal of obtaining information about anisoropy and fracturing inferred from anisotropy led to the use of two orthogonal polarizations of shear waves. Two polarizations of shear waves were obtained with one vibrator by rotating it 90 degrees. The sources are labeled differently at the two offsets because of the nature of their expected particle motion (see Figure 6).

The sources labeled $S H$, and $S V$ are both generated by having the shear vibrator pad move radially toward and away from the well. The term $S H_{r}$ is used at the near offset and the term $S V$ is used at the far offset This terminology is derived from the geometry of the transverse particle motion expected for the first arrival in an isotropic media. The source labeled $S V$ at the far-offset gives an SV type of particle motion (i.e. transverse to the ray path in a vertical plane) because the raypath is not verical. In other words, since the wave propagation vector of the $S V$ source has a horizontal component, there is a vertical component of particle motion for the propagating shear wave. The source labeled $S H_{r}$ at the near-offset has a vertical raypath (and vertical wave propagation
vector), which means the particle motion is in the horizontal plane. The two near-offset shear sources then are $S H$ sources polarized 90 degrees apart, labeled $S H_{t}$ and $S H_{r}$.

The term $\mathrm{SH}_{\mathrm{t}}$ is used at both offsets for the source generated by having the shear vibrator pad move transverse to the radial direction. This source motion, in an isotropic material, gives a wave with particle motion transverse to the raypath in the horizontal plane. Figure 6 shows the direction of particle motion expected for each source in an isotropic medium.

The $S H_{s}$ and $S H_{r}$ source motions are the same as those labeled $S H$ and $S V$, respectively, in surface reflection work. It should be noted that this standard terminology of SH and SV shear waves is strictly valid only for transversely isotropic material with horizontal layering, otherwise the two wave types are not separable. Since the VSP data shows the surrounding material is not transversely isotropic, the simple model of separate SV and SH propagation does not hold. The actual shear-wave first arrivals will be seen to have a complex three-dimensional particle motion. The terms SH and SV will be taken to mean that part of the shear-wave particle motion in the horizontal plane or out of the plane, respectively.

The desire to study the three dimensional nature of the seismic wave field required the use of a three-component borehole geophone. A hydraulic wall-locking mechanism was used for maximum coupling with the well and the surrounding formations. The borehole tool was provided by Seismograph Service Corporation.

The vibrator source sweep was an 8 to 55 Hz upsweep, sixteen seconds long. The choice of sweep was govemed by time limitations, the desire to have as broad a sweep as possible while puting maximum energy in the more easily transmitted lower frequencies, and a desire to avoid 60 Hz noise. This 8-55 Hz sweep is also a standard surface seismic sweep which allows the VSP to be integrated with future reflection studies. The P-wave vibrator used for this survey was a Failing Y600 BD with 7.5 tons of output force and the Shear wave vibrator was a Mertz model 13 with 15 tons of output force. The vibrators were provided by Lawrence Berkeley Laboratory and the wireline service was provided by the USGS. The data were recorded on a Texas Instruments DFS-IV 24-channel seismic
data acquisition system. Only four channels were actually used, one for the source sweep and three for the borehole geophone's three components. The data were recorded uncorrelated at a two millisecond sample rate with a four second listening time following the 16 sec . sweep.

The choice of offsets was limited by the time available. A near-offset source was necessary to obtain accurate velocity information from vertical wave propagation. The near-offset data are in 50 foot depth intervals. While muluple far offsets were hoped for, only one far-offset data set was acquired because of time limitations and tool failure. 75 -foot depth intervals were used at the faroffset, giving less spatial sampling, but using less time. The near offset location was 300 feet from the well, and the far offset was 2300 feet from the well, both on approximately the same azimuth of S45E. Figure 7 shows a schematic of the offsets and a list of the depths surveyed. The offsets' azimuths correspond with the trend of the Salton Trough meaning the shear wave polarizations were approximately parallel and perpendicular to the axis of the trough.

The acquisition procedure was designed for the most efficient use of the available time. The survey was begun at the botum of the interval to be logged. This is done because the depth measurement is more accurate when the wireline is pulling a tool up the well instead of lowering it down. Starting at depth also reduces the impact of high temperatures since the tool is in the hottest part of the well first and as the well, which was cooled for the survey, begins to heat up, the tool is raised to cooler depths. After the tool is locked in place at a given depth, the wireline is slacked to prevent the resonance of a taught wireline. The downhole amplifier, which gives about 60 dB of gain, is then tumed on, providing amplification for the analog geophone output before the signal is sent over the wireline. The sources are alternated with one polarization of shear first, then the P source, then the other polarization of the shear source. By alternating the sources, the shear wave vibrator could be turned while the P wave vibrator is running its sweeps or while the tool is being moved up to the next level. Each source had three to ten sweeps stacked per level, depending on signal-to-noise ratio and on the time available.

The VSP was hampered by severe noise problems at the near-offset. The noise problem was worst between $2500^{\prime}$ and $4000^{\prime}$ and was worse on the horizontal components of the geophone than on the vertical. Possible causes of the noise will be discused in the velocity analysis section. High temperatures in the well stopped an effort to extend the survey to $7100^{\prime}$ for the far offset when the tool failed after two levels ( $7100^{\prime}$ and $7000^{\prime}$ ) were recorded with the $P$ source.


Figure 6. Diagram of expected particle motion for borehole geometry.

## VSP SOURCE LOCATIONS



## VSP GEOPHONE DEPTHS

NEAR OFFSET
P $500^{\circ}, 1000^{\prime}, 1500^{\circ}$
1900 - 5650

SHr 3050. 5650
1975-5650
SHt 500', 1000', 1500' $1900^{\circ}$-5650'

FAR OFFSET
P 1500', 1900 ${ }^{\circ}$ $1975^{\circ}$-5650 7000,7100

SV 1500, $1900^{\circ}$

SHt $\mathbf{1 5 0 0}^{\circ}, 1900^{\circ}$
$1975^{\circ}-5650^{\circ}$

Figure 7. VSP source locations and geophone recording depths.

## CHAPTER 4

## DATA PROCESSING

## 1. EDITING AND SORTING OF DATA

The goal of the editing and sorting process is to produce sections of each geophone component from each of the sources with the best signal to noise ratio possible. The editing and sorting was performed with Digicon's Interactive Seismic COmputer (DISCO) software. This software package is organized into subroutine modules which can be used in any order. Some further processing and display was performed with specially written FORTRAN routines which will be identified when they are discussed. All processing was done on a VAX 11/780 computer at the Center for Computational Seismology at the Lawrence Berkeley Laboratory.

The first step of processing was to demultiplex the data and extract the four desired channels from the 24 channels of data and 4 auxiliary channels which were recorded. This was done with DISCO's 'DEMUX' mutine. The next step was to edit noise out of the uncorrelated recordings. A major noise problem was spikes caused by movement of the borehole geophone while recording. If spikes are left in the data, the correlation process will produce noise at the correlation frequencies across the entire correlated output trace. Ideally, these spikes would be manually edited out with the minimum amount of trace removed. However, the large number of traces to be edited (over 2400 16 -second traces) encouraged the development of an automated editing process from various subroutines available in the DISCO package. After experimenting with various arrangements, final editing was done with the following software:
(1) An eight pole Butterworth bandpass filter (DISCO's 'FILTER' module) with 3 dB points at 8 and 55 Hz to match the sweep. This removed any noise outside the spectrum of interest.
(2) DISCO's 'SPIKEDT' routine which looks for noise bursts or spikes in the following manner: Two consecutive windows of user specified length are passed through the data, if the ratio of signal level in the leading window to signal level in the trailing window exceeds a user
specified value, then the leading window is zeroed. The windows are moved one sample at a time for the whole trace. This routine was used with 300 millisecond windows and a detection ratio of four.
(3) A second run of the 8 to 55 Hz Butterworth filter to remove discontinuities caused by the zeroed section of data, and prevent the associated correlation noise.

After this automatic editing, each trace was correlated with a source sweep recorded with every shot. Each correlated trace was visually inspected and either chosen for use in the stacked records or removed from further processing. At this point a few levels were found to have no usable data.

The sorting of the data was done by giving each trace a separate header value for source type and depth. The geophone components were automatically assigned a channel number by the demultiplexing software. The data were then sorted into groups with the same source and geophone component, in order of increasing depth. Once the data traces were sorted in this manner they could be stacked, with all traces having the same source, geophone component, and depth stacked together. At this point the final stacked sections of the original data were available, but full interpretation could not yet begin. A remaining obstacle was the random orientation of the geophone components.

## 2. ROTATION OF BOREHOLE GEOPHONE COMPONENTS

The data, as recorded in the field, consist of three components of ground motion recorded at each level in the well. These components are oriented orthogonally with one aligned on the tool's axis. When set vertically in a well, the tool has one vertical component and two horizontal components. The orientation of the horizontal components relative to an arbitrary frame of reference is, unfortunately, not known. There are three main reasons for the lack of orientation. The primary problem is the fact that a borehole geophone rotates around its axis on the wireline cable which supports it. This rotation, which occurs while the geophone is moved between levels, causes the horizontal phones to be randomly oriented at each level. A second cause of variation in geophone orientation is the deviation of the well itself, which rarely is vertical. Well deviation can cause a systematic
variation in the orientation of all three components as the tool is moved up the well. For wells with significant deviation, a deviation survey is usually run. This survey, if available, can be used to correct the geophone tilt. Thirdly, orientation error is caused by local variations in an uncased well. A severe washout can cause a tilt up to 10 or 15 degrees, depending on how the phone is locked to the wall. This survey was performed in an essentially undeviated, cased section of the well. However, until all borehole geophones are equipped with an orientation device, such as a gyroscope, the lack of known orientation will remain a source of uncertainty in interpretation.

There are various mathematical methods of determining the orientation of a three component geophone at each depth. In general, these methods depend on the assumption that the first arrival is a P-wave, which exhibits linear particle motion. The orientation of the three geophone components with respect to this particle motion can, in theory, be precisely determined. New data can be generated by projecting the particle motion onto any orthogonal coordinate system. This projection or 'rotation' of coordinates can be defined by two angles termed phi ( $\phi$ ) and theta ( $\theta$ ), and shown in Figure 8. Phi is the angle down from vertical and theta is the angle counter-clockwise around the vertical axis which is positive upwards. Once these two angles are known, the recorded data can be mathematically 'rotated' to recreate the signal which would have been recorded by geophones at any given orientation.

The 'wavefront' coordinate system which is most often used here is shown in Figure 9. This system has coordinates oriented in the directions that the P-wave, SH-wave and SV-wave first arrival's particle motion would occur for an isotropic material. Since the rotation is based on the angle at which the P -wave wavefront impinges on the geophone, and the curvature of the raypath is not known exactly, this wavefront coordinate system is not accurately defined with respect to the surface. If an assumption can be made that the vertical geophone component is truly vertical, and that the rays stay in one plane, then the wavefront rotation will give an absolute orientation with respect to the surface. For the Salton Sea well, which was cased and had deviation less than 10 degrees over the VSP intervals, the assumption of a true vertical component seems valid.

The method used to rotate the SSSDP data is eigenvalue analysis of the covariance matrix as described by Kanasewich (1981). The first step in this procedure is to pick the first P-wave arrival wavelet at each level. This is actually a non-trivial step since the wavelet may be distorted by signal-to-noise problems, and the particle motion is spread on 3 geophone components. As can be seen in Figures 10a and 10b, the vertical component gives good first arrivals for the near-offset data, but varying wavelet character for the far-offset data. The first arrival window for each level was picked manually with knowledge of the source wavelet and consistent decision making. Tests with varying windows show the method only varied five to ten degrees for any reasonable window.

Once the first arrival window has been picked, the next step is computing the covariance matrix. The covariance matrix is formed from N samples of the three components of data within the firstarrival window. The covariance matrix for three components, call them $\mathrm{Z}, \mathrm{X}$ and Y , is

$$
\left[\begin{array}{lll}
\Phi_{z Z} & \Phi_{Z x} & \Phi_{Z, y} \\
\Phi_{x,} & \Phi_{X x} & \Phi_{x, y} \\
\Phi_{Y Z} & \Phi_{Y X} & \Phi_{Y, Y}
\end{array}\right]
$$

where

$$
\Phi_{x_{1}, x_{2}}=\frac{1}{N} \sum_{i=1}^{N}\left(x_{1, i}-\mu_{1}\right)\left(x_{2, i}-\mu_{2}\right)
$$

and

$$
\mu_{1}=\frac{1}{N} \sum_{i=1}^{N} x_{1, i}=\text { mean of component } x_{1}
$$

for
N observations of component $x_{1, j} i=1,2,3, \ldots N$.

Once this matrix is computed, the eigenvalues and their associated eigenvectors can be found. The eigenvector of the largest eigenvalue is a vector which defines the best fit of linear particle motion within the coordinate system of the geophone components. A coordinate rotation is then defined by using this vector as one new coordinate (called 'Radial') and constraining one new
coordinate (called 'SH') to be perpendicular and made up of only the horizontal geophone components. This horizontal constraint assumes the tool is vertical and the 'horizontal' components are truly horizontal. Since the SSSDP well has little deviation, and the VSP was inside casing, this assumption should be quite good. If the well was deviated, a separate deviation survey would be required to properly orient the components. The third new coordinate axis (called 'SV') is simply orthogonal to the first two. The terms SH and SV are used because in an isotropic medium, the SH and SV sources would give first arrivals on these components only.

This eigenvector analysis gives the two rotation angles, $\phi$ and $\theta$, shown in Figure 8. The new data traces in the rotated coordinate system are computed by the following matrix operation.

$$
\left[\begin{array}{c}
R \\
S V \\
S H
\end{array}\right]=\left[\begin{array}{ccc}
\cos \phi & \cos \theta \sin \phi & \sin \phi \sin \theta \\
-\sin \phi & \cos \theta \cos \phi & \sin \theta \cos \phi \\
0 & -\sin \theta & \cos \theta
\end{array}\right]\left[\begin{array}{l}
Z \\
X \\
Y
\end{array}\right]
$$

The accuracy of the rotation is dependent on the quality of the P -wave first arrival. At some levels the signal-to-noise ratio was not good enough to allow a very accurate determination of the particle motion direction. A measure of the particle motion linearity, which is then also a measure of the accuracy of rotation, is the ratio of the largest eigenvalue $\lambda_{1}$ to the second largest eigenvalue $\lambda_{2}$. Kanasewich suggested forming the rectulinearity function

$$
F\left(\lambda_{1}, \lambda_{2}\right)=1-\left(\frac{\lambda_{2}}{\lambda_{1}}\right)
$$

which will be close to one when the particle motion is linear. This function served as a guide to which levels had noise problems. For some levels with severe noise or dead traces it became necessary to use a $\theta$ angle which was an interpolation of nearby levels. In most cases the error of rotation is probably between 5 and 10 degrees. The rotation angles computed for the near-offset P-wave source are given in Table 2, while the far-offset P-wave rotation angles are in Table 3.

For the near-offset data, the P-wave rotation was applied to all three sources. This used the assumption that the shear waves had essentially the same near vertical raypath as the P -waves. For the far-offset data rotation there was a problem. The raypaths were different for $P$ and $S$ waves. The
difference is clear because the shear-wave arrival does not show a continuous increase in travel time with increasing depth while the P -wave travel time does increase continuously with depth. This means that the shear-wave rays, particularly for the far-offset $S H_{t}$ source, turned up at some point and arrived from below the geophone. To accurately rotate the far-offset shear wave data, another rotation method was necessary.

A solution to this dilemma was suggested by the detection of very linear particle motion from the $S V$ source. The particle motion for the first arrival from the $S V$ source was analyzed after the rotation was applied, and it showed linear motion which was rotated off the new 'SV' component by an angle which varied with depth in a consistent manner (see Figure 11). The linear motion implied isotropic propagation, but in that case the first arrival motion should be aligned on the SV component. The angle between the SV component and the axis of the particle motion is interpreted as the error between the P-wave raypath and the shear wave raypath. This 'error' angle delta ( $\delta$ ), is shown in Figure 12. Delta can be found with a two dimensional. application of the covariance matrix algorithm to the R and SV components, finding the eigen vector of the larger eigen value of the matrix

$$
\left[\begin{array}{ll}
\Phi_{R, R} & \Phi_{R, S V} \\
\Phi_{S V R} & \Phi_{S V}, S V
\end{array}\right]
$$

Again, the accuracy of the analysis can be inferred from a rectilinearity function. Delta is then added to phi from the P -wave rotation and the shear wave data are rotated with this new phi angle. The horizontal 'SH' component is not changed since the raypath error is assumed to be in the Radial - SV plane. The error angles and rectilinearity function values for the far-offset are shown in Table 4.

Analysis of arrivals from $\mathrm{SH}_{\mathrm{t}}$ and SV sources showed a difference in their raypaths, but no method was known to recover this error angle. The $S H_{i}$ source was therefore rotated with the same angles as the $S V$ source and these angles are given in Table 5. Fortunately, the error between $S_{t}$ and $S V$ raypaths should be much less than the error between P and $S V$ sources since the error comes from differences in the waves' velocity gradients. The effect of the rotation is seen by comparing the plots in Figures 13a, 13b and 13c with those in Figures 33a, 33b and 33c. These figures show the
unrotated and rotated data, respectively, from the far-offset SV source. Note especially the difference between vertical and radial components.


Figure 8. Diagram of coordinate change for 3-component borehole geophone.

## ROTATED COORDINATE SYSTEMS



Figure 9. Rotated coordinate systems.


Figure 10a. Near-offset P-source, vertical component data.


Figure 10b. Far-offset P-source, vertical component data.

TABLE 2 -- NEAR-OFFSET P-WAVE ROTATION ANGLES

| DEPTH (FEET) | PHI | THETA | F |
| :---: | :---: | :---: | :---: |
| 593 | 66.579 | 147.279 | 0.908 |
| 1860 | 7.132 | -19.796 | 6.993 |
| 1668 | 11.104 | 84.536 | 0.981 |
| 2898 | 0.800 | 006.606 | 0.066 |
| 2060 | 4.638 | -85.856 | 0.994 |
| 2168 | 0.193 | -91.585 | 0.995 |
| 2160 | 6.725 | -82.885 | 0.997 |
| 2186 | 4.828 | -82.876 | 0.997 |
| 2258 | 5.246 | -80.691 | 6.998 |
| 2368 | 3.239 | -53.519 | 6.998 |
| 2350 | 4.363 | -24.881 | 0.975 |
| 2460 | 5.755 | -92.158 | 0.993 |
| 2469 | 3.954 | 64.110 | 0.997 |
| 2590 | 3.725 | 51.288 | 0.997 |
| 2550 | 2.339 | 115.934 | 0.999 |
| 2860 | 5.373 | -139.374 | 0.936 |
| 2880 | 5.833 | 15.498 | 0.739 |
| 2760 | 2.248 | 6.415 | 6.746 |
| 2756 | 7.623 | -134.746 | 0.989 |
| 2880 | 8.633 | -23.467 | 0.984 |
| 2856 | 1.886 | -78.867 | 0.788 |
| 2980 | 7.651 | -89.164 | 0.721 |
| 2960 | 9.361 | 136.676 | 0.910 |
| 30ed | 7.780 | 134.489 | 0.940 |
| 3050 | 9.553 | 44.519 | 0.972 |
| 318 | 3.763 | -99.973 | 0.658 |
| 3150 | 4.630 | 88. 489 | 0.987 |
| 32\% | 7.448 | -32.963 | 6. 824 |
| 3260 | 9.180 | 40.631 | 0.661 |
| 3360 | 1.008 | -92.885 | 0.857 |
| 3366 | 5.256 | -82.482 | 6.757 |
| 34E0 | 4.590 | 100.155 | 0.816 |
| 3458 | 2.759 | 131.692 | 0.989 |
| 3500 | 1.250 | 151.188 | 0.999 |
| 3656 | 0.447 | 79.846 | 6. 998 |
| 36e | 2.111 | 146.112 | 6.989 |
| 3856 | e. 083 | 90.898 | 0.984 |
| 376 | 6.803 | 0.000 | 0.090 |
| 3760 | 1.583 | 0.880 | 0.989 |
| 30e3 | -.984 | 46.791 | 0.985 |
| 3860 | 0.299 | 24.277 | 0.989 |
| 3900 | 7.284 | -178.843 | 6.989 |
| 396 | -. 635 | 157.863 | 0.997 |
| 40e6 | 2.323 | 177.621 | 0.942 |
| 406 | 5.381 | 21.838 | 0.928 |
| 410 | 4.234 | -7. 285 | 0.989 |
| 418 | 8.528 | 14.865 | 0.986 |
| 420 | 0.714 | -187.210 | 0.504 |
| 4260 | 11.493 | -10.787 | 0.988 |
| 430 | 14.384 | 3.896 | 0.981 |
| 4356 | -. 60 | 0.804 | -.806 |
| 446 | 13.562 | 7.044 | 0.986 |
| 4463 | 13.532 | -9.14 | 0.971 |
| 4508 | 17.084 | 6.318 | 6. 949 |
| 4580 | 14.606 | -8.894 | 0.982 |
| 460 | 19.387 | -34.698 | 0.974 |
| 4686 | 18.781 | -28.391 | 0.972 |
| 47 13 | 16.382 | -24.886 | 0.964 |
| 4780 | 22.597 | -29.523 | 6.968 |
| 4806 | 28.301 | -9.132 | 0.961 |
| 4880 | 25.567 | 35.616 | 0.894 |
| 4904 | 17.533 | -6. 851 | 0.931 |
| 4986 | 9.800 | 06.868 | 6. 680 |
| 60es | 19.983 | -5.831 | 6.898 |
| 5660 | 15.563 | 12.542 | 0.781 |
| 618 | 11.113 | 33.339 | 0. 942 |
| 5150 | 7.975 | -98.006 | 0.793 |
| 520 | 18.394 | 64.233 | 0.598 |
| 5280 | 14.594 | 83.094 | 0.875 |
| 5300 | 7.543 | 134.879 | 0.933 |
| 536 | 13.116 | 58.841 | 0.894 |
| 54E3 | 7.288 | 20.488 | 6.988 |
| 5456 | 10.110 | 47.862 | 0.949 |
| SEC | 12.675 | -82.128 | 6.292 |
| 585 | 13.911 | 98.207 | 0.780 |
| 56e | 14.032 | 92.482 | 0.925 |
| 5685 | 5.785 | 185.426 | 6.981 |

TABLE 3 - FAR-OFFSET P-WAVE ROTATION ANGLES

| DEPTH (FEET) | PHI | THETA | F |
| :---: | :---: | :---: | :---: |
| 1500 | 77.888 | -84.603 | 0.971 |
| 1900 | 78.278 | -77.235 | 0.982 |
| 1975 | 89.265 | -101.295 | 0.972 |
| 2056 | 71.454 | -122.580 | g. 982 |
| 2125 | 78.887 | -47.499 | 0.931 |
| 2268. | 76.827 | -41.888 | 6. 918 |
| 2275 | 74.773 | -149.349 | 0.837 |
| 2350 | 89.564 | -145.779 | 0.888 |
| 2425 | 68.698 | -156.367 | 0.933 |
| 2560 | 88.895 | -118.90] | 0.974 |
| 2575 | 77.160 | -148.188 | 0.978 |
| 2856 | 77.498 | -138.849 | 0.946 |
| 2725 | 89.418 | -33.774 | 0.882 |
| 2888 | 66.892 | -185.177 | 0.789 |
| 2875 | 65.453 | -109.951 | 0.776 |
| 2950 | 75.975 | -152.222 | 0.989 |
| 3025 | 78.982 | -188.935 | 0.983 |
| 3166 | 76.566 | -175.589 | 0.983 |
| 3175 | 86.226 | -18.439 | 8. 987 |
| 3250 | 84.517 | -6.165 | 0.889 |
| 3325 | 74.117 | -4.679 | 0.867 |
| 3468 | 72.433 | -9.740 | 0.946 |
| 3475 | 87.139 | -8.939 | 0.973 |
| 3550 | 62.494 | -16.725 | 0.969 |
| 3625 | 59.161 | -12.732 | 0.977 |
| 3760 | 81.511 | -15.914 | 0.977 |
| 3775 | 54.916 | -9.147 | 0.974 |
| 3858 | 58.338 | -9.631 | 0.976 |
| 3925 | 52.985 | -12.871 | 6.939 |
| 4808 | 52.736 | -18.683 | 8.935 |
| 4075 | 48.538 | -29.886 | 0.964 |
| 4150 | 48.367 | -27.724 | 0.886 |
| 4225 | 46.892 | -19.999 | 0.877 |
| 4360 | . 49.731 | -31.431 | 0.932 |
| 4375 | 40.992 | -19.644 | 0.841 |
| 4456 | 41.891 | -22.343 | 0.977 |
| 4526 | 45.464 | -29.440 | 0.975 |
| 4806 | 44.585 | -29.695 | 0.993 |
| 4675 | 37.396 | -31.343 | 0.938 |
| 4750 | 44.978 | -19.141 | 8.953 |
| 4825 | 53.679 | -23.694 | 0.795 |
| 4968 | 41.781 | -14.542 | 0.978 |
| 4975 | 3.288 | -90.600 | 0.638 |
| 5058 | 43.262 | -10.219 | 0.899 |
| 5125 | 48.632 | -5.581 | 8.973 |
| 5298 | 24.916 | -32.841 | 0.808 |
| 5275 | 32.724 | 7.285 | 0.982 |
| 5360 | 38.775 | 2.109 | 0.946 |
| 5425 | 14.342 | 28.346 | 6.936 |
| 5560 | 29.684 | 4.289 | 8.972 |
| 5575 | 18.568 | 22.289 | 0.997 |
| 5858 | 0.060 | 0.068 | 0.000 |

DEPTY

1900.
$3625^{\circ}$

$1975^{\circ}$
$5050^{\circ}$


Figure 11. Particle Motion of far-offset SV source data. Radial component is on vertical axis and SV component is on horizontal axis.

## ROTATION ANGLE CHANGE FOR SV RAY PATH



Figure 12. Rotation error angle $\delta$ seen between $P$ raypath and SV raypath.

TABLE 4 - ERROR ANGLES FOR FAR-OFFSET ROTATION

| DEPTH | (FEET) | DELTA | $F$ |
| :---: | :---: | :---: | :---: |
| 1500 |  | -24.271 | 6.9589 |
| 1908 |  | -22.461 | 6.9759 |
| 1975 |  | -22.223 | 6.9813 |
| 2656 |  | -21.592 | 0.9833 |
| 2125 |  | -28.719 | 6.9784 |
| 2200 |  | -19.378 | 0.9874 |
| 2275 |  | -19.862 | 0.9886 |
| 2350 |  | -19.823 | 0.9852 |
| 2425 |  | -19.239 | 0.9839 |
| 2568 |  | -19.933 | 6.9583 |
| 2575 |  | -26.025 | 6.9441 |
| 2650 |  | -28.129 | 0.9314 |
| 2725 |  | -28.138 | 0.9288 |
| 2866 |  | -19.464 | 0.9259 |
| 2875 |  | -19.461 | 6.9258 |
| 2956 |  | -19.283 | 0.9218 |
| 3625 |  | -19.621 | 6.9188 |
| 3108 |  | -18.831 | 8.9176 |
| 3175 |  | -19.028 | 0.9178 |
| 3256 |  | -19.108 | -.9178 |
| 3325 |  | -19.889 | 0.9181 |
| 3486 |  | -19.656 | 0.9181 |
| 3475 |  | -18.951 | 8.9175 |
| 3566 |  | -18.842 | 6.9171 |
| 3625 |  | -18.780 | 6.9167 |
| 3768 |  | -18.588 | 8.9168 |
| 3775 |  | -18.462 | 0.9188 |
| 3856 |  | -18.328 | 6.9184 |
| 3925 |  | -18.224 | 6.9183 |
| 4065 |  | -18.126 | 6.9182 |
| 4675 | - | -18.643 | 0.9168 |
| 4150 |  | -17.969 | 6.9154 |
| 4225 |  | -17.866 | 0.9149 |
| 4366 |  | -17.777 | 0.9144 |
| 4375 |  | -17.691 | 0.9139 |
| 4456 |  | -17.893 |  |
| 4625 |  | -17.546 | 0.9129 |
| 4806 |  | -17.418 | 6.9124 |
| 4875 |  | -17.332 | 6.9126 |
| 4756 |  | -17.238 | 0.9115 |
| 4825 |  | -17.146 | 6.8188 |
| 4966 |  | -17.841 | 6.9161 |
| 4975 |  | -16.973 | \%.96937 |
| 6856 |  | -16.886 | 0.9687 |
| 6125 |  | -16.888 | 6.9681 |
| 5288 |  | -16.741 | 0.8676 |
| 5276 |  | -16.659 | 0.9685 |
| 5350 |  | -18.681 | 0.9686 |
| 5425 |  | -18.518 | 6.9864 |
| 5560 |  | -16.446 | . 808 |

TABLE 5 - FAR-OFFSET SHEAR-WAVE ROTATION ANGLES

| DEPTH (FEET) | PHI | THETA |
| :---: | :---: | :---: |
| 1508 | 101.68 | -84.063 |
| 1900 | 92.278 | -77.235 |
| 1975 | 91.295 | -161.265 |
| 2056 | 93.454 | -122.589 |
| 2125 | 96.887 | -47.499 |
| 2280 | 89.827 | -41.886 |
| 2275 | 93.773 | -149.349 |
| 2350 | 89.584 | -146.779 |
| 2425 | 87.898 | -156.367 |
| 2500 | 88.895 | -118.990 |
| 2575 | 97.166 | -146.188 |
| 2860 | 97.496 | -138.849 |
| 2725 | 89.418 | -33.774 |
| 2860 | 85.892 | -165.177 |
| 2875 | 85.463 | -109.961 |
| 2950 | 95.975 | -152.222 |
| 3625 | 96.982 | -168.935 |
| 3180 | 95.568 | -175.689 |
| 3175 | 85.228 | -18.439 |
| 3250 | 83.517 | -6.165 |
| 3325 | 93.117 | -4.679 |
| 3489 | 91.433 | -9.746 |
| 3475 | 86.138 | -8.939 |
| 3550 | 80.494 | -18.725 |
| 3625 | 78.181 | -12.732 |
| 3708 | 88.511 | -15.914 |
| 3775 | 73.916 | -9.147 |
| 3856 | 74.336 | -9.031 |
| 3925 | 71.965 | -12.871 |
| 4006 | 71.736 | -16.583 |
| 4675 | 87.538 | -29.680 |
| 4150 | 66.367 | -27.724 |
| 4225 | 84.692 | -19.999 |
| 4360 | 67.731 | -31.431 |
| 4375 | 59.992 | -19.544 |
| 4450 | 59.891 | -22.343 |
| 4525 | 63.404 | -29.440 |
| 4608 | 82.585 | -29.695 |
| 4875 | 54.398 | -31.343 |
| 4750 | 62.970 | -19.141 |
| 4826 | 76.979 | -23.694 |
| 4966 | 58.781 | -14.542 |
| 4975 | 59.288 | -98.606 |
| 5650 | 80.262 | -18.219 |
| 6125 | 57.632 | -6.561 |
| 5208 | 41.916 | -32.841 |
| 5275 | 48.724 | 7.285 |
| 5350 | 47.775 | 2.108 |
| 5425 | 31.342 | 28.346 |
| 5500 | 46.684 | 4.288 |



Figure 13a. Far-offset SV source, vertical component data.

## FRR OFFSET SV SOURCE <br> HORIZONTRL COMPONENT 1



Figure 13b. Far -offset sV source, horizontal component 1.


Figure 13c. Far-offset $s V$ source, horizontal component 2.

## 3. OTHER PROCESSING PACKAGES USED

After the three components were rotated into a consistent coordinate system, more standard processing packages could be used to aid interpretation. One important tool for VSP interpretation is separation of upgoing and downgoing waves. This separation is possible because the vertical receiver geometry gives upgoing waves and downgoing waves opposite apparent velocities across the section. Apparent velocity is seen as moveout or 'dip' of an arrival on the VSP section, and events with differing velocities such as P-waves, shear waves and tube waves will exhibit different dips. A useful tool is some type of dip filter which will attenuate energy with a given dip anywhere on the section. A number of different methods have been proposed and studied for dipfiltering of VSP sections. Two DISCO modules were available for use. One was a user-written module 'DIPFIL', which is a timedistance domain filter. The other is DISCO'S 'COHERE' module which is a frequency-wavenumber domain filter. Tests of these two filters on the same data led to use of the DIPFL module because of noise generated by COHERE. The filter was applied with a 4 ms per trace slope and used a 25 Hz center frequency. The procedure used for dipfiltering included aligning the first arrivals by time shifting all the traces so their first arrival was at the same time. This means the downgoing energy would have zero dip across the section. This improved the filtering process since it was found that the filter works best on events with zero dip. The traces could be shifted back after filtering to preserve the correct time relations. The traces are left aligned when estimating the depth of reflections below the deepest level.

Other processing used standard seismic techniques such as automatic gain control (AGC), bandpass filters (the previously described Butterworth filter with adjustable bandwidth and rolloff), and trace equalization for ploting. Unless stated otherwise, all plots have individual trace normalization using the maximum amplitude on the trace. The particle motion analysis ploting routines used in the anisotropy studies were from specially writuen FORTRAN codes (HODOS.FOR in Appendix 4) used on the VAX $11 / 780$ with a laser writer for output plots.

## CHAPTER 5

## DATA ANALYSIS

## 1. VELOCITY CALCULATION, NOISE PROBLEMS AND ACCURACY

The first step in analyzing the VSP data was calculation of the vertical velocity structure from the travel time of the near-offset first arrivals. The first arrival time was picked on the first peak of the P-wave wavelet and the first trough of the shear-wave wavelet. The shear-wave arrival was picked differently because its trough had consistently better signal-to-noise ratio than its peak. The signal-to-noise ratio of the first arrivals determine the accuracy of the velocity measurements. The quality of near-offset data ranges from excellent for the $P$-wave source to very poor for sections of both shear sources. Many of the shear-wave arrivals were not coherent enough to pick travel times (see Figure 14). A first arrival pick was only made at those depths where the first arrival wavelet had a consistent character.

The P-wave travel time was picked from the rotated radial component, the SH , travel time was picked from the rotated SH component, and the $S H_{r}$ travel time was picked from the rotated SV component. These components were used to measure the velocity of those arrivals having particle motion oriented the same as the source with respect to the raypath. Because of the near vertical raypath, the SV component of near-offset recordings is virtually in the horizontal plane, rotated 90 degrees from the SH component. When calculating velocities, a straight raypath assumption is used to correct the measured depth for the extra propagation distance due to the $300^{\prime}$ source offset. The travel time information and average velocity calculations for $P, S H_{s}$ and $S H_{r}$ sources are shown in Tables 1-1 and 1-2 of Appendix 1.

Interval velocities were calculated from travel times between depth levels. If the level at the top of the interval was not pickable, the next shallowest level with a pickable first amival was used. Therefore, at levels where signals were too noisy to pick, the interval velocity is averaged over a larger interval. The interval depth listed is the center of the interval regardless of interval distance. The interval velocities are listed for $100^{\prime}$ and $500^{\circ}$ intervals in Tables $1-3$ and $1-4$ of Appendix 1.

Using velocities for P and S -waves, the $\mathrm{P} / \mathrm{S}$ ratio and Poisson's ratio ( $\sigma$ ) could be computed. Poisson's ratio is calculated using the relation

$$
\sigma=1 / \frac{\alpha^{2}-2 \beta^{2}}{\alpha^{2}-\beta^{2}}
$$

where $\alpha$ is the $P$-wave velocity and $\beta$ is the shear-wave velocity. Computations were made using both $S H_{t}$ and $S H_{r}$ sources, although the $S H_{r}$ source has fewer measurements. Interval ratios were computed with the same algorithm as the interval velocities. Table 1-5 and 1-6 in Appendix 1 give, respectively, the average and interval calculations for the $\mathrm{P} / \mathrm{S}$ ratios and Poisson's ratio. The interval P/S ratios and Poisson's ratio calculations are given for 500 -foot intervals. Figure 1-1 of Appendix 1 shows the interval Poisson's ratios for the $S H$, source data. The FORTRAN program written to perform velocity and Poisson's ratio calculations, VEL.FOR, is listed in Appendix 4.

Analysis of the velocity data should include consideration of the data quality and the dependability of measurements. Since the P-wave arrivals have uniformly excellent signal-to-noise ratios, the P-wave interval velocities can be considered very precise. The error of the first break pick should be within $\pm 2 \mathrm{~ms}$. The shear-wave data quality is variable (see Figure 14). The $\mathrm{SH}_{\text {, }}$ arrivals are good from surface to $2550^{\prime}$; from $2550^{\prime}$ to $4000^{\prime}$ they are very poor and only a few arrivals were picked. The SH, first arrivals are poor from 3050', the shallowest level, to $4200^{\prime}$; from 4250' to $5650^{\prime}$ the first arrivals are good, with only a few unpickable levels in this zone. Because of the lower frequency content of the shear waves, the first arrival pick is less precise than P-wave arrivals with the same signal-to-noise ratio. The estimated shear-wave travel time error is $\pm 5 \mathrm{~ms}$.

The zone of noisy shear-wave arrivals is caused by large amplitude noise on the borehole tool's horizontal components. In general, the horizontal geophones within a borehole tool are more susceptible to noise vibrations than the vertical component because the weight of the tool tends to wedge it vertically in the well. Also, if the locking-arm is not fully set, the tool may allow horizontal vibrations; if vertical motion is allowed, the tool will tend to slide down the well. The horizontal components of the near-offset P-wave data, which were recorded without relocking the tool, have the same noisy zone.

There does appear to be other noise problems in the SSSDP well. The cement bond log shows a deterioration in the cement bond below 2500'. A poor bond will reduce the signal-to-noise ratio for all the data. A different cause of the noise zone is indicated by the difference in data quality between near-offset and far-offset data. The far-offset shear-wave first arrivals have better data quality at the same depths (see Figure 15). A possible factor is a lower number of sweeps run at each depth for the near-offset sources. There are also indications that a tube wave created low frequency noise coincident with the near-offset shear-wave first arrivals. Figure 16 shows how an event with tube wave velocity (T4), arrives at the same time as the noise begins. Whatever its cause, the poor signal-tonoise ratio of the near-offset shear-wave arrivals reduces the scope and accuracy of data analysis.


Figure 14. Comparison of near-offset $S H_{t}$ source, SH component data (top) with near-offset SH $_{r}$ source, SV component data (bottom).


Figure 15. Comparison of near-offset $S H_{l}$ source, SH component data (top) with far-offset $S H_{1}$ source, SH component data (bottom).


Figure 16. Near-offset $P$ source, vertical component data showing possible generation of noise from a tube wave.

## 2. VELOCITY MODELING AND RAY TRACING

After computing the interval velocities for both P and S waves, a model could be developed for use in ray tracing. It was desirable to smooth the calculated interval velocities in order to reduce the number of velocity layers used for computation by the ray tracing software. Using the velocities from $100^{\prime}$ intervals, zones where the velocity did not show sharp discontinuities were averaged, and models for P and S waves were developed. The S -wave model came from the $S H_{\mathrm{t}}$ velocities. Figure 17a shows the computed interval velocities at $100^{\prime}$ intervals and Figure 17 b shows the velocity models.

The simplified velocity models were used in ray tracing software to help understand the propagation of the observed arrivals. The ray tracing, shown in Figures 18 a and 18 b , used an algorithm which propagated straight-rays within each constant velocity layer. The first observation is that the near-offset data do represent vertical propagation for all but the shallowest levels. The far-offset data, however, have strongly bent raypaths caused by the velocity variation with depth. The ray tracing for the far-offset shear waves (Figure 18b) shows discrete changes in ray path for the shallow velocity layers. A change in ray path is the most likely explanation of the change in wavelet character seen at 2275' for the far-offset $S H_{\text {t }}$ source. The far-offset shear waves actually have a turning point and arrive at some depths from below the geophone. This effect is seen on the rotated first arrivals for far-offset $S H_{t}$ source in Figure 31a which have their earliest arrival at approximately $2800^{\circ}$.


Figure 17a. Interval velocity measurements for $\mathbf{P}$ and SH waves.


Figure 17b. Velocity models used for ray tracing.


Figure 18a. P-wave ray tracing for near-offset source (top) and far-offset source (bottom).


Figure 18b. Shear-wave ray tracing for near-offset source (top) and far-offset source (bottom).

## 3. EVENT IDENTIFICATION - REFLECTED AND TRANSMITTED ENERGY

### 3.1. NEAR-OFFSET P-SOURCE DATA

The first data traces to be analyzed for reflected energy were from the near-offset $P$-wave survey. Figures 19a, 19b and 19c shows the 3 rotated components. The SH and SV components show shear arrivals between 1 and 2 seconds and some $P$ energy between .25 and .75 seconds. The radial component shows the P -wave first arrivals, but their large amplitude has left the rest of the trace uninterpretable because of the trace normalization used. Figure 20a shows a plot of the vertical component of the near-offset $P$ source with AGC. An AGC was applied to enhance the later events and the vertical component was used because P reflections should be traveling nearly vertically. The first arrival is clearly visible as are the numerous downgoing multiples. With horizontal layering, the downgoing multiples will parallel the first arrivals (Balch and Lee 1984, Hardage 1985).

Also identifiable on this section are two twbe waves (labeled T1 and T2), the second of which may be a multiple of the first. A tube wave is a low frequency, dispersive, Stoneley wave which propagates in the borehole fluid and decays exponentially away from the well (Toksoz and Stewart, 1984). The tube wave's average velocity was computed to be $4800 \mathrm{ft} / \mathrm{sec}$, slower than $P$ or shear waves. The zero frequency velocity of the tube wave $\left(V_{T}\right)$ is related to the fluid velocity ( $\alpha_{1}$ ) and fluid density ( $\rho_{1}$ ), and the surrounding rock's shear velocity $\left(\beta_{2}\right)$ and density $\left(\rho_{2}\right)$ by the following relation (Hardage, 1985):

$$
V_{T}=\left[\frac{1}{\alpha_{1}^{2}}+\frac{\rho_{1}}{\rho_{2} \beta_{2}^{2}}\right]^{-1 / 2}
$$

Tube waves can be analyzed for the information they provide through velocity, attenuation and particle motion as well as depth of generation (Toksoz and Stewart 1984, Biot 1952, Cheng and Toksoz 1981, Hardage 1985, Huang and Hunter 1981). The strength of the tube waves observed in the SSSDP VSP indicates further analysis could be profitable, however a thorough study of the tube waves is outside the limits of this paper.

Other identifiable events on Figure 20a include a shear-wave arrival which seems to have an associated reflection. There are suggestions of P-wave reflections, but they are mostly obscured by the downgoing energy. The enhancement of upgoing reflected energy is the next step.

To emphasize the reflected energy, a dipfilter was applied which attenuated downgoing events and therefore enhanced the upgoing events. The dipfiltered vertical component section is shown in Figure 20b. Here the previously identified events stand out, but also seen clearly is reflected energy, labeled event E , from below the deepest level at 5650 . Event E is clearly the strongest reflection and it is impossible to say if the later events are multiples or separate events, but a strong reflection is likely to have associated multiples. As can be seen, the central section of the data, from 3000' to 4000', is apparently blank. This effect is caused by the high amplitude noise associated with the shear arrival between 1250 ms and 1500 ms . The traces are normalized to the highest amplitude so the earlier, lower ampliude arrivals are not clearly visible. Fortunately, the noise is dominandy low frequency, so a high pass filter will reduce the noise. The highpass filtered section is shown in Figure $20 c$.

- An estimate can be made of the depth at which a reflection is being generated. The procedure is to project the event down the section until it intersects a first arrival (the first arrivals can be aligned at one time, or not aligned, but the analysis is easier if they are aligned). The depth at which the projected reflection event intersects the first arrivals is the depth of generation (Hardage, 1985). If the reflection is from below the deepest level, this procedure assumes a constant velocity between the deepest level and the depth of refiection generation. Figure 21 shows the results of this interpretation. Event $E$ has an estimated depth of generation of $7000^{\prime}$, the later event $F$ is estimated at $7600^{\circ}$. Event D is not a very coherent reflection because it only appears strongly on four traces, however it does have the moveout of a reflection and its depth of generation is estimated at $6650^{\circ}$.

The depth of generation of the tube wave can be identified in the same manner, except it is projected up the well. The estimated depth of tube wave T1 generation is $1800^{\prime}$. The second tube wave, T2, is estimated to be generated $500^{\prime}$ above the surface, meaning the estimate is wrong by $500^{\prime}$ and it
then may be generated by surface waves incident on the borehole, or it is a multiple of the first tube wave. Another tube wave, T3, is apparently being generated by the shear wave at about 3000' although this may be a simple tube wave multiple.

Other identifiable events include a shear reflection, event $G$, being generated at approximately 4100'. This reflection appears to be a mode-converted P-wave. The evidence for this interpretation is the apparent velocity of the reflection. The estimated velocity of the reflection between $2000^{\prime}$ and $3000^{\prime}$ is $10,500 \mathrm{ft} / \mathrm{sec}$ while the measured P -wave velocity in this zone is $9,700 \mathrm{ft} / \mathrm{sec}$ and the shear wave velocity is $6000 \mathrm{ft} / \mathrm{sec}$. Allowing for error in estimating the reflection velocity, event $G$ appears to be P-wave energy. Also seen on this section are P reflections generated at approximately 2900', 4400', and 5450', labeled R2, R3 and R4 respectively.


Figure 19a. Near-offset $\mathbf{P}$ source, radial component data.

## NEAR OFFSET P SOURCE <br> SV COMPONENT



Figure 19b. Near-offset $P$ source, SV component data.


Figure 19c. Near-offset $\mathbf{P}$ source, SH component data.


Figure 20a. Near-offset $P$ source, vertical component data with AGC.


Figure 20b. Near-offset $P$ source, vertical component data, dipfiltered.


Figure 20c. Near-offset $P$ source, vertical component data, dipfiltered with 25 Hz high pass filter.


Figure 21. Identification of events on near-offset $\mathbf{P}$ source, vertical component data with dipfilter, high pass and aligned first arrivals.

### 3.2. NEAR-OFFSET SHt SOURCE DATA

The near-offset shear surveys were, as discussed, hampered by severe noise problems. The traces between $2500^{\prime}$ and $4000^{\prime}$ are essentially useless. This gap in the data made the dipfiltering process noisy. The best sections to interpret were simply high pass filtered at 25 Hz with an AGC applied. This processing is shown for all three components in Figures 22a, 22b and 22c. The SH component data (Fig. 22a) shows two strong multiples, labeled M1 and M2 following the primary shear event. Since these multiples are not evenly spaced in time, they are probably being generated at different depths, but the data are not complete enough or quiet enough to identify their source depth. The SV component data (Fig. 22b) shows strong shear-wave arrivals and some multiples as well as indications of tube wave energy, event T2. The implications of strong first arrivals on the SV component will be discussed in conjunction with anisotropy detection. The radial component data (Fig. 22c) show a P-wave event, probably generated at the surface by the vibrator. Also seen is the tube wave event (T2) identified from the P-wave source. Weak indications of a reflector, labeled R1, can be seen on the radial component. The depth of generation of R1 is estimated at $4100^{\prime}$, but accurate determination is not possible.

### 3.3. NEAR-OFFSET SH ${ }_{r}$ SOURCE DATA

The $S H_{r}$ source suffered from the same poor signal- to-noise ratio as the $S H_{1}$ source as can be seen in Figures 23a, 23b and 23c which show the three components of the SH $_{r}$ survey. Again, the dipfiltering process was not found to provide interpretable events, so the sections shown are high pass filtered at 25 Hz ( with a 6 pole Butterworth filter) and displayed with AGC applied. The first shear arrival is good from $4150^{\prime}$ to $5650^{\circ}$ on the SV and SH components (Figs. 23a and 23b). The radial and SV components (Figs. 23c and 23a) show tube wave energy (event T2), probably generated by surface waves at the top of the well's fluid column. The radial component also has converted P-wave arrivals with associated multiples. Shear-wave multiples can be seen on the SV and radial components, although they are not as strong as those from the $\mathrm{SH}_{\mathrm{r}}$ source.


Figure 22a. Near-offset $S H_{\text {, }}$ source data, SH component.


Figure 22b. Near-offset $S H_{\text {; }}$ source data, SV component.


Figure 22c. Near-offset $S H_{H}$, source data, radial component.


Figure 23a. Near-offset $S H_{r}$ source data, SV component.


Figure 23b. Near-offset $S H_{r}$, source data, SH component.


Figure 23c. Near-offset $S H_{r}$, source data, radial component.

### 3.4. FAR-OFFSET P SOURCE DATA

The data traces recorded by the far-offset $P$ source are probably the most interesting and unusual in the survey. Figures $24 \mathrm{a}, 24 \mathrm{~b}$ and 24 c show the data from the rotated components. While the radial component shows good P-wave first arrivals, it is the SV component which is quite anomalous. A number of unexpected downgoing events, including the two labeled $A$ and $B$, and one strong reflection event labeled C , dominate the first second of data on the SV component. If these events were reflections or multiples from horizontal layers they would have appeared on the nearoffset data sets. Keeping in mind the near-offset P-wave survey, which did not show such strong events on any component, it is clear that a simple horizontal layer model could not explain these events. Also identifiable is a source-generated shear arrival at 1.6 seconds, identified because the faroffset shear sources' first arrivals have the same travel time.

The downgoing events seen on the SV component are not simple downgoing multiples of the P-wave since they do not parallel the downgoing first arrivals. Their apparent velocity, or moveout across the section, would be the same as the first arrivals if they were generated by a horizontal bed. If they were generated by a dipping bed, they would show a different moveout but they would still be observed on the near-offset survey. It was the analysis of the apparent velocity of these events which provided a key to understanding their generation.

The downgoing events $A$ and $B$ and the reflection, $C$, have the same moveout as the near-offset P-wave first arrivals. This observation implies these events are vertically travelling P-waves. The reason they are seen strongest on the SV component is explained by looking at Table 3 which has the rotation angle $\phi$ for the far-offset $P$ source. For the first 30 levels, where these events are seen, phi is between 50 and 85 degrees. This means the radial component is close to true horizontal and the SV component is seeing most of the vertical particle motion. If these events are vertically travelling P . waves, whether upgoing or downgoing, they would have a vertical particle motion. In order to best see a vertically polarized wave, the original unrotated vertical geophone component should be used.

Figure 25 shows the original vertical component. Events A, B and C are still coherent even though the direct P ray has oblique incidence and the first arrival is weak on many traces. The event labeled F is more coherent here than on the SV component (Fig. 24b). Given the observations, the best interpretation of the events labeled A, B, C, and F is that they are caused by P-wave energy which has been scattered near the well. The depth of generation can be estimated by looking at events B and C which appear to be upgoing and downgoing waves from the same source. They have opposite polarity and can be traced back to a common point in time at the 3000' depth where they no longer are seen on the section. The time at which they appear is after the first arrival, implying they were generated at some distance from the well, causing a delay before they arrived above and below the depth of generation. These events are then explained as being generated at a depth of approximately $3000^{\circ}$, and at some offset from the well. Figure 26 shows a schematic of this proposed scattering.

A likely cause of P-wave scattering, given the known geologic environment, is a localized zone of fracturing or perhaps a vertical boundary between open fractured rock and impermeable sealed rock. This explanation is suggested by suidies of cores taken between 3012' and 3020'. These were found to have fractures which are "presently open and permeable" in a matrix of "chloritized and epidotized crossbedded sandstone and shales" (McKibben and Andes, 1986). This depth is also the only zone above $6000^{\circ}$ which had fluid loss during drilling and was considered a possible flow zone. In fact, it was thought that a reservoir might be found near this depth because of projections made from wells in the SSGF, but the low rate of fluid loss and low permeability led to a decision to not test this zone (Sass, 1987, personal communication).

In order to look for other reflections, the vertical component section was dipfiltered to enhance upgoing energy. Figure 27 shows the dipfilered vertical component. Notice that while the first arrivals are attenuated, the downgoing scattered energy (events A and B) is still apparent. Since the dipfiltering process removes any downgoing energy caused by interbed multiples from horizontal reflectors - these events have the same apparent dip as the first arrivals - the indication is again that the downgoing energy associated with events $A$ and $B$ is scattered and not reflected. A reflection
from below the data set can also be seen on the dipfiltered section: In fact three reflections can be identified. Figure 28 shows the estimated depth of these reflections at $6100^{\prime}, 6900^{\prime}$, and $7900^{\prime}$. Again these are estimates which assume a constant velocity below the deepest level and the later events may be multiples. The later upgoing events (after 1.1 seconds) actually appear to be primary reflections of the scattered energy as well as multiple reflections. There is also an indication of a reflection from approximately $2800^{\prime}$ labeled event $D$. The different moveout between events D and C illustrates the difference between scattered and reflected energy.

In order to improve the resolution and separation of events. the dipfiltered vertical section was filtered with a 25 Hz high pass Butterworth filter. This filtered section, shown in Figure 29, shows the pairs of upgoing and downgoing events centered about 2950'. Events $A$ and $A^{\prime}$ are the first pair which appear to be generated very near the well. In fact, the first wavelets for $2875^{\prime}$ and 2950 ' show a reversal in polarity. The other labeled events, $B, B \prime, F$ and $F$ ', all appear to be generated between 2800' and 3100'. There are also events following the shear wave arrival which appear to be scattered from the same depth zone.

To confirm the scattering model, the apparent moveout of events A and A' was computed for comparison with the vertical velocity. Event A has an apparent velocity of $13,000 \mathrm{ft} / \mathrm{sec}$ between $3025^{\prime}$ and $3925^{\prime}$ while the vertical velocity is $11,800 \mathrm{ft} / \mathrm{sec}$ and the far-offset first arrival's moveout is $19,200 \mathrm{ft} / \mathrm{sec}$. These velocities show that event $\mathbf{A}$ is related to the vertical velocity; its faster moveout velocity can be explained by allowing for some horizontal offset of the scattering source. Event $A^{\prime}$ has an apparent velocity of $9600 \mathrm{ft} / \mathrm{sec}$ between $2050^{\circ}$ and $2950^{\circ}$, while the vertical velocity is 9000 $\mathrm{ft} / \mathrm{sec}$ and the far-offset first arrival's moveout is $29,000 \mathrm{ft} / \mathrm{sec}$. Event $A^{\prime}$ is apparently vertical propagation and not an interbed multiple which would have a moveout close to the first arrivals.

The radial component in Figure 30a, shows good first arrivals which have an unexpected double peak character. The expected wavelet is the same as the near-offset P source (Figure 19a). The two deepest traces, which are check shots from $7000^{\prime}$ and $7100^{\circ}$, show the expected wavelet without the double peak. The double peak wavelet seen on the radial component is probably caused by a
vertically traveling P-wave following the direct arrival close enough that their wavelets merge. At the $7000^{\prime}$ depth the two events have either separated enough so that there is no interference, or the secondary event has attenuated enough to reduce its effect. A more complete data set is necessary to understand the change between $5650^{\circ}$ and $7000^{\circ}$. A 25 Hz high pass filter was applied to the radial component to increase the resolution and see if a second event was interfering with the direct arrival. This section, Figure 30b, does indeed show a second event following the first arrival between 3175' and $4675^{\prime}$. It is the low frequency components of these two events which are overlapping and causing the double peak character seen in Figure 30a.

The SH component (Figure 24c) shows some of the characteristics of the SV component but it is more noisy. The downgoing scattered events can be seen, implying they have a horizontal component of motion. The shear arrival at 1600 ms shows a reflection, event H , whose depth of generation is estimated at 2875 ', possibly the same zone which is generating the scatured energy.

## FAR OFFSET P SOURCE ROTATED RADIAL COMPONENT



Figure 24a. Far-offset $\mathbf{P}$ source, radial component data.


Figure 24b. Far-offset $\mathbf{P}$ source, SV component data.

## FAR OFFSET P SOURCE <br> ROTATED SH COMPONENT



Figure 24c. Far-offset $\mathbf{P}$ source, $\mathbf{S H}$ component data.


Figure 25. Far-offset $\mathbf{P}$ source, vertical component data.

## MODEL OF SCATTERED P-WAVES



Figure 26. Model of scattered P-waves.


Figure 27. Far-offset $\mathbf{P}$ source, vertical component data, dipfiltered.

## FAR OFFSET P SOURCE ALIGNED DOWNGOING



Figure 28. Reflector depth estimation for far-offset $\mathbf{P}$ source, vertical component data, dipfiltered.


Figure 29. Far-offset $\mathbf{P}$ source, vertical component data, dipfiltered with $\mathbf{2 5 ~ H z}$ high pass filter.


Figure 30a. Far-offset $\mathbf{P}$ source, radial component data.


Figure 30b. Far-offset $P$ source, radial component data with 25 Hz high pass filter.

### 3.5. FAR-OFFSET SH, SOURCE DATA

Figures 31a, 31b and 31c show the rotated sections for the far-offset $S H_{\mathrm{t}}$ source. The shear wave arrival can be seen on all three components between 1.5 and 2.0 seconds. Between $2500^{\prime}$ and 3500' there is very little moveout of the first arrival because the rays are tuming and the angle of shear-wave incidence is near 90 degrees in this zone (see Table 5). The presence of a strong first arrival on both the SH and SV components is evidence of shear wave splitting, which will be discussed with the anisotropy effects. There is a discrete change in the character and time of the first arrival between $2200^{\circ}$ and $2350^{\prime}$ which is seen on the SH component (Figure 31a). This is probably an effect of raypath since the ray tracing (Figure 18b) shows a change in ray path near this depth with shadow zones between 2000' and $3000^{\circ}$.

The wavelet changes again at 3100' with a downgoing event emerging with different moveout than the direct shear arrival (event A). This event shows the same apparent velocity as the scattered P-wave energy seen for the far-offset P source. Apparently, the zone at $3000^{\circ}$ is scattering P waves from incident shear waves. These scattered waves are seen more clearly on the SV component (Figure 31b), which has nearly vertical orientation.

Both the SH and SV components show downgoing multiples following the first arrivals. The radial component shows a poor first arrival, as expected, but there is low frequency downgoing energy below 3000' which is probably also associated with the P-wave scattering. An interesting effect is seen on the shallow traces at about 1.3 seconds where a coherent event is arriving before the first shear anrival. A possible interpretation is that this is also P-wave energy which has been scattered at some distance from the well, possibly by fracture zones, and has moved out ahead of the shear wave. Also seen on the radial component is a reflection from approximately $2500^{\prime}$ (event J).

Deep reflection events are somewhat visible on the radial component. Inspection of the various components shows these reflections are strongest on the original vertical component. An estimate of the depth of reflection was made using the vertical component. Figure 32 shows an aligned dipfiltered section of the vertical component. Estimates were made for three events at $6800^{\circ}, 7500^{\circ}$ and $8600^{\circ}$.

Later events can be seen, but they are probably multiples. In fact, the $8600^{\prime}$ could easily be a multiple between the $6800^{\circ}$ and $7500^{\prime}$ reflectors, only its strength indicates it is a separate event.

Figure 31a. Far-offset $S H_{\text {t }}$ source, SH component data.


Figure 31b. Far-offset $S H_{t}$ source, SV component data.


Figure 31c. Far-offiset $S H$, source, radial component data.

```
FAR OFFSET SH SOURCE ALIGNED DOWNGOING
```



Figure 32. Estimated reflector depth from Far-offset $S_{4}$, source, vertical component data with dipfilter and aligned first arrivals.

### 3.6. FAR-OFFSET SV SOURCE DATA

Figures 33a, 33b and 33c show the rotated component plots for the far-offset $S V$ source. The first thing to notice is the strong first arrival and good signal to noise ratio on the SV component. Also, the SV component shows the first arrival having nearly simultaneous arrivals between 2000' and 2500', showing again the upturning of the ray path due to a steep velocity gradient. A possible reflection can be identified on the deep traces. The SH component shows the shear wave arrival but it is noiser with a high frequency event labeled A causing some interference. The event A may be scattered P energy propagating with a horizontal component.

The radial component shows some energy with the main shear arrival, and it also has two separate events before the shear arrival (labeled K and L ). These events are probably scattered P waves generated away from the well which have moved out in front of the shear waves. An estimate can be made of the distance from the well at which these scattered $P$ waves were generated by using the known P and S velocities and the travel time difference. By assuming a straight ray path one can use the relation,

$$
X=\frac{\Delta T V_{P} V_{S}}{V_{P}-V_{S}}
$$

where

$$
\begin{aligned}
& \mathrm{X}=\text { Distance of scatterer from well } \\
& \Delta V=\text { Travel ime difference } \\
& V_{P}=\text { P-wave average velocity } \\
& V_{S}=S \text {-wave average velocity }
\end{aligned}
$$

Using this equation, the distance from the well to the scatter source was estimated at $1600^{\prime}$ for event K , which would be roughly half the distance from the source to the $2000^{\circ}$ depth receiver. Given the oblique incidence of the raypaths at this depth, it seems reasonable to assume the scatterer is mostly offset horizontally, and the lack of moveout of the scattered event also suggests horizontal
offset. A separate scattered event is seen deeper on the radial component (event L). The same analysis gives a distance to this scatterer of $1500^{\prime}$ from the receiver at $3000^{\prime}$, which is probably also a dominantly horizontal offset.

The radial component of the SV source shows a strong reflection from below the deepest geophone depth. This reflection actually is stronger on the original vertical component, so this component is used for depth estimation. Figure 34 shows an aligned, dipfiltered section of the vertical component. Two reflector depths are estimated at $6800^{\circ}$ and $8000^{\circ}$. Later upgoing events can be seen, but these are probably multiples.


Figure 33a. Far-offset $S V$ source, SV component data.


Figure 33b. Far-offset $S V$ source, SH component data.


Figure 33c. Far-offset SV source, radial component data.

## FAR OFFSET SV SOURCE

ALIGNED DOWNGOING


Figure 34. Estimate of reflector depth for far-offset SV source, vertical component data with dipfilter and aligned first arrivals.

## CHAPTER 6

## ANISOTROPY ANALYSIS

## 1. VELOCITY ANISOTROPY - NEAR-OFFSET DATA

As stated before, the near-offset shear sources are two polarizations of SH waves, labeled $\mathrm{SH}_{r}$ and $\mathrm{SH}_{\mathrm{t}}$. With these two polarizations of SH waves, it becomes possible to look for anisotropy in the horizontal plane, information rarely measurable in-situ. Two separate methods were used to look for anisotropy. The first is a simple travel time measurement used to look for velocity differences between orthogonal polarizations. This was done by picking first arrival times for both sources on the receiver component with the same polarization as the source, and comparing those times at depths where both were pickable. The first arrival was picked on the first trough for both sources. Figure 35 shows a plot of the travel time difference. The data available, from $4250^{\prime}$ to $5650^{\prime}$, shows that $S H_{r}$, is faster with an apparent trend of increasing separation between $S H_{i}$ and $S H_{r}$. The average travel time difference in this depth range is 9 msec which is approximately $0.5 \%$ velocity anisorropy. While this is not a large amount of anisotropy, it is consistent enough over this depth interval of 1400 feet to show the presence of an azimuthal polarization dependence of shear wave velocity in the horizontal layers of this area.

## 2. VELOCITY ANISOTROPY - FAR-OFFSET DATA

The far-offset SV source generates particle motion in the vertical plane. In fact, the oblique incidence of the raypaths indicate most of the motion is in the vertical plane. By comparing the data from the $S H_{1}$ source with the SV source, any velocity difference between SH and SV polarizations can be detected. As with the near-offset data, the first step is to look at simple travel time differences between $S H_{\text {t }}$ and $S V$ sources. Again the travel time for each source was picked on the rotated coordinate on which the wave would arrive in an isotropic material, i.e. the $S H_{t}$ was picked on the SH component and the SV source was picked on the SV component. Figure 36 shows the travel time
difference at each depth where both sources could be picked, points are to the right of the zero line when $S H_{t}$ is faster and to the left when $S V$ is faster. The error of any one point is estimated at $\pm 5$ ms.

The dominant feature is a trend of SV motion becoming increasingly faster than SH motion. Another surprising feature is the crossover in travel time difference with the SV source having longer travel times shallow and $S H_{\mathrm{t}}$ source having longer travel times at depth. For the shallowest levels, from $1500^{\prime}$ to $2100^{\prime}$, the $S H_{\text {t }}$ source becomes increasingly faster, but at $2275^{\prime}$ there is a discrete jump to equal travel time. This jump corresponds to the changes in wavelet character, raypath and travel time for the $S H$, source data. From 2275' to $3550^{\prime}$ the travel times are approximately equal with the exception of an anomalous zone around $3250^{\prime}$. In this zone the $S H_{t}$ motion is again faster. From 3250' to $5500^{\prime}$ the $S V$ motion becomes increasingly faster than the $S H$, motion. The total velocity difference of 16 ms . at $5500^{\circ}$ represents about $1 \%$ velocity anisotropy. As with the $0.5 \%$ horizontal plane anisotropy seen at the near-offset, this is a small but measurable amount. The overall trend of increasing separation between SH and SV is very evident.


Figure 35. Travel time difference for near-offset shear sources. Data points are $S H$, time minus $S H_{\text {, }}$ time.


Figure 36. Travel time difference for far-offset shear sources. Data points are $S V$ time minus $S H_{t}$ time.

## 3. SHEAR-WAVE SPLITTING

It is the observation of particle motion which provides the second method of analyzing anisotropy. Given the observed velocity anisotropy, shear-wave splitting should be seen on plots of the first shear arrival's particle motion. The use of particle motion plots to sudy anisotropy effects has been well documented (Crampin 1985, Majer, et. al. 1986, etc). A three component geophone provides the ability to look at the particle motion within any 2-D plane or in its complete 3-D motion. Figure 37 illustrates the splitting of shear wave particle motion when a shear wave enters an anisotropic region. As described by Crampin (1985):

A shear wave entering an anisotropic region necessarily splits into the two or more fixed polarizations which can propagate in the particular ray direction. These split phases propagate with different polarizations and different velocities, and on reentering an isotropic region the original waveform can not be reconstructed.

Figure 38 shows particle motion within the first arrival window for all three sources from the far offset at the 1900 ' level. These plots show the three 2-D slices which describe the complete particle motion plotted on the faces of a cube whose axes are the three components. The vertical axis is the radial component while the horizontal axes are the SH and SV components. The three plots on the cube faces are Radial vs SH, Radial vs SV and SH vs SV. This means the pure shear motion, SH vs SV, is shown on the bouom of the cube. While the SH vs SV plot is often the only one displayed in other studies, the radial component can provide much information about the three-dimensional nature of particle motion. It would be possible to plot the particle motion as a single threedimensional pach, but in practice this type of plot is difficult to interpret since it is significantly dependent on the angle at which it is viewed. Figure 38 is representative of the particle motion one would expect from isotropic propagation. The P-source first arrival has mostly radial motion, the SH -source arrival has mostly SH motion and the SV-source arrival has mostly SV motion.

The window used for these hodographs is approximately 70 milliseconds, although each arrival has a slighty different window length. The points plotted are two milliseconds apart. The first three
points in time are circled to provide the correct identification of first motion. The software used for this display was a FORTRAN program, HODOS.FOR listed in Appendix 4. Since these plots are sensitive to signal-to-noise problems at any one level, their analysis should avoid individual levels which do not show correlation with nearby levels. Very good data quality is necessary before interpreting a single anomalous particle motion plot. The levels which were obviously distorted by noise are not displayed.

Appendix 2 contains the particle motion plots for all the sources and all depths with sufficient signal-to-noise ratio. A quick inspection shows the P-wave source arrivals are, as theory predicts, confined to the radial component. It is the particle motion of the shear arrivals which contains information about anisotropic propagation along the raypath. The following Tables summarize the important characteristics of the shear-wave particle motion plots displayed in Appendix 2, emphasizing both the SH-SV plane motion and the amount of radial motion.

## MODEL OF SHEAR-WAVE SPLITIING



Figure 37. Model of shear-wave splitting in an anisotropic media. After Crampin, 1985.

## THREE COMPONENT PARTICLE MOTION

1900 FT


Figure 38. Three component particle motion for far-offset $\mathbf{P}, \mathbf{S H}$, and $\mathbf{S V}$ sources at 1900' depth level.

## 4. NEAR-OFFSET PARTICLE MOTION DESCRIPTIONS

4.1. $S H_{1}$ SOURCE

|  | NEAR-OFFSET SH, SOURCE |
| :--- | :--- |
| DEPTH (FEET) | DESCRIPTION |
| 500 | The SH-SV motion is fairly linear with dominant SH orientation. A <br> strong radial component is seen in the Radial vs SH plane. |
| $1000-2200$ | The first motion is SH, but the arrival then forms an ellipse rotated <br> approximately 45 degrees from the SH axis. |
| $2300-2500$ | The first motion is generally SH, but the dominant direction is SV <br> with the motion more circular at 2450' and 2500 . Very litule radial <br> motion is observed. |
| $2550-4000$ | These traces are too noisy to use particle motion analysis. |
| $4050-4350$ | The motion is elliptical in the SV-SH plane with a dominant SH <br> direction and no radial motion. |
| $4400-4700$ | The SV and SH motion begins to be out of phase with each other <br> and the amount of SV motion at each level decreases until at 4700 <br> there is linear SH particle motion. |
| $54500-5450$ | The shear motion is again elliptical with a dominant SH motion and <br> little radial motion. |
|  | The SH and SV motions are about equal, though they are out of <br> phase giving odd shapes. There is some radial motion at these levels <br> and the first motion is always SH. |

### 4.2. NEAR-OFFSET SH, SOURCE

| NEAR-OFFSET SH, SOURCE |  |
| :--- | :--- |
| DEPTH (FEET) | DESCRIPTION |
| $3050-4200$ | These traces are too noisy for particle motion analysis. |
| $4250-4350$ | The first motion is in the SV direction. At 4250' the SH motion is <br> as large as the SV motion. The amount of SH motion decreases to a <br> minimum at 4350'. |
| $4400-4550$ | The SH motion increases to give circular polarization at 4550' with <br> some radial energy. |
| $4600-4800$ | The motion is split into two orthogonal polarizations, both rotated <br> approximately 45 degrees from the SH and SV axis. |
| $4850-5200$ | The two polarizations are approximately aligned with the SH and SV <br> axis. |
| $5350-5650$ | The motion takes on a figure 8 pattern with dominant polarization <br> which rotates from aligned on SH to aligned on SV. |

## 5. FAR-OFFSET PARTICLE MOTION DESCRIPTIONS

5.1. $\mathrm{SH}_{\mathrm{I}}$ SOURCE

|  | FAR-OFFSET SH, SOURCE |
| :--- | :--- |
| DEPTH (FEET) | DESCRIPTION |
| 1500 | The main arrival has very linear SH motion. Late in the arrival there <br> is more elliptical motion. |
| $1900-2200$ | Very linear SH motion with litule radial motion. |
| $2350-2875$ | The motion develops a significant SV component of motion. SH <br> dominant elliptical motion continues to 2875' with radial motion at <br> 2725 and 2800'. |
| $2950-3100$ | The motion is nearly circular with all three components nearly equal <br> in amplitude. |
| $3175-3500$ | The radial motion decreases and the ellipse becomes more linear in <br> the shear plane, with an SH direction. |
| $3550-3775$ | The shear motion ellipse broadens as more SV motion is seen. |
| $3925-4225$ | The first motion is SH, but the wavelet takes on SV and radial <br> motion, and then retums to SH motion. |
| $4300-4675$ | There is no radial motion, and the motion is elliptical with SV orien- <br> tation. <br> re |
| $4750-5275$ | The motion is still elliptical, but the direction of the orientation is <br> rotated about 45 degrees from the axis and more radial motion is <br> seen. |
| $5350-5425$ | The shear motion is roughly a figure 8 aligned mainly with SH and <br> with the later part having large radial motion. |
|  | The shear motion is mainly aligned on the SH axis with strong SV <br> and radial components. |

### 5.2. FAR-OFFSET SV SOURCE

| FAR-OFFSET SV SOURCE |  |
| :--- | :--- |
| DEPTH (FEET) | DESCRIPTION |
| $1500-2050$ | The shear motion is a fairly linear ellipse dominant SV motion. |
| $2125-2500$ | The ellipse begins to rotate off of the SV axis and develop into two <br> polarizations which are about 45 degrees apart. The later polariza- <br> tion has some radial motion. |
| $2500-2800$ | These levels are too noisy for particle motion analysis. |
| $2800-5500$ | The shear motion for all these depths is mostly linear SV motion. <br> There is significant radial motion around 3025', 4525' and between <br> 4900' and 5050'. |

## CHAPTER 7

## INTERPRETATION OF ANISOTROPY ANALYSIS AND EVENT IDENTIFICATIONS

The interpretation of the various observations made with the VSP data is somewhat difficult because seismic data interpretation needs good geologic control, especially with the more experimental techniques such as particle motion analysis. Currently, only preliminary findings about the SSSDP are available. The available geologic information is shown in Figure 5 which has the fractional parts of sandstone, siltstone and claystone with the amount of alteration from sulfides, anhydrite, chlorite and epidote. Gaps such as the one from $6700^{\prime}$ to $6900^{\prime}$ are due to lost circulation which prevented mud logging and core samples. The interpretation of the seismic data will attempt to link the anisotropy analysis and the event identification. Among the factors which may affect the VSP data are the local stratigraphy, fracturing with or without fluid, and regional tectonic stress. Particular attention is paid to indications of fracture zones since this is one of the goals of the VSP. The interpretation separates the discussion of the horizontal plane anisotropy data from the $S V$ vs $S H_{t}$ anisotropy data.

## 1. HORIZONTAL PLANE ANISOTROPY - NEAR-OFFSET DATA

As discussed previously, the near-offset shear wave data represent a horizontal plane anisotropy experiment. Figure 39 shows a summary of the near-offset anisotropy data. Unfortunately, the data are incomplete since the $S H_{r}$, source has no shallow data. The left column of Figure 39 shows the velocity information, the center and right hand columns shows the generalized particle motion description for first arrivals from the $S H_{t}$ source and $S H_{r}$ source, respectively.

There is a problem with interpreting near-offset particle mocion analysis. The determination of the horizontal rotation angle theta is inaccurate for vertically incident waves. Therefore, the absolute orientation of the rotated SH and SV components may vary with depth. However, since the components are still orthogonal, non-linear motion implies there is shear wave splitting. Only the absolute orientation of the motion is unknown. For instance, the variation in ellipse orientation seen on $\mathrm{SH}_{\text {t }}$ source plots between $2000^{\prime}$ and $2050^{\prime}$ in Appendix 2 for may be due to incorrect rotation, but the
non-linear motion still indicates anisotropy.

Between the surface and $2500^{\circ}$, the elliptical nature of the particle motion plots shows shearwave splitting within the horizontal plane. The observed spliting must be caused by horizontal plane anisotropy. A possible cause of this anisotropy is depositional alignment of mineral grains due to the dominate Northwestern flow of sediments from the Colorado river to the Salton Sea and its predecessors. Another possible cause is variation of the stress field within the Salton Trough caused by tectonic forces.

Below 4000' more information is available. The travel time information, which begins at 4250', shows the $S H_{r}$-generated wave is slower than the $S H_{r}$-generated wave. The particle motion plots for both sources show shear-wave spliting with two dominant polarizations of motion which appear to change phase with each other. A change in shape is interpreted as a phase change. The phase changes are probably due to the varying difference in travel time of each polarization. The changes in orientation of a given shape may be the rotation error.

## NEAR OFFSET ANISOTROPY SUMMARY



Figure 39. Near-offset anisotropy data summary.

## 2. SV VS SH, ANISOTROPY - FAR-OFFSET DATA

The far-offset anisotropy data provide a much more complete data set. The results are outlined in Figure 40 which gives the travel time and particle motion summaries. The shallow zone above $1500^{\prime}$ appears to be fairly isotropic with equal travel time to $1500^{\prime}$ and linear particle motion for both sources' arrivals at 1500 '. From 1900' to $2200^{\prime}$ the $S H_{\mathrm{t}}$-generated waves are increasingly faster than waves from the SV source. Between $2200^{\prime}$ and $2500^{\prime}$ the particle motion shows evidence of anisotropy, with the first shear arrival from the $S H_{1}$ source developing SV motion and the SV source first arrival spliting into two polarizations. The travel time and particle motion information both indicate a change near 2200'. Velocity increases are seen for both $P$ and $S$ waves near 2200 ' with the $S$ velocity showing a large increase. There are indications of reflections near 2200', but identification was not possible because of the lack of data between 1500' and 1900'. There does appears to be a general transition zone around $2200^{\prime}$. Geologically, from $1700^{\prime}$ to $2100^{\prime}$ there is the first anhydrite alteration, possibly marking the "cap rock" seen in the SSGF. The seismic effects may be related to the transition from anhydrite alteration in a mostly claystone layer to an alteration free sandstone layer at 2200'.

From 2200' to 3000' the $S V$ - and $S H_{1}$-generated arrivals have approximately equal travel time indicating a relatively isotropic region. The $S H_{1}$ source first arrival gives elliptical particle motion in this region with SV motion increasing with depth. SV motion in the $S H_{1}$-generated arrival is probably due to the local high of SV velocity (seen at $2200^{\circ}$ in Figure 36 ) which causes SV motion to move into the $S H_{1}$-generated arrival. From $2500^{\prime}$ to $2800^{\prime}$ the poor signal-to-noise in the SV source data prevent seeing the transition from split SV-generated arrivals to linear SV-generated arrivals.

The zone from 2900' to 3200' has a number of anomalous properties. At 3000' a local high in the $\mathrm{SH} / \mathrm{SV}$ velocity ratio develops with a peak $\mathrm{SH} / \mathrm{SV}$ ratio at $3250^{\circ}$. The SH $_{\mathrm{t}}$-generated first arrival gives nearly circular polarization in this zone, and both sources give a large amount of radial motion. Anomalous vertically scattered P -waves seen from the far-offset P source (Figure 27) appear to originate from about $2900^{\circ}$ and indications of scattering from $3000^{\prime}$ are seen on the $\mathbf{S H}_{r}$ source data.

The near-offset P source data show a reflection at $2900^{\circ}$, and there is a high P velocity zone at $3000^{\circ}$ to $3050^{\prime}$ followed by a low velocity zone at $3200^{\prime}$ to $3300^{\prime}$.

Figure 41 shows the hodographs for the $S H_{\mathrm{t}}$ source's first arrivals from 2875' to 3175 '. The 2875' hodograph shows a narrow ellipse in the SH-SV plane with dominant SH motion. The amount of Radial and SV motion increases in the hodographs at 2950' and 3025', and then decreases at 3175' leaving the motion similar to the 2875 ' level. A possible interpretation of the increased radial energy is P -wave scattering. If the radial motion within these shear wave arrivals is caused by P -wave scattering, it would have to be caused by an inhomogeneity near the well since the higher P-wave velocity would cause the scattered P -wave energy to separate from the shear wave arrival.

A key finding relating to this anomalous zone is that the core from $3012^{\prime}-3020^{\prime}$ had an "open and permeable fracture zone " with indications that "this fracture zone presently contains a brine" (McKibben and Andes, 1986). There is anhydrite alteration above 2900' which may serve as an impermeable cap, and the nuclear porosity $\log$ showed an increase in percent porosity between $2900^{\circ}$ and 3100' (Paillet, 1986). The lithology in this zone is mostly sandstone from 3000' to 3200' with some shale, siltstone and claystone. The anomalous seismic effects are most likely caused by a fluid. filled fracture zone near the well.

Below 3300' the SH velocity decreases with respect to SV , and below 3500' the $\mathrm{SH}_{\mathrm{r}}$-generated wave has more total travel time than the $S V$-generated wave. The $S H_{\text {t }}$ source first arrival particle motion becomes more linear from $3200^{\circ}$ to $3500^{\circ}$. The effect of motion becoming more linear may be caused by phase changes between SH and SV motion since the major axis of the ellipse moves off the SH axis and then moves back to the SH axis as the motion becomes more circular from 3500' to 3800'.

After 3800' the particle motion of the $S H_{i}$-generated arrival develops an early SH motion and a later SV motion. This is seen on the hodographs from 3925' to 4150' in Appendix 2. A radial component of motion is seen developing from 4075' to 4225'. The separation of SH and SV motion is probably a local effect, possibly from aligned fractures which have slowed the SV motion more than
that seen from the total travel time measurement. The observed radial motion could also result from fracturing which causes P-wave scattering. The P-wave scattering could be caused by SV motion since the radial motion and the SV motion occur at the same time within the first arrival. There are refiections from the near-offset shear-wave data generated between 4100' to 4200' (events R1 and G on Figures 22c and 21). The zone from 3900' to 4200' is associated with large anhydrite alteration and some epidote alteration which indicates hydrothermal activity and fracturing.

For the $\mathrm{SH}_{\mathbf{t}}$ data below 4300', radial motion stops and shear motion becomes more circular. At 4600' and 4675' the shear motion ellipse actually takes on an SV orientation as the SH motion decreases. This effect is probably due to a local high in SV/SH velocity ratio at 4675'. At 4300' Poisson's ratio reaches a minimum and begins increasing after decreasing continuously.

From 4800' to $5650^{\prime}$ both SH and SV motion are seen on arrivals from the $S H_{\mathrm{t}}$ source with the phase changing with depth giving varying shapes from circular to 'figure 8' motion at 5350' and $5425^{\prime}$. Large radial motion at $5500^{\prime}$ and $5575^{\prime}$ may be caused by local fracturing, and the radial motion again appears at the same time within the arrival window as SV motion. A P-wave reflection was also seen at 5450'. The seismic effects between $4500^{\prime}$ and $5500^{\circ}$ may be related to the epidote alteration which begins at 4400' and is indicative of strong hydrothermal process. The epidote alteration has a maximum of $1 \%$ at $4800^{\prime}$, decreases to minimal amounts from 4200' to 4400' and increases to nearly $1 \%$ again from 5450' to 5600' and then is not detected until 6100' (see Figure 5).

One observation which may have regional implications is the consistent particle motion seen in the $S V$ source first arrivals. Waves generated by the $S V$ source motion are a stable propagation mode which only shows anisoropic effects in the $2100^{\circ}$ to $2500^{\prime}$ range. Below $2800^{\prime}$ the SV source data have consistently linear SV motion in the SH-SV plane with some radial motion seen at varicus depths. The nearly isotropic propagation of $S V$ polarized waves may be a regional effect since this polarization is oriented along the axis of the spreading center while the $\mathrm{SH}_{8}$ orientation is perpendicular to the axis. Another possible explanation is that all the anisotropy is in relatively thin horizontal beds to which the SV motion, with its long vertical wavelength, is not sensitive. More tests would be
necessary to confirm either hypothesis. One such test would be conducting another VSP with a faroffset 90 degrees around the well from ours, giving the SV and $\mathrm{SH}_{\mathrm{t}}$ sources opposite polarization from this survey, and then looking for linear SH or SV motion. If the SH motion is linear it would indicate a regional cause, and if the SV motion was still linear it would indicate a local bedding cause.

FAR OFFSET ANISOTROPY SUMMARY


Figure 40. Far-offset anisotropy data summary.

## three component particle motion

far offset sh-source


Figure 41. Three component particle motion from far-offset $S H$, source near 3000' depth zone.

## CHAPTER 8

## FRACTURE ORIENTATION USING POLARIZATION DIRECTION

In the previous section the particle motion of the first shear-wave arrival was used to analyze anisotropy. The presence of anisotropy is inferred by any variation in the direction of motion from the isotropic components of radial, SH and SV motion. Within the SSGF the anisotropy is probably affected by fracturing associated with hydrothermal alteration and fluid circulation. Since the specific orientation of any one fracture zone is unknown, a method of analyzing the VSP data to detect this orientation was used. The direction of shear-wave polarization after passing through an aligned set of fractures will be parallel and perpendicular to the plane of the fracture set (Crampin 1985, Majer et al. 1986). Using this information, an atempt was made to detect zones in which the first arrival particle motion became aligned in two directions while passing through some zone in the well. This requires computing the particle motion orientation as a function of time within the first arrival wavelet.

The data used in the previous analysis of particle motion could not be analyzed directly because the coordinate system used was the wavefront system whose orientation varied as a function of depth. What is needed to detect fracture orientation is a "borehole" coordinate system which has the same coordinate orientation at every depth. Such a coordinate system has components vertical (V), horizontal towards the source (H1), and horizontal transverse to the source (H2) as shown in Figure 9. The data traces were rotated into this system by the same rotation algorithm used previously. For borehole coordinates, the original vertical component is not altered while the two horizontals are rotated into H 1 and H 2 (see ROTBOR.FOR in Appendix 4 for FORTRAN code).

The data from the far-offset shear sources were rotated into borehole coordinates for orientation analysis. Once the traces were in borehole coordinates, the polarization direction could be compared from level to level. The computation of polarization direction as a function of time was accomplished with a modification of the covariance marrix analysis used to find the geophone rotation. A window of 15 samples length ( 30 msec ) was passed through the first arrival wavelet moving 2 samples ( 4
msec ) at a time. Each time the window moved, the direction of polarization within the window was computed. Because each depth had a different first arrival window length, the number of polarization computations varied with depth. The rectilinearity function was used to estimate the degree of accuracy of each calculation.

The complete listing of polarization directions with their rectilinearity function values for the far-offset shear sources is given in Appendix 3. The angle $\phi$ is in degrees down from vertical, and the angle $\theta$ is in degrees counterclockwise from the source azimuth, which is approximately S45E. The angle $\theta$ is non-unique since $\theta+180$ degrees is also an eigen vector solution (in other words, the fracture orientation forms a plane not a vector).

Careful reading of the polarization direction data shows a number of depths which have a unique direction with high rectilinearity value. It may be that some or all of these are related to localized fracture orientation, but it is also possible they are simply errors in rotation or thin bed multiple noise. There is one zone which gives some appearance of spliting the far-offset $\mathrm{SH}_{\mathbf{r}}$-generated arrival into distinct polarization directions over a range of depths. This zone is centered at $4000^{\prime}$. The data are displayed by ploting the polarization directions on a separate borehole axis system for each level. A unit length vector is ploued for each window's polarization computation. Figure 42 shows the orientations from 3775' to $4150^{\circ}$ for arrivals from both the $S H_{1}$ and $S V$ sources.

The arrows in Figure 42 indicate the directions of polarization which develop for the SH, $_{4}$ source at $3925^{\prime}$ and $4000^{\circ}$. Notations in Appendix 3 show how the wavelet switches between these two directions of polarizations (labeled D1 and D2) at progressively earlier times within the wavelet. This is the effect one would expect to see if one polarization direction represented a faster direction of propagation. The rectilinearity function also changes from nearly one at early times, to low values as the direction switches and then retums to nearly one late in the arrival wavelet. This indicates two coherent polarization directions which overlap within the wavelet. With no other evidence of a fracture zone at this depth, it is difficult to say what this polarization analysis represents. The consistent data over a range of depths would rule out rotation error or random noise. Interbed multiple
interference could still be a source of error.

The polarization direction analysis provides another use. Figure 42 clearly shows the difference between arrivals from the $S H_{t}$ source and the $S V$ source. While the $S H_{\mathrm{t}}$-generated arrivals have a varied polarization direction, the $S V$-generated arrivals have a very consistent direction at every depth. This is a display of the isotropic propagation of the $S V$-generated waves juxtaposed with the anisotropic propagation of the $S H_{\mathrm{t}}$-generated waves. Figure 43 shows polarization direction plots for shallower depths. The difference between $S H_{t}$ and $S V$ data is again shown. The 1500' level in Figure 43 shows how the wavelet is polarized in an isotropic zone; the $S V$-generated motion is tilted because of the incident ray angle while the $\mathrm{SH}_{\mathrm{t}}$-generated motion is on the $\mathrm{H}_{2}$ axis which is parallel to the source motion.
DEPTH SU SOURCE

Figure 42. Direction of particle motion polarization for a 30 msec window within the first shear arrival for far-offset $S H$, and $5 V$ sources near 4000' depth.

## direction of particle motion polarration <br> 30 mSEC MOVING WINDOW WITHIN THE FIRST ARRIVAL

DEPTH 2800

Figure 43. Direction of particle motion polarization for a 30 msec mindow within the first shear arrival for far-offset $S H$, and $S V$ sources.

## CHAPTER 9

## CONCLUSIONS - SUMMARY

This section is a summary of data analysis results for each of the stated goals of this thesis. The areas to be summarized are standard VSP analysis, anisotropy analysis, fracture detection and orientation analysis. At times these areas overlap and results are found to apply to more than one conclusion. All the results are brought together here in a summary of the seismic wave propagation effects seen with the SSSDP VSP.

## 1. STANDARD VSP ANALYSIS

The standard VSP analysis includes velocity analysis and identification of events seen on various profiles. The velocity analysis proved adequate for the modeling and event identification analysis used in this study. The P-wave velocities were accurate when compared to the other well information available. The S-wave velocities were sufficient for modeling although the poor data quality between $2500^{\prime}$ and $4000^{\prime}$ proved a hindrance especially when analyzing the anomalous zone around $3000^{\prime}$. The lack of accuracy in the near-offset $\theta$ determination may also have created some error in the first break picks. An error in $\theta$ will cause a phase shift on each of the three components. A phase shift would cause the first trough to be at the wrong time. The consistency of the first anrival indicates this error is small, however a change in wavelet character is seen on the last 4 traces of the shear source sections. This change may be a rotation error. The difficulty in determining $\theta$ from a near offset increases with depth as the raypath approaches vertical. The shear velocities determined from these deepest 4 levels and from the noisy levels between $2500^{\prime}$ and $4000^{\prime}$ are most likely to be inaccurate.

The identification of events was successful with all coherent events identifiable as reflections, multiples, tube waves, or the more unusual scattered waves. There are relatively few reflectors within the depths covered with the VSP and none were strong enough to be seen by all of the sources .

All the VSP sections did show reflections from below the deepest VSP level. The shallowest and most often seen reflection is from the $6700^{\circ}$ to $6900^{\circ}$ zone where the lost circulation indicates a fracured reservoir is located. The far-offset $S H$, and $S V$ data have a reflection from 6800' while the near-offset P-source data have a reflection from $7000^{\circ}$ and the far-offset P-source data have a reflection from 6900'. The strength and consistent appearance of this reflection indicates that the fracture zone has an areal extent of at least a few hundred feet and that it is a relatively horizontal feature. A steeply dipping reflector would be seen to have curved moveout across the VSP section. Deeper reflectors were indicated on all the profiles, but their identification as reflections is not definite and their depth of origin can not be accurately assessed.

The detection of vertically scatered P-waves from the far-offset P-source was an indicative of a heterogeneity near $3000^{\prime}$. The scauered waves were the most obvious indication of the $3000^{\prime}$ zone which also shows a reflection on the near-offset P-source survey and shows radial particle motion within the shear arrival. The anomalous seismic effects of this zone indicate fracturing.

## 2. ANISOTROPY ANALYSIS

The analysis of travel time for orhogonal polarizations of shear sources did find anisotropic propagation. The near-offset data, which was only useful over a limited depth range, shows anisotropy within the horizontal plane. The particle motion analysis of this data was ambiguous because the geophone component rotation is inaccurate at close offsets. However, both near-offset shear source first arrivals show non-linear particle motion, an effect which is only dependent on having orthogonal components which the near-offset rotation does provide. This non-linear particle motion shows horizontal plane anisotropy exists. The lack of accurately oriented components prevented analysis for any axis of symmetry within the horizontal plane. While the absolute orientation of the near-offset shear wave particle motion is questionable, spliting of the wave form is occurring. The horizontal plane anisotropy needs to be measured with an oriented geophone to obtain more information.

Anisotropy seen with the far-offset data is the more standard SH vs SV propagation anisotropy. The travel time data shows a crossover in velocity structure with SH propagation faster shallow and SV propagation faster at depth. There is also some localized structure seen in the travel time difference plot. Most notable is the local high in SH propagation velocity from 3000' to 3500'. This may be related to the anomalous zone around $3000^{\circ}$ with ray path coverage spreading the effect to deeper levels.

The particle motion analysis showed the SV propagation mode to be an axis of symmetry with spliting of the SV motion only seen from $2100^{\prime}$ to $2500^{\circ}$. Some zones did show significant radial motion in the shear arrival from the $S V$ source which may be scattering of P-wave energy within the shear wavelet.

The $\mathrm{SH}_{\mathrm{t}}$-generated waves showed a complicated particle motion having SH, SV and radial motion whose relative amplitudes varied with depth. The SH $_{2}$-generated particle motion is indicative of anisotropic propagation. The amount of travel time difference between arrivals from $S H_{t}$ and $S V$ sources (up to $1 \%$ ) can explain the shear wave spliting seen in the data from the $S H_{\text {, source. }}$

Given the $S H$, data, the lack of splitting in the $S V$-generated first arrival is unusual. There is not enough separation in time to cause SH motion generated by the SV source to move out of the first arrival. The maximum travel time difference is 15 msec while the wavelets were analyzed with windows at least 50 msec long. The linearity of the motion in arrivals from the $S V$ source does show that the $S V$ source motion is an axis of anisotropy symmetry.

The splitting and travel time difference measured in the horizontal plane does indicate that the horizontal plane anisotropy needs to be included in future anisotropy studies. The $S V$-generated firstarrival particle motion has a component of motion with the polarization of the $S H_{r}$ source. Strong anisotropy in the horizontal plane could be misinterpreted as SV vs SH anisotropy. This is especially true for surface reflection data using SH and SV sources because they have no horizontal plane information. In this survey, the horizontal plane anisotropy has half the effect of the SH-SV anisotropy (0.5\% vs $1.0 \%$ ).

## 3. FRACTURE DETECTION

The first discovery of interest to fracture detection is the strong reflection from the fractured reservoir between 6700' and 6900'; all the sections which could be dipfiltered showed a reflection from this zone. The strength of the reflections indicates that a surface seismic survey could track this event and possibly delineate the reservoir. The relative strength of the shear wave reflections on the vertical component points out the need for three component recording. It is unfortunate that VSP data could not be recorded through this zone to surdy its effects on seismic wave propagation.

The anomalous zone at $3000^{\prime}$ shows the effects of fracturing. It is scattering P-waves from both P and S wave sources. The shear velocity anisotropy shows a local high just below this zone, an effect which indicates localized fracturing. The scattered P-wave events on the radial component of the $S V$ source data (events $K$ and $L$ on Figure 33c), could lend themselves to inversion schemes to identify their source location with more accuracy.

The appearance of radial motion within the shear wave first arrival may also be an indication of P-wave scatering from fractures. Assuming the rotation is reasonably accurate, it would take severe anisotropy to cause any significant shear motion along the propagation direction. The scattering of P-energy by the shear wave seems a most likely explanation of the large amount of radial motion seen within the shear arrival at some depths.

## 4. FRACTURE ORIENTATION ANALYSIS

The attempt to find the orientation of some fracture zone with this data set was limited by the lack of a known zone within the depths surveyed. Given that limitation, the procedure of analyzing the particle motion orientation as a function of time within the first arrival did produce some interesting results. First, the display of particle motion direction proved to be a good graphic for the linearity of the SV source first arrival motion as compared to the $S H_{t}$ source's first arrival motion. Within the $S H_{t}$ 's first arrival motion, one zone of split orientation was found. This zone, from approximately 3900' to 4200', gave the results which were expected for two polarizations of differing velocity. An
estimate could be made of the absolute orientation of these two polarizations. No core data showed fracturing in this zone. However, this is a zone with anhydrite alteration indicating hydrothermal processes which could have associated fracturing outside the borehole. While no definitive conclusions could be reached about fracture orientation, the analysis technique does show enough promise to warrant further study in a more controlled situation.

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## APPENDIX 1 <br> VELOCITY INFORMATION

All the tables and plots in appendix 1 use the following units:

Depths - Feet<br>Travel Time - Seconds<br>Velocity - Feet per Second<br>Ratios - unitless

TABLE 1-1 TRAVEL TME MEASUREMENTS
Values Of 0.00 Indicate No Data

| DEPTH | P travel time | Shr travel time | Sht travel time |
| :---: | :---: | :---: | :---: |
| 603. | 0.116 | 0.000 | 0.863 |
| 1000. | 0.197 | -.000 | 0.969 |
| 1600. | 0.272 | 0.000 | 1.908 |
| 2900. | -. 0 . 3 | -.000 | 1.226 |
| 2060. | 0.339 | 0.808 | 0.608 |
| 2180. | -. 346 | c.800 | 1.281 |
| 2164. | 0.362 | 0.608 | 1.209 |
| 2204. | 0.367 | 0.800 | 1.279 |
| 2200. | 0.363 | 0.000 | 1.287 |
| 2300. | 8. 388 | 0.000 | - 800 |
| 2360. | -.374 | 0.000 | - 0e0 |
| 2400. | -. 380 | 0.808 | 8.000 |
| 2460. | -. 388 | 0.000 | 1.320 |
| 2600. | 0.391 | 0.090 | 1.329 |
| 2660. | 0.397 | 0.800 | 1.336 |
| 2800. | 0.492 | 9.000 | -. 800 |
| 2656. | 8.488 | 0.080 | 0.096 |
| 27104. 2760. | -.412 | 0.808 | -.800 |
| 2900. | -. 421 | -.000 | 0.000 |
| 2860. | 0.425 | -. 0 e | 0.000 |
| 2ues. | 0.431 | 0.000 | 0.000 |
| 2960. | 0.436 | 0.000 | -.80e |
| 3003. | 4.04* | 0.000 | 0.000 |
| 3460. | 3.444 | 0.000 | - . 0 |
| 3180. | -. 44.4 | 0.800 | -.000 |
| 3160. | -.463 | 0.608 | 1.4.46 |
| 3250. | S. 403 | 0.804 | -.800 |
| 3300. | -. 407 | 0.000 | -. 030 |
| 3360. | 0.412 | 0.000 | -.604 |
| jave. | 0.418 | 0.048 | 1.413 |
| 3460. | 0.419 | -. 080 | 1.489 |
| 360e. | 0.483 | -.000 | . 0.000 |
| 3666. | 0.467 | 0.80 | -. 0 ees |
| 3660. 3660. | 0.496 | - . 0 en | 1.681 1.689 |
| 3100. | -. 498 | -.800 | -.800 |
| 3760. | -. 603 | -.cee | -. 000 |
| 380. | 0.607 | -0.000 | -.090 |
| 3060. | 0.512 | 9.0ee | 0.800 |
| 3000. | 0.618 | -. 0 | 0.800 |
| 3960. | 0.620 | 0.0ee | 6.046 |
| 4 ces . | 0.623 | - .800 | 0.000 |
| 4060. | 0.627 0.632 | -.808 | 1.680 |
| 4160. | -. 636 | 1.686 | 1.577 |
| 4200. | 0.639 | 1.614 | -.409 |
| 4260. | 0.642 | 1.680 | 1.647 |
| 4360. | 0. 544 | 1. 580 | 1.503 |
| 4360. | 0.600 | 1.601 | 1.600 |
| 4460. | -.663 | 1.698 1.602 | 2.803 1.800 |
| 460. | 0.650 | 1.6e7 | 1.814 |
| $4{ }^{4} 56$. | 0.683 | 1.612 | 1.014 |
| 4 40se. | -. 668 | 1.614 | 1.026 |
| 4060. | -. 660 | 1.623 | 1.031 |
| 4100. | 0.672 6.675 | 1.636 | 1.631 |
| 4160. | 6.676 6.679 | 1.636 1.840 | 1.842 1.648 |
| 4060. | -.643 | 1.646 | 1.064 |
| 4900. | -. 686 | 1.660 | 1.868 |
| 4960. | - . 0 en | 1.666 | 1.864 |
| beee. | 0.603 | 1.061 | 1.610 |
| 6060. | 0. 600 | 1.066 | 1.076 |
| 6190. | 0.698 | 1.671 | 1.082 |
| 6160. 6200. | 5.602 | 0.806 | - 000 |
| 6280. 6269. | 0.086 | 1.000 | 1.694 |
| 6269. 6300. | -. 609 | 1.601 | $1.69 \%$ 1.185 |
| 8360. | -. 615 | 1.890 | 1.710 |
| 6490. | 0.619 | 1.705 | 1.716 |
| 6460. | 0.023 | 1.711 | 1.723 |
| 6640. | 0.626 | -.060 | 1.728 |
| 5650. | 0.029 | 1.723 | 1.733 |
| 5086. | -. 632 | 1.730 | 1.738 |
| 6650. | 0.636 | 1.713 | 1.743 |

TABLE 1-2 AVERAGE VELOCITY CALCULATIONS
Values Of 0.00 Indicate No Data

| DEPTM | P VELOCITY | SHP VELOCITY | SHE VELOCITY |
| :---: | :---: | :---: | :---: |
| 680. | 6027. | 0 - | 893. |
| 1006. | 5346. | 6. | 1689. |
| 1500. | 6824. | - | 1390. |
| 2000. | 6. | 6. | 1858. |
| 2856. | 8112. | 6. | 0. |
| 2100. | 8149. | e. | 1682. |
| 2168. | 6167. | 0. | 1711. |
| 2286. | 8219. | 6. | 1736. |
| 2268. | 6263. | e. | 1784. |
| 2380. | 6303. | c. | 6. |
| 2363. | 8334. | 6. | 0. |
| 2406. | 6306. | - | - |
| 2464. | 0396. | e. | 1876. |
| 250. | 6446. | 6. | 1896. |
| 2668. | 8467. | 6. | 1922. |
| 208. | 8611. | 4. | - |
| 2868. | 0689. | 8. | $\theta$ - |
| 2100. | 6694. | 6. | e. |
| 276 . | 686. | e. | 0. |
| 280 | A18. | e. | - |
| 2860. | 6727. | E. | 0. |
| 290. | 6744. | - | 6. |
| 2960. | 4803. | - | 6. |
| 3000. | 6862. | ¢ | 4. |
| 3000 . | s903. | e. | - |
| 3100. | 6962. | - | - |
| 3160. | A9+3. | * | 228. |
| 328. | 7133. | 0. | e. |
| 3260. | 7849. | e. | $\theta$ e. |
| 3380. | 1890. | 6. | 0. |
| 3360 | 1126. | - | 0. |
| 340. | 7111. | - | 2317. |
| 3460 . | 723 . | e. | 2346. |
| 360. | 7273. | - | 4. |
| 3650. | 7316. | - | $\theta$ - |
| 384. | 7372. | e. | 2417. |
| 106s. | 7100. | - | 2427. |
| 310. | 7464. | - | - |
| 3768. | 7419. | 6. | $s$ |
| sues. | 1514. | e. | e. |
| 3064. | 7642. | - | - |
| 390. | 7580. | - | 0. |
| 1064. | 7418. | - | - |
| 4 Ane. | 167 . | e. | 8 8. |
| 4*Ft. | 1700. | 6. | 6. |
| 41410 | 7127. | E. | 2817. |
| 41,0 | 1171. | 2859. | 2038. |
| 4200. | 7812. | 2615. | 0. |
| 4260 | 7 76. | 2607. | 2086. |
| 43 E . | 7808. | 2710. | 278. |
| 4364. |  | 2741. | 2729. |
| 4404. | 7976. | 2783. | 2781. |
| 4460. | -222. | 2784. | 2774. |
| 4600. | pert. | 2asa. | 2784. |
| 4666. | eses. | 2829. | 2818. |
| 4004. | 8144. | 284. | 2837. |
| 466*. | 8140. | 2871. | 2467. |
| 476 | 2234. | 2893. | 2971. |
| 476. | 0263. | 2011. | 2497. |
| 486 . | 8308. | 2933. | 2918. |
| 4860. | 8335. | 2964. | 2036. |
| 4980. | 8377. | 2976. | 2960. |
| $4959 .$ | \% | 2096. | 2988. |
| 6 wer. | 8447. | 3815. | 2809. |
| 6ibl. | 8489. | 3037. | 302 . |
| 6100. | 8620. | 3157. | 3031. |
| 5ibe. | 8689. | \%. | 30. |
| 624. | 840s. | e. | 3476. |
| 5260. | 4836. | 3117. | 3896. |
| 63400. | 8674. | 3137. | 3113. |
| 6368. | 4713. | 3168. | 3134. |
| 6498. | 0737. | 3172. | 3162. |
| 5464. | 8781. | 3100. | 3168. |
| $560{ }^{6}$. | 4799. | - | 318 d . |
| 6668. | 4836. | 3228. | 3287. |
| 588. | 8873. | 3242. | 3227. |
| 6868. | 0016. | 3286. | 3246. |

TABLE 1-3 100' INTERVAL VELOCITY CALCULATIONS
Values Of 0.00 Indicate No Data

| DEPTM | $P$ | DEPTH | SMr | OEPTM | SHE |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 102. | cese. | 4. | $\cdots$ | 1004. | 2131. |
| 6. | - | - | $\cdots$ | 150 | 1994. |
| 1776. | eeer. | - | - | E. | - |
| 1 Hes. | 8184. | + | 6. | 2066. | 2321. |
| 2200. | 7615. | e. | - | 2875. | 3462. |
| 2164. | 8263. | e. | e. | 2160. | 66.2 |
| 22 . | 9-3. | 0. | 6. | 22 . | 6606. |
| 2286 . | 911. | 4. | - | - | 6. |
| 234. | 916. | 6 | - | - | - |
| 2364. | 8268. | e. | e. | 6. | E. |
| 2443. | 2289. | - | e. | 2350. | $0 \times 12$. |
| 24661. | 023. | E. | - | 2376. | 68. |
| 260. | -28. | e. | - | 2840. | 32 Cl . |
| 258.6 | e02\%. | - | - | B. | - |
| 2048. | 11010. | - | e. | - | - |
| 2 t 8. | 0931. | e. | - | e. | E |
| 2 'tis. | 9039. | $\bullet$ | - | - | e. |
| 2765 | 11048. | - | s. | -. | - |
| $20 *$. | 0943. | e. | e. | - | - |
| 285 . | 0046. | - | 0. | E. | - |
| 20. | 9047. | e. | - | e. | e. |
| 2968. | 11850. | e. | B. | E. | E. |
| 3 Cx . | 12431. | - | - | $\theta$ e. | - |
| 3060. | 1244. | - | E. | 6. | - |
| 3100. | 1109. | e. | e. | 2060. | 5097. |
| 3168. | 11081. | 4. | 0. | 0. | e. |
| $32 *$. | 896. | - | e. | e. | e. |
| 3264. | 0067. | - | - | - | - |
| 330 . | 1108 . | - | e. | e. | - |
| 3360. | 1107 \% | - | -. | 3276. | 6720. |
| 3464. | 14230. | - | - | 124. | 6700. |
| $34 b e \text {. }$ | 44232. | - | - | - | E. |
| $16$ | 12464. | - | - | - | - |
| 3668. | 14236. | - | e. | 3526. | 7117. |
| 3493. | 12467. | - | - | 386. | 1872. |
| $3 \times 6 .$ | 12469. | - | - | - | - |
| $37 \mathrm{e} \text {. }$ | 12469. | e. | e. | e. | - |
| 3768. | 1167 . | e. | e. | e. | - |
| 3ne. | 11077. | - | $\theta$ - | - | - |
| 3960. | 13177. | - | - | - | - |
| 300. | 82483. | - | 6. | - | . |
| 3956. | 14246. | - | e. | 0. | - |
| ace. | 1424. | - | E. | - | - |
| 4054. | 110er. | - | - | 3976. | 7230. |
| 418. | 124ar. | *. | 0. | $30 \times 5$. | 7311. |
| 4160. | 14248. |  |  | $6$ | \% |
| $420$ | 142 Es . | $4204$ | $8868 .$ | $425$ | 9876. |
| 4280. | $14251$ | $4250$ | $0113 .$ | $4226 .$ | $0362 .$ |
| - | $0$ | $430$ | sere. | $43 E$ | eas. |
| 4369. | 14252. | 4160. | 9076. | 4366. | 9076. |
| 4316. | 14086. | 44. | W076. | 44 Cl | 0977. |
| 4463. | 18420. | 4468. | 9076. | 4460. | wre. |
| 467. | 14264. | 460. | 9414. | 460. | 081. |
| 4664. | 14266. | 4568. | 0071. | 4661. | 971. |
| dece. | 18631. | 48w. | 0872. | $46+5$. | 8316. |
| 4463. | 18332. | 4868. | 0078. | 406 . | 0318. |
| 476 . | 14237. | 470. | 0317. | 47 C | 0387. |
| 476. | 14260. | 4764. | 8317. | 4760. | 073. |
| 4eet. | $1425 \%$. | 4898. | 008. | 48 4. | cels. |
| 486. | 14250. | 4060. | 908. | 4463. | 9814. |
|  | $\theta .$ | 400. | 0081. | 490. | 901. |
| 4056. | 14289. | 4068. | $0 \times 74$. | 4960. | 9874. |
| 4976. | 14972. | 6une. | 9074. | 503. | 074. |
| 645. | 18937. | 64e. | gen2. | 6054. | -210. |
| $61{ }^{\text {6 }}$ | 10030. | - | - | * | 4. |
| 6160. | 10030. | - | - | 6160. | 0319. |
| 628. | 14202. | 6116. | 0364. | 6176. | 8480. |
| 6260. | 14263. | 62 . | 068. | 626 . | D18. |
| 530. | 10642. | 6308. | 917. | $63 *$. | 9617. |
| 6364. | 14282. | 6360. | 780 | 6360. | 078. |
| $64 \times 2$. | 12480. | 64t. | 7 tes. | 640. | 1880. |
| 5468. | 14285. |  | - | 646 . | -321. |
| 654. | 18641. | 560. | 8321. | 664. | 9046. |
| 666 . | 10442. | 6525. | 7803. | 6568. | 9906. |
|  | 18044. | sate. | 90* | Gex. | 0004. |

TABLE 1-4 500' INTERVAL VELOCITY CALCULATIONS
Values Of 0.00 Indicate No Data

| OEPTM | $P$ | OEPTH | SHP | DEPTH | 5HE |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1426. | 8022. | 6. | - | B. | 6. |
| 17 er | 7612. | 0 | 6. | e. | 4. |
| 1976. | 8231. | 0. | 0. | 1976. | 4190. |
| 2496. | 0304. | 0. | 0. | 2268. | 4911. |
| 2380. | 1647. | 4. | 0. | 2276. | 4088. |
| 2368. | 181. | 0. | 0. | - | 0. |
| 24tes. | 9147. | - | e. | e. | - |
| 246 . | 0423. | 0. | *. | $\bullet$ | 6. |
| 280 | 930. | - | - | - | 6. |
| 2650. | 0340. | d. | *. | - | 6. |
| 2086. | 0551. | $\theta$ - | - | - | 6. |
| 28be. | 0741. | a. | $\theta$. | 6. | - |
| 27 es. | 0031. | 6. | 0. | e. | 8. |
| 2160. | 10143. | 6. | 0. | 6. | 0. |
| $2{ }^{2} 80$. | 18877. | 6. | 6. | $\bullet$ - | - |
| 2858. | 18808. | 4. | 0. | e. | - |
| 2094. | 18601. | 0. | 0. | 288. | 6087. |
| 2060. | 11054. | * | 4. | - | - |
| 304. | 1058. | 0 | 6. | - | - |
| 3 186. | 1017. | - | - | - | - |
| 11 ce. | 1018. | e. | - | 4 | - |
| 3160. | 11081. | $\theta$ c | e. | 2976. | 6172. |
| 324. | 11577. | 0. | $\theta$. | 3010. | $6210 .$ |
| 3250. | 11676. | E. | 0. | - | 0. |
| 3388. | 1450 . | 0. | 0. | e. | $\theta$ - |
| 3360. | 14867. | - | 6. | 3075. | 6333. |
| 3406. | 11859. | 0. | © | 3404. | 083. |
| 3456. | 12140. | 0. | e. | 4 | 0 |
| $36 \%$. | 12484. | - | - | - | - |
| 3664. | 124E5. | 3. | 6. | e. | 6. |
| 360. | 12467. | 0. | 0. | 0. | $\theta$ - |
| 3060. | 12469. | 0. | 0. | 0 | 0. |
| 3780. | 12166. | 4. | 0 | - | - |
| 3766. | 12468. | 0. | $\omega$ | - | - |
| 3000. | 12481. | 0. | 0. | 0. | - |
| 3156. | 11860. | $\theta$ e. | $\theta \cdot$ | 3060... | 7121. |
| 300. | 12463. | e. | 4. | 301 | 7331. |
| 3968. | 12168. | 0. | $\theta$. | 6. | \%. |
| 4 ce. | 12764. | 4. | - | 3961. | 1816. |
| 4ete. | 12706. | 0. | 0. | 3976. | 7716. |
|  | - | - | 0. | 4093. | 7043. |
| 4160. | 13474. | 6. | E. | 4426. | 1968. |
| 424 . | 13853. | - | - | 4060. | 8869. |
| 4264. | 13184. | e. | - | 4075. | 1073. |
| 43 ¢ | 13156. | - | 3 | 410 . | 1160. |
| 4360. | 14671. | - | e. | 4350. | 9231. |
| 446. | 14872. | 4480. | 8681. | 4460. | 9230. |
| 4458. | 16117. | $446 \cdot$. | 0230. | 4426. | 9146. |
| 460. | 14673. | 4580. | 0871. | 468. | 8 800. |
| 4558. | 15110. | 4560. | 9230. | 4660. | 0,71. |
| 4576. | 1413. | 4890. | 9246. | 4890. | 8910. |
| 4650. | $15120 .$ | $4068 \text {. }$ | 024. | 4668. | 6צ10. |
| ©. | E. | $4700 .$ | 9416. | 476. | 8016. |
| 4760. | 14877. | 4768. | 4241. | 4768. | 8911. |
| -1at. | 16122. | 4000. | 0241. | 4y0. | 8911. |
| 4864. | 16423. | 486 | 9410. | 406. | 9185. |
| 400. | 16121. | 0. | ©. | 0. | E. |
| 4061. | 16124. | - | - | 4969. | 8768. |
| Seles. | 15124. | 6404. | 0690. | 6ate . | 0912. |
| 6860. | 16125. | 5068. | 0690. | 5864. | 0760. |
| 610. | 16508. | 618. | 9418. | 6108. | 8913. |
| 6160. | 15820. | 6160. | 0875. | 6168. | 8767. |
| 6116. | 14644. | 620. | 0914. | \$2as. | 4461. |
| $6250 .$ | 16127. | 0. | $0 .$ | 5269. | 8087. |
| $63(2$ | 16127. | 6380. | ar6e. | 6300. | wet. |
| 636 | 16120. | 6350. | 8481. | 5360. | 0916. |
| 6ate. | 1612. | 6315. | ges7. | 6376. | 9enz. |

TABLE 1-5 AVERAGE VELOCITY RATIOS
Values Of 0.00 Or 1.00 Indicate No Data


TABLE 1-6 $500^{\prime}$ INTERVAL VELOCITY RATIOS
Values Of 0.00 Indicate No Data

| OEPTM | P/SMr | SHP | POISSON | DEPTM | P/SME | SHE POISSON |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| - | - . en |  | e. 00 | - | - . | 0. 088 |
| - | - 0 . 0 er |  | - . 19 | - | -.804 | - 0.80 |
| 1976. | 1.968 |  | 6. 326 | c. | c. 600 | -. 088 |
| 2004. | 1.969 |  | 0.32. | c. | - . 000 | - . 000 |
| 2026. | 1.920 |  | 0.314 | 4. | - . 000 | - .104 |
| - | 1.800 |  | - . 0ad | 0. | -. 0 . ${ }^{\text {er }}$ | c. |
| ¢ | 0. 0.00 |  | - . 080 | $\bigcirc 0$. | - . 000 | 0.000 |
| - | - 0.60 |  | - 0.83 | 0. | - 0.80 | P. 018 |
| - | -.00 |  | e. 080 | 0 | c. 080 | 8. dos |
| - | c.as |  | t. 8 . ${ }^{\text {ces }}$ | 6 | - . 0 - | - . 중사 |
| 0. | - del |  | - . 46 | *. | - . 300 | - bidt |
| 0. | 0.081 |  | 0. 004 | 4. | - . 000 | - . 0808 |
| 0 | 0.8en |  | 0. 086 | 3 . | 8.802 | 0.090 |
| 0. | - 0.08 |  | 4.096 | - | *. 800 | 6. 800 |
| - | 0.8en |  | - 0.318 | * | - . -3 | 6.808 |
| - | -.800 |  | 0.80 | c. | - . - | -.ctes |
| 2164. | 1.780 |  | c. 272 | 4. | - . 0 er | - . oce |
| . 0 | 6.808 |  | -.0es | c. | - . 800 | 6.880 |
| e. | - 0 - |  | c.04 | 0. | - . 803 | - 0.000 |
| - | 6.80 |  | - . 89 | 4. | 4.004 | 4.006 |
| e. | 3.804 |  | 8.893 | 6. | 0. 363 | - . 408 |
| 2976. | 1.734 |  | -. 261 | 0. | - . $0^{0}$ | 0. 008 |
| 198. | 1.750 |  | 6. 280 | 6. | -.tee | -.tes |
| - | 1.04 |  | 9.096 | - | - . ${ }^{\text {a }}$ | - . 086 |
| - | 0. An |  | e.en | - | - . sed | 0.880 |
| 3076. | 1.774 |  | 0.287 | 4. | - . 080 | 0.000 |
| 34ts. | 1.731 |  | -. 263 | 0. | 0.0ec | 8.006 |
| - | -.ter |  | -. 0et | $\theta$ * | - . 0 | -.000 |
| - | e.ce |  | - .cer | - | -. 0 | 0.080 |
| - | c.en |  | 0.691 | - | 0.8es | -. 000 |
| - | 0.0ce |  | 0.09 | 6. | 6.40 | 0.00 |
| - | -. 80 |  | - . 60 | - | - .0er | 0.600 |
| - | -. An |  | 0.003 | - | - .800 | 0.003 |
| - | 0.8ee |  | B.8es | e. | - . A0 | 0.000 |
| - | - . 0 cer |  | 0.00 | $\theta$ - | - .80 | - . 000 |
| 3068. | 1.801 |  | 0.210 | 6 | - bee | 0.000 |
| 30. | 1.74 |  | c. 236 | 0. | -. 0 - | 0.00 |
|  | -. EC |  | - . 0 | 0. | 1.40 | 0.000 |
| 306. | 1.86 |  | 0.216 | 420. | - . ese | 0.090 |
| 1916. | 1.847 |  | - 208 | 4260. | - 0.80 | - . 000 |
| - | - 000 |  | - . 0 | 4 . | 0.000 | 0.008 |
| 4026. | 1.021 |  | -. 193 | 4360 . | 0.000 | 0.000 |
| 4 406. | 1.623 |  | 0.194 | 4376. | 0.09 | 0.060 |
| 416. | 1.641 |  | B. 294 | 445 . | 6.0ec | 0.080 |
| 4180. | 1.818 |  | c.191 | 4602. | 6.003 | 0.000 |
| 4369. | 1.689 |  | 0.112 | 4653. | - .ever | 0.008 |
| 44. | 1.681 |  | 0.172 | 4A0). | 1.706 | 0.238 |
| 4425. | 1.622 |  | ©. 193 | 4464. | 1.036 | 0. 202 |
| 4680. | 1.647 |  | 0.208 | 4680. | 1.610 | 0.151 |
| 4660. | 1.837 |  | 0.210 | 4650. | 1.838 | 0.20? |
| $4676 \text {. }$ | 1.649 |  | 0.280 | 4676. | 1. 696 | 4. 176 |
| 4650. | 1.897 |  | 0.214 | 4456. | 1.836 | 0.282 |
| E. | - 0.00 |  | *. 606 | 0. | - 0.04 | -. 000 |
| 4760. | 1.847 |  | - 238 | 4759. | 1.688 | 0.172 |
| 4 -18. | 1.697 |  | d.234 | 4883. | 1.836 | 0.282 |
| 4050 | 1.727 |  | - 2.24 | 4150. | 1.000 | 0.163 |
| 13. | - . 0 |  | - 0.00 | 0. | - . ${ }^{\text {a }}$ | -. 000 |
| 4950. | 1.727 |  | - 24. | $\theta$ E | - . 800 | 6.803 |
| 6rye. | 1.897 |  | -. 234 | 6and. | 1.678 | c. 103 |
| Fese. | 1.727 |  | - 248 | 5069. | 1. 576 | 0.163 |
| 510. | 1.760 |  | 3.250 | 610. | 1.866 | 0.213 |
| 5169. | 1.727 |  | 0.248 | 6163. | 1.687 | 0.219 |
| 6176. | 1.738 |  | 0.249 | 6176. | 1. 449 | 0. 209 |
| 6260. | 1.75 ¢ |  | -.261 | 6. | 0.400 | 6. 080 |
| 5380. | 1.760 |  | - 281 | 63es. | 1.727 | 0.248 |
| 6364. | 1.807 |  | -.234 | 635 . | 1.7et | 0.272 |
| 6376. | 1.804 |  | -.213 | 6376. | 1.722 | - 246 |



Figure 1-1 Interval Poisson's Ratio Calculated From SH, Data With $500^{\prime}$ Interval Spacing.

## APPENDIX 2

## PARTICLE MOTION HODOGRAPHS

The hodographs in this appendix are listed by level number instead of depth. The following tables cross-reference level number and depth for the various data sets. Table 3-1 is for all the faroffset data, while table 3-2 is for the near-offset $P$ and $S H_{1}$ source data and table 3-3 is for the nearoffset $S H_{r}$ source data.

| Table 2-1 |  |
| :---: | :---: |
| Far-Offset Data, All Sources |  |
| Level | Depth |
| 1 | 1500 |
| 2 | 1900 |
| 3 | 1975 |
| 4 | 2050 |
| 5 | 2125 |
| 6 | 2200 |
| 7 | 2275 |
| 8 | 2350 |
| 9 | 2425 |
| 10 | 2500 |
| 11 | 2575 |
| 12 | 2650 |
| 13 | 2725 |
| 14 | 2800 |
| 15 | 2875 |
| 16 | 2950 |
| 17 | 3025 |
| 18 | 3100 |
| 19 | 3175 |
| 20 | 3250 |
| 21 | 3325 |
| 22 | 3400 |
| 23 | 3475 |
| 24 | 3550 |
| 25 | 3625 |
| 26 | 3700 |
| 27 | 3775 |
| 28 | 3850 |


| Table 2-1 |  |
| :---: | :---: |
| Far-Offset Data, All Sources |  |
| Level | Depth |
| 29 | 3925 |
| 30 | 4000 |
| 31 | 4075 |
| 32 | 4150 |
| 33 | 4225 |
| 34 | 4300 |
| 35 | 4375 |
| 36 | 4450 |
| 37 | 4525 |
| 38 | 4600 |
| 39 | 4675 |
| 40 | 4750 |
| 41 | 4825 |
| 42 | 4900 |
| 43 | 4975 |
| 44 | 5050 |
| 45 | 5125 |
| 46 | 5200 |
| 47 | 5275 |
| 48 | 5350 |
| 49 | 5425 |
| 50 | 5500 |
| 51 | 5575 |
| 52 | 5650 |


| Table 2-2 |  |
| :---: | :---: |
| Near-Offset P and $\mathrm{SH}_{1}$ Data |  |
| Level | Depth |
| 1 | 500 |
| 2 | 1000 |
| 3 | 1500 |
| 4 | 2000 |
| 5 | 2050 |
| 6 | 2100 |
| 7 | 2150 |
| 8 | 2200 |
| 9 | 2250 |
| 10 | 2300 |
| 11 | 2350 |
| 12 | 2400 |
| 13 | 2450 |
| 14 | 2500 |
| 15 | 2550 |
| 16 | 2600 |
| 17 | 2650 |
| 18 | 2700 |
| 19 | 2750 |
| 20 | 2800 |
| 21 | 2850 |
| 22 | 2900 |
| 23 | 2950 |
| 24 | 3000 |
| 25 | 3050 |
| 26 | 3100 |
| 27 | 3150 |
| 28 | 3200 |
| 29 | 3250 |
| 30 | 3300 |
| 31 | 3350 |
| 32 | 3400 |
| 33 | 3450 |
| 34 | 3500 |
| 35 | 3550 |
| 36 | 3600 |
| 37 | 3650 |
| 38 | 3700 |
| 39 | 3750 |
| 40 | 3800 |

Table 2.2
Near-Offset P and SH, Data

| Level | Depth |
| :---: | :---: |
| 41 | 3850 |
| 42 | 3900 |
| 43 | 3950 |
| 44 | 4000 |
| 45 | 4050 |
| 46 | 4100 |
| 47 | 4150 |
| 48 | 4200 |
| 49 | 4250 |
| 50 | 4300 |
| 51 | 4350 |
| 52 | 4400 |
| 53 | 4450 |
| 54 | 4500 |
| 55 | 4550 |
| 56 | 4600 |
| 57 | 4650 |
| 58 | 4700 |
| 59 | 4750 |
| 60 | 4800 |
| 61 | 4850 |
| 62 | 4900 |
| 63 | 4950 |
| 64 | 5000 |
| 65 | 5050 |
| 66 | 5100 |
| 67 | 5150 |
| 68 | 5200 |
| 69 | 5250 |
| 70 | 5300 |
| 71 | 5350 |
| 72 | 5400 |
| 73 | 5450 |
| 74 | 5500 |
| 75 | 5550 |
| 76 | 5600 |
| 77 | 5650 |


| Table 2-3 |  |
| :---: | :---: |
| Near-Offset SH, Data |  |
| Level | Depth |
| 1 | 3050 |
| 2 | 3100 |
| 3 | 3150 |
| 4 | 3200 |
| 5 | 3250 |
| 6 | 3300 |
| 7 | 3350 |
| 8 | 3400 |
| 9 | 3450 |
| 10 | 3500 |
| 11 | 3550 |
| 12 | 3600 |
| 13 | 3650 |
| 14 | 3700 |
| 15 | 3750 |
| 16 | 3800 |
| 17 | 3850 |
| 18 | 3900 |
| 19 | 3950 |
| 20 | 4000 |
| 21 | 4050 |
| 22 | 4100 |
| 23 | 4150 |
| 24 | 4200 |
| 25 | 4250 |
| 26 | 4300 |
| 27 | 4350 |
| 28 | 4400 |
| 29 | 4450 |
| 30 | 4500 |
| 31 | 4550 |
| 32 | 4600 |
| 33 | 4650 |
| 34 | 4700 |
| 35 | 4750 |
| 36 | 4800 |
| 37 | 4850 |
| 38 | 4900 |
| 39 | 4950 |
| 40 | 5000 |

Table 2-3

| Near-Offset SH Data |  |
| :---: | :---: |
| Level | Depth |
| 41 | 5050 |
| 42 | 5100 |
| 43 | 5150 |
| 44 | 5200 |
| 45 | 5250 |
| 46 | 5300 |
| 47 | 5350 |
| 48 | 5400 |
| 49 | 5450 |
| 50 | 5500 |
| 51 | 5550 |
| 52 | 5600 |
| 53 | 5650 |



## FAR OFFSET P-WAVE SOURCE


particle motion at level 10

particle motion at level 11

particle motion at level 12


particle motion at level 15


PARTICLE MOTION AT LEVEL 16


PARTICLE MOTION AT LEVEL 17


PARTICLE MOTION AT LEVEL 18


## - <br> FAR OFFSET P-WAVE SOURCE


particle motion at level 21

particle motion at level 22

particle motion at level 23

particle motion at level 24


## FAR OFFSET P-WAVE SOURCE



PRRTICLE MOTION AT LEVEL 26


PARTICLE MOTION AT LEVEL 28

particle motion at level 29


PARTICLE MOTION AT LEVEL 30


FAR OFFSET P- WAVE SOURCE



PARTICLE MOTION AT LEVEL 38


PARTICLE MOTION AT LEVEL 40


PARTICLE MOTION AT LEVEL 41


PARTICLE MOTION AT LEVEL. 42


## FAR OFFSET P-WAVE SOURCE


particle motion at level 45


PARTICLE MOTION AT LEVEL 46


PARTICLE MOTION AT LEVEL 47


PARTICLE MOTION AT LEVEL 48


## FAR OFFSET. P-WAVE SOURCE


particle motion at level 50

particle motion at level 51

particle motion at level 52

particle motion at level 53

particle motion at level 54



PRRTICLE MOTION AT LEVEL 4


PARTICLE MOTION AT LEVEL 5


PARTICLE MOTION AT LEVEL 6


## FAR OFFSET SV SOURCE



PARTICLE MOTION AT LEVEL 9


PRRTICLE MOTION AT LEVEL 10


PARTICLE MOTION AT LEVEL 11

particle motion at level 12


## FAR OFFSET SV SOURCE



PRRTICLE MOTION AT LEVEL 14


PARTICLE MOTION AT LEVEL 15


PARTICLE MOTION AT LEVEL 16


PARTICLE MOTION AT LEVEL 17


PARTICLE MOTION AT LEVEL 18



FAR OFFSET SV SOURCE


PARTICLE MOTION AT LEVEL 26


PRRTICLE MOTION RT LEVEL 27


PARTICLE MOTION AT LEVEL 28

particle motion at level 29


PARTICLE MOTION AT LEVEL 30


FAR OFFSET SV SOURCE


PARTICLE MOTION AT LEVEL 32


PARTICLE MOTION AT LEVEL 33


PARTICLE MOTION RT LEVEL 34


PARTICLE MOTION AT LEVEL 35


PARTICLE MOTION AT LEVEL 36


FAR OFFSET SV SOURCE


PARTICLE MOTION AT LEVEL 38


PARTICLE MOTION AT. LEVEL 39


PARTICLE MOTION AT LEVEL 40


PARTICLE MOTION AT LEVEL 41


PARTICLE MOTION AT LEVEL 42



PARTICLE MOTION AT LEVEL 46


PARTICLE MOTION AT LEVEL 47


PARTICLE MOTION AT LEVEL 48



PRRTICLE MOTION AT LEVEL 50


FAR OFFSET SH ${ }_{t}$ SOURCE






PARTICLE MOTION AT LEVEL 5




FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE


PARTICLE MOTION AT LEVEL 8


PARTICLE MOTION AT LEVEL 9



PARTICLE MOTION AT LEVEL 11


PARTICLE MOTION AT LEVEL 12



## FAR OFFSET SH $_{t}$ SOURCE



PARTicle motion at level 20

particle motion at level 21

particle motion at level 22

particle motion at level 23

particle motion at level 24


FAR OFFSET SH $_{t}$ SOURCE

particle motion at level 26


PARTICLE MOTION AT LEVEL 27


PRRTICLE MOTION AT LEVEL 28

particle motion at level 29


PRRTIICLE MOTION AT LEVEL 30


FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE


PARTICLE MOTION AT LEVEL 32


PARTICLE MOTION AT LEVEL 33



PARTICLE MOTION AT LEVEL 35


PARTICLE MOTION AT LEVEL 36


## FAR OFFSET SH $_{t}$ SOURCE



PARTICLE MOTION AT LEVEL 38



PARTICLE MOTION RT LEVEL 41


PARTICLE MOTION AT LEVEL 42


FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE


PARTICLE MOTION AT LEVEL 44


PRRTIClE MOTION AT LEVEL 45



PARTICLE MOTION AT LEVEL 47


PARTICLE MOTION AT LEVEL 48


FAR OFFSET SH ${ }_{t}$ SOURCE


PARTICLE MOTION AT LEVEL 50


## - <br> NEAR OFFSET P-WAVE SOURCE



PARTICLE MOTION RT LEVEL 3


PARTICLE MOTION RT LEVEL 4


PRRTICLE MOTION RT LEVEL. 5


PARTICLE MOTION AT LEVEL 6



NEAR OFFSET P-WAVE SOURCE

particle motion at level 16

particle motion at level 17
PARTICLE MOTION AT LEVEL


15





PARTICLE MOTION AT LEVEL 26

particle motion at level 27 Coses)

PARTICLE MOTION AT LEVEL 28


PARTICLE MOTION AT LEVEL 29

particle motion at level 30


particle motion at level 33

particle motion at level 34

particle motion at level 35

particle motion at level 36



## - NEAR OFFSET P-WAVE SOURCE



PRRTICLE MOTION AT LEVEL 44


PRRTICLE MOTION RT LEVEL 45


PARTICLE MOTION AT LEVEL 46


PRRTICLE MOTION RT LEVEL 47


PARTICLE MOTION AT LEVEL 48



## - NEAR OFFSET P-WAVE SOURCE

particle motion at level 55

particle motion at level 57

particle motion at level 58

particle motion at level 59

particle motion at level 60 Cos:


## - <br> NEAR OFFSET P-WAVE SOURCE



PARTICLE MOTION AT LEVEL 69 Coses:
particle motion at level 70

particle motion at level 71

particle motion at level 72



NEAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE


PARTICLE MOTION AT LEVEL 2


PARTICLE MOTION AT LEVEL 5


PARTICLE MOTION AT LEVEL 6


## NEAR OFFSET SH $_{t}$ SOURCE


particle motion at level 9


PARTICLE MOTION RT LEVEL 10


PRRTICLE MOTION AT LEVEL 11

particle motion at level 12


## NEAR OFFSET SH ${ }_{t}$ SOURCE


particle motion at level 14

particle motion at level 15

particle motion at level 16

particle motion at level 17

particle motion at level 18



$$
\text { PARTICLE MOTION AT LEVEL } 44
$$



PARTICLE MOTION AT LEVEL 45


PARTICLE MOTION AT LEVEL 46


PARTICLE MOTION AT LEVEL 47


PARTICLE MOTION AT LEVEL 48


NEAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE


PARTICLE MOTION AT LEVEL SO

particle motion at level 51


PARTICLE MOTION AT LEVEL 52


PRRTICLE MOTION AT LEVEL 53


PRRTICLE MOTION AT LEVEL 54


NEAR OFFSET SH ${ }_{t}$ SOURCE


PARTICLE MOTION AT LEVEL 57



PARTICLE MOTION AT LEVEL 59


PARTICLE MOTION RT LEVEL 60


## PRRTICLE MOTION AT LEVEL 61



PRRTICLE MOTION AT LEVEL 62


PARTICLE MOTION AT LEVEL 64

particle motion at level 65


PARTICLE MOTION AT LEVEL 66


## NEAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE



PRRTICLE MOTION AT LEVEL 68


PARTICLE MOTION RT LEVEL 69


PARTICLE MOTION AT LEVEL 70


PARTICLE MOTION AT LEVEL 71

particle motion at level 72


## NEAR OFFSET SH ${ }_{t}$ SOURCE

PARTICLE MOTION AT LEVEL 73


PRRTICLE MOTION RT LEVEL 74


PARTICLE MOTION AT LEVEL 75


PARTICLE MOTION AT LEVEL 76

particle motion at level 77



particle motion at level 29


PARTICLE MOTION AT LEVEL 30


- NEAR OFFSET SH r SOURCE


PARTICLE MOTION RT LEVEL 32


PRRTICLE MOTION AT LEVEL 33



PARTICLE MOTION AT LEVEL 35


PARTICLE MOTION AT LEVEL 36


NEAR OFFSET $\mathrm{SH}_{\mathrm{r}}$ SOURCE


PARTICLE MOTION AT LEVEL 38


PARTICLE MOTION AT LEVEL 39



PARTICLE MOTION AT LEVEL 41


PARTICLE MOTION AT LEVEL 42


- NEAR OFFSET $\mathrm{SH}_{r}$ SOURCE
particle motion at level 43




## NEAR OFFSET SH ${ }_{r}$ SOURCE



## APPENDIX 3

## POLARIZATION DIRECTION -- DATA AND PLOTS

The data in this appendix are listed by level number instead of depth. The following table cross-references level number and depth for the far-offset shear sources used in the polarization analysis.

| Table 3-1 |  |
| :---: | :---: |
| Far-Offset Data |  |
| Level | Depth |
| 1 | 1500 |
| 2 | 1900 |
| 3 | 1975 |
| 4 | 2050 |
| 5 | 2125 |
| 6 | 2200 |
| 7 | 2275 |
| 8 | 2350 |
| 9 | 2425 |
| 10 | 2500 |
| 11 | 2575 |
| 12 | 2650 |
| 13 | 2725 |
| 14 | 2800 |
| 15 | 2875 |
| 16 | 2950 |
| 17 | 3025 |
| 18 | 3100 |
| 19 | 3175 |
| 20 | 3250 |
| 21 | 3325 |
| 22 | 3400 |
| 23 | 3475 |
| 24 | 3550 |
| 25 | 3625 |
| 26 | 3700 |
| 27 | 3775 |
| 28 | 3850 |
| 29 | 3925 |
|  |  |

Table 3-1

| Far-Offset Data |  |
| :---: | :---: |
| Level | Deph |
| 30 | 4000 |
| 31 | 4075 |
| 32 | 4150 |
| 33 | 4225 |
| 34 | 4300 |
| 35 | 4375 |
| 36 | 4450 |
| 37 | 4525 |
| 38 | 4600 |
| 39 | 4675 |
| 40 | 4750 |
| 41 | 4825 |
| 42 | 4900 |
| 43 | 4975 |
| 44 | 5050 |
| 45 | 5125 |
| 46 | 5200 |
| 47 | 5275 |
| 48 | 5350 |
| 49 | 5425 |
| 50 | 5500 |
| 51 | 5575 |
| 52 | 5650 |

## FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE

| LEVEL | WINDOW | PHI | THETA | $F$ |
| :---: | :---: | :---: | :---: | :---: |
| 1 | 1 | 84.272 | 77.929 | 0.980 |
| 1 | 2 | 84.734 | 77.979 | 0.979 |
| 1 | 3 | 84.782 | 78.375 | 0.989 |
| 1 | 4 | 84.511 | 79.982 | 6.979 |
| 1 | 5 | 84.720 | 84.288 | -.97\% |
| 1 | 6 | 87.483 | 89.875 | 0.969 |
| 1 | 7 | 89.240 | -98. 142 | 0.983 |
| 1 | 8 | 89.080 | -92.833 | 0.992 |
| 1 | 9 | 89.683 | 87.296 | 0.984 |
| 1 | 10 | 88.637 | -92.369 | 0.958 |
| 1 | 11 | 84.882 | -91.556 | 0.955 |
| 1 | 12 | 86.161 | -90.662 | 0.984 |
| 1 | 13 | 89.446 | 90.450 | 0.929 |
| 1 | 14 | 86.566 | 91.698 | 0.787 |
| 1 | 16 | 88.348 | 92.183 | 0.653 |
| 1 | 18 | 82.681 | -88.765 | 0.408 |
| 1 | 17 | 87.289 | -87.381 | 0.505 |
| 2 | 1 | 83.687 | -102.498 | 0.974 |
| 2 | 2 | 83.417 | -103.259 | 0.984 |
| 2 | 3 | 83.309 | -103.292 | 0.990 |
| 2 | 4 | 83.148 | -163.113 | 0.994 |
| 2 | 5 | 82.886 | -103.118 | 0.996 |
| 2 | 6 | 82.694 | -103.366 | 0.994 |
| 2 | 7 | 82.822 | -163.178 | - . 993 |
| 2 | 8 | 83.493 | -189.968 | 0.989 |
| 2 | 9 | 84.568 | -97.165 | 0.998 |
| 2 | 10 | 85.086 | -96.815 | 0.996 |
| 2 | 11 | 04.676 | -96.846 | 0.996 |
| 2 | 12 | 83.688 | -98.262 | 0.992 |
| 2 | 13 | 82.625 | -99.690 | 0.986 |
| 2 | 14 | 82.837 | -99.361 | 0.979 |
| 2 | 15 | 81.978 | -99.469 | 0.974 |
| 2 | 18 | 82.858 | -188.214 | 0.978 |
| 2 | 17 | 02.064 | -101.558 | 0.984 |
| 2 | 18 | 83.099 | -182.793 | 0.956 |
| 2 | 19 | 87.538 | -182.869 | 0.931 |
| 2 | 29 | 88.950 | 78.046 | 0.946 |
| 2 | 21 | 86.155 | 78.811 | 0.968 |
| 3 | 1 | 84.444 | -98.847 | 0.994 |
| 3 | 2 | 84.642 | -97.016 | 0.988 |
| 3 | 3 | 84.541 | -96.719 | 0.982 |
| 3 | 4 | 84.52\% | -96.160 | 0.975 |
| 3 | 6 | 84.999 | -94.986 | 0.972 |
| 3 | 6 | 86.810 | -94.468 | 0.974 |
| 3 | 7 | 86.656 | -93.463 | 0.981 |
| 3 | 8 | 80.308 | -02.476 | 0. 989 |
| 3 | 9 | 85.601 | -91.726 | 0.990 |
| 3 | 15 | 84.591 | -91.269 | 0.988 |
| 3 | 11 | 83.931 | -91.170 | 0.986 |
| 3 | 12 | 84.140 | -91.168 | 0.980 |
| 3 | 13 | 84.678 | -98. 843 | 0. 977 |
| 3 | 14 | 84.736 | -94.747 | 0.975 |
| 3 | 15 | 84.813 | -91.487 | 0.976 |
| 3 | 10 | 87.557 | -94.358 | 0.951 |
| 3 | 17 | 85.648 | 82.858 | 0.841 |
| 4 | 1 | 87.680 | 81.997 |  |
| 4 | 2 | 07.429 | 82.362 | $0.894^{\circ}$ |

## FAR OFFSET SH $_{\mathrm{t}}$ SOURCE

| LEVEL | WINDOW | PHI | theta | F |
| :---: | :---: | :---: | :---: | :---: |
| 4 | 3 | 87.533 | 81.589 | 0.889 |
| 4 | 4 | 87.846 | . 86.644 | 0.902 |
| 4 | 5 | 87.485 | 81.530 | 0.924 |
| 4 | 8 | 86.982 | 83.887 | 0.937 |
| 4 | 7 | 86.929 | 85.287 | 0.938 |
| 4 | 8 | 87.611 | 85.844 | 0.939 |
| 4 | 9 | 88.369 | 83.747 | 0.951 |
| 4 | 10 | 87.956 | 83.899 | 0.989 |
| 4 | 11 | 86.783 | 85.443 | 0.988 |
| 4 | 12 | 85.573 | 87.349 | 8.983 |
| 4 | 13 | 84.193 | 87.116 | ©. 988 |
| 4 | 14 | 80.119 | 82.868 | 6.947 |
| 5 | 1 | 83.896 | 78.978 | 6. 979 |
| 5 | 2 | 84.117 | 77.142 | 9.979 |
| 5 | 3 | 84.544 | 77.786 | 0.989 |
| 5 | 4 | 84.711 | 78.402 | 0.981 |
| 5 | 5 | 84.966 | 78.888 | 0.988 |
| 5 | 6 | 86.148 | 79.338 | 6.978 |
| 5 | 7 | 88.642 | 79.513 | 0.986 |
| 5 | 8 | 89.646 | -160.739 | 0.986 |
| 5 | 9 | 88.533 | -160.646 | 0.979 |
| 5 | 10 | 89.335 | -99.857 | 0.974 |
| 8 | 11 | 89.894 | 79.444 | 0.975 |
| 6 |  | 72.953 | 78.626 | 0.952 |
| 6 | 2 | -72.395 | 78.675 | 0.947 |
| 6 | 3 | 72.372 | 78.091 | 0.946 |
| 6 | 4 | 73.858 | 77.971 | 0.925 |
| 6 | 5 | 75.536 | 77.293 | 0.885 |
| 6 | - | 82.162 | 70.320 | 0.821 |
| 6 | 7 | 80.613 | -103.206 | 0.798 |
| 6 | 8 | 75.360 | -101.284 | 0.873 |
| 6 | 9 | 70.722 | -99.886 | 0.965 |
| 6 | 16 | 68.412 | -160.384 | 0.976 |
| 7 | 1 | 56.803 | 134.898 | 0.928 |
| 7 | 2 | 67.360 | 138.893 | 0.930 |
| 7 | 3 | 56.839 | 144.261 | 0.874 |
| 7 | 4 | 64.744 | 149.848 | 0.783 |
| 7 | 5 | 61.084 | 165.343 | 0.736 |
| 7 | 8 | 45.761 | 161.567 | 0.732 |
| 7 | 7 | 38.332 | 171.160 | 0.229 |
| 7 | 8 | 63.460 | 29.331 | 0. 426 |
| 7 | 9 | 67.185 | 29.485 | 0.574 |
| 7 | 18 | 56.768 | 33.613 | 0.636 |
| 7 | 11 | 38.869 | 55.886 | 0.689 |
| 7 | 12 | 37.437 | 71.746 | 0.818 |
| 7 | 13 | 39.982 | 89.644 | 0.822 |
| 7 | 14 | 41.820 | 88.117 | 0.884 |
| 7 | 15 | 42.281 | 96.888 | 0.822 |
| 7 | 18 | 43.176 | 98.171 | 0.881 |
| 7 | 17 | 45.949 | 97.362 | 0.771 |
| 7 | 10 | 54.488 | 99.448 | 0.395 |
| 8 | 1 | 48.978 | 85.554 | 0.641 |
| 8 | 2 | 56.217 | 86.767 | 0.742 |
| 8 | 3 | 61.626 | 89.132 | 6.897 |
| 8 | 4 | 69.824 | 96.177 | 0.947 |
| 8 | 5 | 57.337 | 91.777 | 0.958 |

## FAR OFFSET SH ${ }_{t}$ SOURCE

| LEVEL | WINDOW | PHI | THETA | F |
| :---: | :---: | :---: | :---: | :---: |
| 8 | 6 | 56.292 | 93.380 | 6.938 |
| 8 | 7 | 57.540 | 93.759 | 0.893 |
| 8 | 8 | 60.111 | 93.699 | 0.885 |
| 8 | 9 | 61.495 | 92.252 | 0.849 |
| 8 | 16 | 60.264 | 91.665 | 0.868 |
| 9 | 1 | 46.319 | 144.824 | 0.612 |
| 9 | 2 | 53.182 | 184.925 | 8.515 |
| 9 | 3 | 56.169 | 91.425 | 0.886 |
| 9 | 5 | 68.612 | 88.496 | 0.701 |
| 9 | 6 | 69.246 | 96.855 | 0.887 |
| 9 | 6 | 68.238 | 92.876 | 0.951 |
| 9 | 7 | 64.212 | 94.733 | 0.968 |
| 9 | 8 | 61.365 | 96.509 | 0.898 |
| $\bigcirc$ | 9 | 62.453 | 96.996 | 0.789 |
| $\bigcirc$ | 18 | 67.548 | 95.581 | 0.712 |
| $\bigcirc$ | 11 | 71.684 | 94.146 | 0. 699 |
| 9 | 12 | 76.412 | 93.643 | 0.873 |
| 9 | 13 | 69.621 | 92.009 | 0.628 |
| 9 | 14 | 77.684 | -187.970 | 0.386 |
| 9 | 15 | 68.982 | -116.125 | 6.813 |
| 10 | 1 | 71.785 | 98.178 | 6.997 |
| 10 | 2 | 72.744 | . 98.329 | 0.998 |
| 16 | 3 | 72.723 | 98.517 | 0.998 |
| 10 | 4 | 71.586 | 99.181 | 6.993 |
| 10 | 6 | 71.829 | 99.219 | - . 959 |
| 16 | 6 | 77.481 | 98.523 | 0.914 |
| 10 | 7 | 01.876 | 93.846 | 0.937 |
| 16 | 8 | 80.737 | 93.918 | -. 956 |
| 18 | 9 | 77.898 | 98.115 | 0.961 |
| 10 | 10 | 75.606 | 96.211 | 0.930 |
| 10 | 11 | 75.616 | 96.464 | 0.898 |
| 16 | 12 | 76.829 | 98.066 | 0.879 |
| 16 | 13 | 76.258 | 96.266 | 0.877 |
| 10 | 14 | 74.568 | 97.646 | 0.870 |
| 10 | 16 | 71.739 | 99.907 | 0.829 |
| 10 | 16 | 09.994 | 101.341 | 0.719 |
| 18 | 17 | 74.876 | 97.540 | - . 583 |
| 10 | 18 | 78.790 | 93.930 | 0.621 |
| 10 | 19 | 73.681 | 95.186 | 0.699 |
| 11 | 1 | 75.545 | 94.462 | 0.998 |
| 12 | 2 | 74.536 | 95.945 | 0.996 |
| 11 | 3 | 73.132 | 97.222 | 0.985 |
| 11 | 4 | 72.934 | 97.362 | 0.952 |
| 11 | 5 | 70.625 | 94.749 | 0.982 |
| 11 | 6 | 01.032 | 91.669 | 0.897 |
| 11 | 7 | 03.678 | 09.346 | 0.919 |
| 11 | 8 | 83.192 | 89.226 | 0.929 |
| 11 | 9 | 81.698 | 89.760 | 0.912 |
| 11 | 10 | 81.464 | 98.271 | 0.881 |
| 11 | 11 | 86.273 | 94.314 | 0.788 |
| 11 | 12 | 86.821 | 96.148 | 0.724 |
| 11 | 13 | 79.567 | 99.769 | 0.670 |
| 11 | 14 | 75.601 | 93.343 | 0.692 |
| 11 | 15 | 65.988 | 160.290 | 6. 478 |
| 11 | 10 | 54.533 | 110.779 | 0.386 |
| 11 | 17 | 64.446 | 110.842 | 0.318 |
| 11 | 18 | 66.421 | 98.998 | 6. 467 |

## FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE

| LEVEL | WINDOW | PHI | THETA | $F$ |
| :---: | :---: | :---: | :---: | :---: |
| 11 | 19 | 88.828 | 98.491 | 0.844 |
| 12 | 1 | 78.734 | 97.889 | 8.994 |
| 12 | 2 | 77.531 | 97.755 | 0.998 |
| 12 | 3 | 78.368 | 97.764 | 8.992 |
| 12 | 4 | 77.266 | 97.375 | 6. 989 |
| 12 | 5 | 82.238 | 96.394 | 0.948 |
| 12 | 6 | 86.521 | 95.196 | 6. 962 |
| 12 | 7 | 87.828 | 93.736 | 0. 983 |
| 12 | 8 | 87.866 | 91.815 | 0. 984 |
| 12 | 9 | 87.273 | 89.852 | 0.938 |
| 12 | 16 | 87.918 | 88.982 | 0.876 |
| 12 | 11 | 89.318 | 88.974 | 0.822 |
| 12 | 12 | 89.839 | -02.212 | 0.794 |
| 12 | 13 | 89.879 | 81.983 | 6.752 |
| 12 | 14 | 86.448 | 68.942 | 0.888 |
| 12 | 15 | 81.441 | 62.298 | 0.868 |
| 12 | 16 | 79.888 | 47.768 | 6. 683 |
| 13 | 1 | 78.117 | 67.432 | 0.942 |
| 13 | 2 | 76.039 | 64.898 | 0.080 |
| 13 | 3 | 74.451 | 60.819 | 6.974 |
| 13 | 4 | 73.291 | 58.925 | 6.982 |
| 13 | 6 | 76.383 | 69.942 | 0.929 |
| 13 | 6 | 81.898 | 61.793 | 0.928 |
| 13 | 7 | 85.838 | 61.887 | 0.938 |
| 13 | 8 | 86.847 | 61.545 | 0.949 |
| 13 | 9 | 86.341 | 61.869 | 0.944 |
| 13 | 16 | 87.468 | 02.642 | 0.903 |
| 13 | 11 | 89.897 | -116.533 | 0.824 |
| 13 | 12 | 88.873 | -110.812 | 8.788 |
| 13 | 13 | 87.134 | -117.768 | 0.719 |
| 13 | 14 | 88.462 | -118.205 | 0.571 |
| 13 | 18 | 47.887 | 72.516 | 0.863 |
| 13 | 18 | 18.528 | 116.483 | 0.481 |
| 13 | 27 | 37.136 | 93.887 | 6.358 |
| 14 | 1 | 77.931 | 88.915 | 0.918 |
| 14 | 2 | 73.984 | 86.264 | 0.916 |
| 14 | 3 | 76.496 | 82.808 | 0.916 |
| 14 | 4 | 67.601 | 80.746 | 0.903 |
| 14 | 5 | 88.982 | 79.186 | 0.860 |
| 14 | 8 | 73.334 | 88.896 | 6. 748 |
| 14 | 7 | 86.180 | 84.496 | 0.756 |
| 14 | 8 | 89.872 | 83.849 | 6. 836 |
| 14 | 9 | 89.147 | 89.641 | 6.889 |
| 14 | 18 | 88.283 | 76.746 | 0. 834 |
| 14 | 11 | 83.288 | 74.143 | 0.748 |
| 14 | 12 | 82.916 | 73.864 | e. 0.43 |
| 14 | 13 | 84.982 | 73.584 | 0.888 |
| 14 | 14 | 86.373 | 6 0. 818 | 0.628 |
| 14 | 15 | 83.521 | 64. 449 | 0.861 |
| 14 | 10. | 82.534 | 53.473 | 0. 062 |
| 14 | 17 | 84.896 | 61.834 | 0. 654 |
| 16 | 1 | 88.788 | -97.420 | 0.998 |
| 15 | 2 | 79.888 | -97.496 | 0.988 |
| 16 | 3 | 79.808 | -97. 548 | 0. 982 |
| 16 | 4 | 77.853 | -97.836 | 0.972 |
| 16 | 5 | 77.832 | -98.852 | 0.939 |

## FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE

| LEVEL | WINDOW | PHI | theta | $F$ |
| :---: | :---: | :---: | :---: | :---: |
| 15 | 6 | 83.107 | -99.449 | 0.910 |
| 15 | 7 | 86.819 | -97.687 | 0.940 |
| 16 | 8 | 85.133 | -96.828 | 0.957 |
| 16 | 9 | 82.042 | -94.763 | 0.962 |
| 15 | 10 | 79.646 | -94.174 | 6. 936 |
| 15 | 11 | 78.449 | -94.875 | 0.986 |
| 16 | 12 | 78.185 | -94.066 | 0.894 |
| 15 | 13 | 77.003 | -93.619 | 0.896 |
| 15 | 14 | 73.667 | -92.684 | 0.896 |
| 15 | 15 | 68.949 | -91.773 | 0.888 |
| 15 | 16 | 67.324 | -92.363 | 0.885 |
| 15 | 17 | 78.997 | -84.029 | 0.888 |
| 16 | 18 | 73.844 | -94.819 | 0.913 |
| 16 | 1 | 71.056 | 96.483 | 0.982 |
| 10 | 2 | 68.432 | 89.519 | 0.915 |
| 16 | 3 | 06.431 | 89.276 | 0.864 |
| 18 | 4 | 84.416 | 89.614 | 0.824 |
| 16 | 5 | 61.175 | 88.905 | 0.780 |
| 16 | 6 | 56.855 | 82.760 | 0.622 |
| 18 | 7 | 64.555 | 67.086 | 6. 468 |
| 16 | 8 | 72.945 | 65.931 | 0.595 |
| 10 | 9 | 69.460 | 73.735 | 0.767 |
| 18 | 10 | 63.592 | 86.884 | 0.748 |
| 16 | 11 | 54.324 | 84.931 | 0.746 |
| 18 | 12 | 54.387 | 89.181 | 0.721 |
| 16 | 13 | 51.776 | 02.710 | 0.683 |
| 16 | 14 | 58.194 | 95.226 | 0.648 |
| 18 | 15 | 49.184 | 97.513 | 0.623 |
| 16 | 16 | 48.971 | 180.388 | 0.697 |
| 18 | 17 | 49.950 | 161.796 | 0.547 |
| 16 | 18 | 53.217. | 97.693 | 0.474 |
| 17 | 1 | 72.463 | 83.744 | 8.971 |
| 17 | 2 | 69.031 | 81.315 | 0.943 |
| 17 | 3 | 66.615 | 79.588 | 0.884 |
| 17 | 4 | 86.354 | 79.858 | 0.812 |
| 17 | 5 | 68.415 | 78.285 | 0.757 |
| 17 | 8 | 71.143 | 74.718 | 0.729 |
| 17 | 7 | 72.573 | 88.867 | 0.718 |
| 17 | 8 | 71.447 | 62.491 | 8.725 |
| 17 | 9 | 08.249 | 62.184 | 0.753 |
| 17 | 10 | 64.761 | 04.884 | 0.797 |
| 17 | 11 | 61.314 | 86.916 | 0.831 |
| 17 | 12 | 68.157 | 67.718 | 0.840 |
| 17 | 13 | 56.231 | 67.668 | 0.820 |
| 17 | 14 | 57.258 | 07.428 | 0.782 |
| 17 | 15 | 68.846 | 68.389 | 0.751 |
| 17 | 18 | 84.891 | 89.982 | 0.737 |
| 17 | 17 | 09.120 | 76.646 | 0.723 |
| 10 | 1 | 72.224 | 75.631 | 0.947 |
| 18 | 2 | 76.653 | 73.108 | 0.928 |
| 18 | 3 | 76.949 | 73.250 | 0.872 |
| 18 | 4 | 73.813 | 74.976 | 0.839 |
| 18 | 5 | 76.843 | 78.290 | 0.844 |
| 18 | 6 | 77.748 | 73.631 | 0.857 |
| 18 | 7 | 76.466 | 76.630 | 0.858 |
| 18 | 8 | 73.883 | 67.643 | 0.852 |
| 18 | 0 | 71.888 | 64.478 | 0.846 |

## $\because$ FAR OFFSET SH ${ }_{t}$ SOURCE

| LEVEL | WINOOW | PHI | ThETA | F |
| :---: | :---: | :---: | :---: | :---: |
| 18 | 10 | 70.816 | 61.784 | 0.847 |
| 18 | 11 | 69.961 | 00.200 | 0.849 |
| 18 | 12 | 89.103 | 59.724 | 0.888 |
| 18 | 13 | 69.944 | 57.235 | 0.896 |
| 18 | 14 | 73.103 | 52.058 | 0.914 |
| 18 | 15 | 77.580 | 46.784 | 0.917 |
| 19 | 1 | 89.873 | 91.458 | 0.964 |
| 19 | 2 | 68.684 | 86.985 | 0.968 |
| 19 | 3 | 86.128 | 85.453 | 0.948 |
| 19 | 4 | 68.946 | 84.272 | 0.919 |
| 19 | 5 | 77.423 | 79.973 | 0. 905 |
| 19 | 6 | 79.815 | 78.138 | 0.938 |
| 19 | 7 | 78.844 | 79.069 | 0.946 |
| 19 | 8 | 75.981 | 80.760 | 0.968 |
| 19 | 9 | 76.320 | 81.740 | 0.816 |
| 19 | 18 | 79.223 | 88.786 | 8.744 |
| 19 | 11 | 89.834 | 79.698 | 0.746 |
| 19 | 12 | 76.565 | 81.679 | 0.734 |
| 19 | 13 | 67.811 | 89.502 | 0.646 |
| 19 | 14 | 53.638 | 114.719 | 0. 561 |
| 20 | 1 | 64.717 | 180.211 | 0.988 |
| 20 | 2 | 66.288 | 93.893 | 0.965 |
| 20 | 3 | 86.322 | 90.842 | 0.963 |
| 28 | 4 | 65.271 | 91.234 | 0.962 |
| 20 | 5 | 71.603 | 84.622 | 0.861 |
| 20 | 6 | 79.236 | 75.380 | 0.899 |
| 20 | $?$ | 78.188 | 76.296 | 0.942 |
| 29 | 8 | 76.200 | 80.179 | 0.906 |
| 20 | 9 | 73.002 | 82.379 | 0.757 |
| 20 | 10 | 76.622 | 78.630 | 6.685 |
| 20 | 11 | 79.687 | 74.629 | 0.657 |
| 29 | 12 | 78.429 | 75.196 | 0.566 |
| 20 | 13 | 72.838 | 82.681 | 6.467 |
| 20 | 14 | 58.386 | 123.322 | 6.329 |
| 21 | 1 | 68.716 | 99.799 | 0.981 |
| 21 | 2 | 66.107 | 96.604 | 0.972 |
| 21 | 3 | 65.461 | 91.686 | 6.963 |
| 21 | 4 | 65.684 | 01.512 | 0.947 |
| 21 | 5 | 69.594 | 86.968 | 0.934 |
| 21 | 6 | 72.688 | 77.624 | 0.961 |
| 21 | 7 | 76.898 | 77.035 | 0.982 |
| 21 | 8 | 68.594 | 79.774 | 0.962 |
| 21 | 9 | 68.148 | 80.784 | 8.876 |
| 21 | 10 | 69.918 | 76.138 | 0.784 |
| 21 | 11 | 76.842 | 73.616 | 0.783 |
| 21 | 12 | 09.992 | 73.978 | 0.783 |
| 21 | 13 | 67.946 | 76.237 | 0.703 |
| 21 | 14 | 64.963 | 01.784 | 0.446 |
| 22 | 1 | 65.928 | 183.317 | 0.991 |
| 22 | 2 | 66.193 | 182.893 | 6.994 |
| 22 | 3 | 06.391 | 99.836 | 0.978 |
| 22 | 4 | 60.298 | 96.741 | 0.941 |
| 22 | 5 | 67.134 | 96.317 | 0.926 |
| 22 | 6 | 68.739 | 96.686 | 0.946 |
| 22 | 7 | 08.082 | 84.648 | 0.971 |
| 22 | - | 60.894 | 84.637 | 0.984 |

## FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE

| LEVEL | WINDOW | P.HI | theta | $F$ |
| :---: | :---: | :---: | :---: | :---: |
| 22 | 9 | 68.363 | 86.809 | 0.968 |
| 22 | 10 | 66.318 | 85.697 | 0.898 |
| 22 | 11 | 66.188 | 81.386 | 0.861 |
| 22 | 12 | 66.278 | 86.728 | 0.875 |
| 22 | 13 | 86.481 | 81.878 | 0.857 |
| 22 | 14 | 86.179 | 83.037 | 0.749 |
| 22 | 15 | 66.775 | 85.061 | 0.434 |
| 22 | 16 | 72.376 | 4.683 | 0.141 |
| 23 | 1 | 74.678 | 186.237 | 0.998 |
| 23 | 2 | 74.691 | 185.671 | 0.994 |
| 23 | 3 | 74.686 | 162.462 | 0.958 |
| 23 | 4 | 74.369 | 189.446 | 0.888 |
| 23 | 5 | 71.874 | 103.856 | 0.844 |
| 23 | 6 | 87.365 | 95.477 | 0.986 |
| 23 | 7 | 66.534 | 87.611 | 0.946 |
| 23 | 8 | 67.414 | 86.662 | 0.962 |
| 23 | 9 | 68.693 | 87.542 | 0.942 |
| 23 | 10 | 66.836 | 84.146 | 0.867 |
| 23 | 11 | 66.601 | 79.871 | 8.856 |
| 23 | 12 | 66.365 | 79.936 | 0.867 |
| 23 | 13 | 67.295 | 79.489 | - 0.798 |
| 23 | 14 | 67.461 | 75.932 | 0.647 |
| 23 | 15 | 61.981 | 13.883 | 0.166 |
| 24 | 1 | 02.765 | 112.683 | 0.956 |
| 24 | 2 | 03.047 | 113.080 | 0.977 |
| 24 | 3 | 63.884 | 111.172 | 0.982 |
| 24 | 4 | 05.228 | 189.218 | 0.929 |
| 24 | 6 | 66.180 | 111.484 | 0.868 |
| 24 | 0 | 68.652 | 119.248 | 0.783 |
| 24 | 7 | 52.397 | 111.361 | 0.828 |
| 24 | 8 | 64.903 | 96.446 | 6.871 |
| 24 | 9 | 59.804 | 92.889 | 0.962 |
| 24 | 10 | 62.820 | 92.579 | 0.916 |
| 24 | 11 | 62.894 | 93.136 | 0.989 |
| 24 | 12 | 61.098 | 89.969 | 0.856 |
| 24 | 13 | 61.047 | 87.243 | 0.860 |
| 24 | 14 | 63.277 | 87.299 | 0.846 |
| 24 | 15 | 66.188 | 96.425 | 0.739 |
| 24 | 16 | 72.209 | 108.713 | 0.496 |
| 24 | 17 | 86.392 | 138.098 | 0.463 |
| 25 | 1 | 69.262 | 189.192 | 0.967 |
| 25 | 2 | 76.808 | 109.671 | 0.977 |
| 25 | 3 | 71.146 | 108.513 | 0.978 |
| 25 | 4 | 71.122 | 108.698 | 0.916 |
| 25 | 5 | 65.682 | 116.258 | 0.822 |
| 26 | 6 | 58.296 | 119.190 | 0.873 |
| 28 | 7 | 68.689 | 110.392 | 0.898 |
| 25 | 8 | 03.813 | 102.273 | 0.890 |
| 26 | - | 88.848 | 180.320 | 0.803 |
| 28 | 10 | 76.032 | 101.647 | 0.868 |
| 25 | 11 | 09.278 | 99.104 | 0.846 |
| 25 | 12 | 69.102 | 96.134 | 0.839 |
| 25 | 13 | 72.893 | 93.433 | 0.837 |
| 25 | 14 | 77.096 | 96.382 | 0.764 |
| 26 | 15 | 83.818 | 107.898 | 0.671 |
| 28 | 1 | 64.493 | 114.692 | -. 968 |

## FAR OFFSET SH $_{t}$ SOURCE

| LEVEL | WINDOW | PHI | THETA | $F$ |
| :---: | :---: | :---: | :---: | :---: |
| 28 | 2 | 85.945 | 114.332 | - 0.978 |
| 28 | 3 | 67.869 | 113.385 | 0.989 |
| 28 | 4 | 69.631 | 112.256 | 0.912 |
| 26 | 5 | 88.838 | 113.633 | 0.774 |
| 26 | 6 | 62.031 | 121.934 | 0.651 |
| 26 | 7 | 57.127 | 127.199 | 0.769 |
| 26 | 8 | 61.238 | 119.027 | 0.755 |
| 26 | 9 | 69.017 | 110.047 | ¢. 770 |
| 26 | 18 | 74.812 | 106.512 | 0.775 |
| 26 | 11 | 76.841 | 106.628 | 0.776 |
| 28 | 12 | 76.359 | 166.757 | 0.899 |
| 26 | 13 | 77.031 | 106.384 | 0.873 |
| 26 | 14 | 79.861 | 184.691 | 0.923 |
| 26 | 15 | 83.347 | 186.714 | 6.911 |
| 26 | 16 | 84.809 | 116.178 | 0.754 |
| 27 | 1 | 68.436 | 118.183 | 0.928 |
| 27 | 2 | 62.364 | 114.692 | 0.947 |
| 27 | 3 | 84.581 | 111.958 | 0.946 - |
| 27 | 4 | 67.487 | 110.746 | 0.928 - $0^{\text {a }}$ |
| 27 | 5 | 69.458 | 110.010 | 0.874 - 0.8 |
| 27 | 6 | 69.877 | 189.208 | 0.738 |
| 27 | 7 | 63.281 | 112.329 | 0.627 |
| 27 | 8 | 58.891 | 125.754 | 0.512 |
| 27 | 9 | 65.578 | 125.651 | . .683 |
| 27 | 18 | 69.949 | 112.214 | 0.609 |
| 27 | 11 | 82.642 | 182.463 | c. 641 |
| 27 | 12 | 87.964 | 98.198 | 0.667 |
| 27 | 13 | 88.289 | 98.859 | 0.693 |
| 27 | 14 | 87.481 | 97.216 | ©. 764 |
| 27 | 15 | 88.312 | 93.849 | 0.869 |
| 27 | 16 | 89.728 | -87.034 | $0.807]$ |
| 27 | 17 | 87.900 | -85.120 | $0.725] \square<$ |
| 28 | 1 | 54.240 | 116.689 | 0.861 |
| 28 | 2 | 58.870 | 110.829 | 0.851 |
| 28 | 3 | 63.479 | 187.236 | 0.895 D |
| 28 | 4 | 66.308 | 104.617 | 0.860 D |
| 28 | 5 | 61.883 | 165.164 | 0.378 |
| 28 | 6 | 34.887 | 133.386 | 0.274 |
| 28 | 7 | 34.721 | 144.568 | 0.421 |
| 28 | 8 | 47.778 | 128.222 | 0.374 |
| 28 | 9 | 86.169 | 182.661 | 0.364 |
| 28 | 18 | 86.193 | -87.981 | 0.520 |
| 28 | 11 | 84.634 | -88.504 | 0.685 |
| 28 | 12 | 86.695 | -87.663 | 0.647 D |
| 28 | 13 | 87.783 | -90.031 | 0.761 - |
| 28 | 14 | 87.468 | -92.769 | 0.878 |
| 28 | 15 | 86.450 | -92.530 | 0.868 ] |
| 28 | 16 | 85.314 | -92.625 | 0.490 |
| 29 | 1 | 80.846 | 104.989 | 0.919 |
| 29 | 2 | 64.684 | 99.367 | 0.828 D |
| 29 | 3 | 66.565 | 97.649 | 0.050 D |
| 29 | 4 | 58.364 | 103.342 | 6.417 - |
| 29 | 5 | 46.338 | 126.607 | 6.395 |
| 29 | 6 | 36.878 | 125.097 | 0.469 |
| 29 | 7 | 46.262 | 120.666 | 0.427 |
| 29 | 8 | 02.472 | 104.466 | 0.217 |
| 29 | 0 | 79.161 | -93.672 | 0. 346 |

## FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE



## FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE

| LEVEL | WINDOW | PHI | THETA | F |
| :---: | :---: | :---: | :---: | :---: |
| 33 | 10 | 56.718 | -55.757 | 0.495 |
| 33 | 11 | 66.385 | -66.514 | 0.754 |
| 33 | 12 | 64.981 | -53.989 | 0.919 |
| 33 | 13 | 65.512 | -51.481 | 0.914 |
| 34 | 1 | 72.232 | 97.129 | 0.598 |
| 34 | 2 | 88.408 | -93.485 | 0.481 |
| 34 | 3 | 85.217 | -166.777 | 0.367 |
| 34 | 4 | 49.178 | -121. 612 | 0. 342 |
| 34 | 5 | 45.291 | -127.253 | 0.325 |
| 34 | 6 | 48.252 | -121.379 | 0.373 |
| 34 | 7 | 47.292 | -117.498 | 0.504 |
| 34 | 8 | 43.262 | -118.415 | 0.584 |
| 34 | 9 | 48.487 | -120.367 | 0.542 |
| 34 | 10 | 45.812 | -116.987 | 0.365 |
| 34 | 11 | 79.474 | -88.624 | 0.504 |
| 34 | 12 | 71.141 | -87.486 | 0.771 |
| 34 | 13 | 67.161 | -91.826 | 0.698 |
| 35 | 1 | 60.361 | -116.836 | 0.374 |
| 35 | 2 | 46.411 | -132.641 | 0.390 |
| 35 | 3 | 41.985 | -142.317 | 0.366 |
| 36 | 4 | 45.194 | -136.868 | 0.346 |
| 36 | 6 | 47.444 | -131.836 | 0.486 |
| 36 | 6 | 44.386 | -136.964 | 0.646 |
| 35 | 7 | 41.418 | -146.749 | 0.717 |
| 35 | 8 | 45.727 | -163.779 | 0.661 |
| 36 | 9 | 46.675 | -146.883 | c. 395 |
| 36 | 1 | 00.471 | -93.241 | 0.491 |
| 36 | 2 | 63.681 | -108.089 | 0.574 |
| 36 | 3 | 47.818 | -114.444 | 0.583 |
| 36 | 4 | 48.112 | -114.579 | 0.568 |
| 36 | 5 | 61.861 | -108.996 | 0. 688 |
| 36 | 8 | 51.838 | -168.618 | 0.713 |
| 36 | 7 | 48.765 | -107.761 | 0.767 |
| 36 | 8 | 45.162 | -109.441 | 0.760 |
| 36 | 9 | 46.156 | -103.789 | e. 52\% |
| 37 | 1 | 62.746 | -111.138 | 0.562 |
| 37 | 2 | 56.826 | -121.338 | 0.585 |
| 37 | 3 | 63.103 | -126.843 | 0. 596 |
| 37 | 4 | 62.906 | -127.802 | 0.618 |
| 37 | 6 | 63.114 | -126.433 | 0.682 |
| 37 | 6 | 51.768 | -127.238 | 0.743 |
| 37 | 7 | 49.112 | -129.697 | 0.745 |
| 38 | 1 | 27.343 | -147.773 | 0.431 |
| 38 | 2 | 39.924 | -137.333 | 0.831 |
| 36 | 3 | 33.432 | -141.674 | 0.618 |
| 38 | 4 | 36.236 | -147.262 | 0. 689 |
| 38 | 5 | 38.290 | -168.835 | 0.709 |
| 38 | 6 | 38.910 | -153.382 | 0.731 |
| 38 | 7 | 38.773 | -158.019 | 0.711 |
| 38 | 8 | 39.327 | -166.328 | 0.811 |
| 39 | 1 | 41.443 | -180.176 | 0.784 |
| 39 | 2 | 41.541 | -140.361 | 0.739 |
| 39 | 3 | 42.858 | -141.463 | 8.776 |
| 39 | 4 | 44.316 | -146.497 | 0.822 |

## FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE

| LEVEL | WINDOW | PHI | Theta | F |
| :---: | :---: | :---: | :---: | :---: |
| 39 | 5 | 45.810 | -141.513 | 0.853 |
| 39 | 6 | 46.938 | -142.347 | 0.877 |
| 39 | 7 | 47.272 | -143.190 | 0.909 |
| 39 | 8 | 46.948 | -144.966 | 0.915 |
| 39 | 9 | 46.688 | -147.998 | 0.913 |
| 39 | 10 | 46.318 | -151.771 | 0.892 |
| 40 | 1 | 37.621 | -174.248 | 0.449 |
| 40 | 2 | 46.394 | 154.624 | 0.423 |
| 46 | 3 | 56.334 | 136.988 | 0. 222 |
| 40 | 4 | 69.443 | -120.564 | -. 330 |
| 48 | 5 | 59.378 | -130.484 | 8.623 |
| 46 | 6 | 57.833 | -136.614 | ¢. 725 |
| 46 | 7 | 56.535 | -137.823 | 0.753 |
| 40 | 8 | 66.636 | -137.035 | 0.765 |
| 46 | 9 | 67.388 | -134.651 | 0.780 |
| 40 | 10 | 67.164 | -134.719 | 0.786 |
| 40 | 11 | 65.251 | -137.865 | 0.738 |
| 41 | 1 | 67.227 | -96.168 | ¢. 159 |
| 41 | 2 | 86.616 | -87.037 | 0.436 |
| 41 | 3 | 80.371 | -94.436 | ¢. 599 |
| 41 | 4 | 72.807 | -184.566 | 0.859 |
| 41 | 5 | 03.047 | -114.436 | 0.069 |
| 41 | 6 | 57.690 | -120.886 | 0.867 |
| 41 | 7 | 54.824 | -123.632 | 0.060 |
| 41 | 8 | 65.146 | -124.251 | 0.671 |
| 41 | 9 | 55.440 | -126.138 | 0.721 |
| 41 | 18 | 53.194 | -126.822 | 0.744 |
| 42 | 1 | 70.358 | -126.941 | ¢. 392 |
| 42 | 2 | 77.383 | -126.472 | 0.638 |
| 42 | 3 | 72.168 | -127.992 | 0.786 |
| 42 | 4 | 67.097 | -133.894 | 0.831 |
| 42 | 5 | 63.597 | -136.378 | 0.828 |
| 42 | 6 | 83.299 | -136.595 | ©. 810 |
| 42 | 7 | 65.893 | -135.642 | 0.823 |
| 42 | 8 | 66.474 | -136.288 | 0.856 |
| 42 | $\bigcirc$ | 64.898 | -136.281 | 0.866 |
| 43 | 1 | 75.598 | -64.740 | 0.774 |
| 43 | 2 | 73.903 | -61.586 | ©. 877 |
| 43 | 3 | 76.139 | -69.022 | 0.915 |
| 43 | 4 | 67.613 | -67.769 | 0.966 |
| 43 | 5 | 67.410 | -57.248 | 0.872 |
| 43 | 8 | 68.494 | -56.382 | $0.86 \%$ |
| 43 | 7 | 67.889 | -52.255 | 0.864 |
| 43 | 8 | 64.873 | -46.915 | 6.867 |
| 44 | 1 | 74.513 | 183.687 | 0.749 |
| 44 | 2 | 77.841 | 153.849 | 0.680 |
| 44 | 3 | 82.738 | 144.615 | 0.498 |
| 44 | 4 | 07.813 | -53.661 | 0.289 |
| 44 | 5 | 74.685 | -78.856 | 6.176 |
| 44 | $\bigcirc$ | 68.767 | -91.681 | 0. 164 |
| 44 | 7 | 70.725 | -76.367 | 0.181 |
| 44 | 8 | 70.667 | -89.368 | 0.396 |
| 44 | 9 | 76.635 | -81.564 | 0.488 |
| 44 | 16 | 78.883 | -73.410 | 0.439 |

## FAR OFFSET SH ${ }_{\mathrm{t}}$ SOURCE

| LEVEL | WINDOW | PHI | Theta | $F$ |
| :---: | :---: | :---: | :---: | :---: |
| 46 | 1 | 62.625 | -143.842 | 0. 540 |
| 45 | 2 | 73.176 | -136.296 | 0.689 |
| 46 | 3 | 72.946 | -132.056 | 0.812 |
| 45 | 4 | 78.185 | -136.169 | 0.854 |
| 45 | 5 | 87.353 | -138.871 | 0.847 |
| 45 | 8 | 88.721 | -138.736 | 0.811 |
| 46 | 7 | 89.416 | -136.198 | 6. 799 |
| 46 | 8 | 71.816 | -134.980 | 0.836 |
| 46 | 9 | 89.696 | -136.783 | 0.848 |
| 46 | 10 | 67.369 | -137.397 | 0.824 |
| 46 | 1 | 75.488 | 173.749 | 6.427 |
| 48 | 2 | 87.478 | -98.887 | 0.691 |
| 46 | 3 | 76.992 | -181.377 | 0.736 |
| 48 | 4 | 71.835 | -183.453 | 0.869 |
| 46 | 5 | 68.176 | -167.677 | 0.823 |
| 46 | 6 | 69.824 | -104.278 | 0.714 |
| 46 | 7 | 78.582 | -100.462 | 0.739 |
| 46 | 8 | 68.780 | -188.856 | 0.775 |
| 47 | 1 | 89.792 | 72.873 | 0.623 |
| 47 | 2 | 78.532 | -122.459 | 0.676 |
| 47 | 3 | 72.868 | -128.645 | 0.753 |
| 47 | 4 | 70.082 | -138.047 | 0.778 |
| 47 | 6 | 72.122 | -129.661 | 6.772 |
| 47 | 0 | 75.840 | -125.819 | 0.823 |
| 47 | 7 | 78.639 | -126.273 | 0.868 |
| 47 | 8 | 74.607 | -126.868 | - 0.81 |
| 48 | 1 | 78.612 | 99.114 | 0.680 |
| 48 | 2 | 86.297 | 98.910 | 0.742 |
| 48 | 3 | 85.298 | -94. 213 | 0.863 |
| 48 | 4 | 78.388 | -96.898 | 0.833 |
| 48 | 6 | 74.724 | -95.381 | 0.838 |
| 48 | 6 | 75.460 | -97.143 | 0.832 |
| 48 | 7 | 77.138 | -102.495 | 0.849 |
| 48 | 8 | 75.778 | -182.834 | 0.878 |
| 48 | 9 | 73.295 | -97.066 | 0.848 |
| 48 | 10 | 71.880 | -85.366 | 0.682 |
| 48 | 11 | 76.921 | -52.910 | 0.383 |
| 49 | 1 | 77.846 | 162.936 | 0.728 |
| 49 | 2 | 89.283 | 89.873 | 0.613 |
| 49 | 3 | 79.789 | -99.963 | 0.594 |
| 49 | 4 | 75.112 | -108.866 | 0.026 |
| 49 | 5 | 74.251 | -102.641 | 0.686 |
| 49 | 6 | 72.947 | -110.321 | 0.797 |
| 49 | 7 | 70.973 | -113.896 | 0.888 |
| 49 | 0 | 60.948 | -107.736 | 0.818 |
| 49 | 9 | 71.136 | -86.837 | 0.578 |
| 49 | 16 | 81.362 | -61.249 | 6.613 |
| 49 | 11 | 83.987 | -56.155 | 0.761 |
| 40 | 12 | 85.880 | -64.039 | 0.695 |
| 49 | 13 | 67.775 | -53.086 | 0.748 |
| 50 | 1 | 86.855 | -47.955 | 0.938 |
| 50 | 2 | 83.367 | -46.186 | 0.938 |
| 68 | 3 | 86.826 | -43.868 | 0.927 |
| 56 | 4 | 89.398 | -44.637 | 0.917 |
| 56 | 5 | 86.356 | -44.466 | 0.922 |

## FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE

| LEVEL | WINDOW | PHI | theta | F |
| :---: | :---: | :---: | :---: | :---: |
| 50 | 6 | 79.202 | -44.233 | 0.876 |
| 50 | 7 | 76.816 | -44.334 | 0.794 |
| 56 | 8 | 73.350 | -42.681 | 0.753 |
| 60 | 9 | 86.144 | -26.013 | 0.710 |
| 60 | 18 | 84.436 | 166.438 | 0.838 |
| 56 | 11 | 83.349 | 187.183 | 0.874 |
| 58 | 12 | 83.280 | 106.991 | 0.877 |
| 60 | 13 | 83.639 | 187.285 | 0.969 |
| 51 | 1 | 89.889 | 123.366 | 0.738 |
| 51 | 2 | 83.177 | -69.180 | 0.831 |
| 51 | 3 | 78.374 | -68.287 | 0.878 |
| 51 | 4 | 75.398 | -59.568 | 0.891 |
| 51 | 5 | 74.625 | -57.816 | 0.909 |
| 51 | 6 | 74.526 | -57.656 | 0.932 |
| 51 | 7 | 72.147 | -59.946 | 0.944 |
| 51 | 8 | 67.501 | -61.498 | 0. 947 |
| 51 | 9 | 63.668 | -80.218 | 0.939 |
| 51 | 16 | 61.315 | -67.518 | 0.913 |
| 61 | 11 | 66.612 | -65.238 | -. 886 |
| 52 | 1 | 83.043 | 141.851 | 0. 938 |
| 52 | 2 | 88.244 | 137.488 | - . 947 |
| 52 | 3 | 86.028 | -43.528 | 0.946 |
| 52 | 4 | 83.555 | -41.852 | 0.943 |
| 52 | 5 | 82.856 | -39.757 | 0.946 |
| 52 | 6 | 81.471 | -41.665 | 0.899 |
| 62 | 7 | 78.184 | -46.982 | 0.876 |
| 52 | 8 | 75.825 | -48.260 | 0.893 |
| 52 | 9 | 75.873 | -45.278 | 0.886 |
| 52 | 10 | 77.752 | -40.891 | 0.836 |
| 52 | 11 | 80.036 | -38.112 | 0.756 |
| 52 | 12 | 02.222 | -39.844 | 0.760 |

## FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE

LEVEL 1


LEVEL 2


LEVEL 3


LEVEL 5


LEVEL 4


LEVEL 6


FAR OFFSET SH $_{t}$ SOURCE


LEVEL 8


LEVEL 9


LEVEL 12


FAR OFFSET SH $_{t}$ SOURCE


LEVEL 16


LEVEL 17


LEVEL 15


LEVEL 18


FAR OFFSET SH $_{\mathrm{t}}$ SOURCE


LEVEL 21


LEVEL 24


FAR OFFSET SH ${ }_{t}$ SOURCE

LEVEL 25


LEVEL 28



LEVEL 29


LEVEL 27


LEVEL 30


FAR OFFSET SH $_{\mathrm{t}}$ SOURCE


LEVEL 34


LEVEL 32


LEVEL 35


LEVEL 33


LEVEL 36


FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE


LEVEL 41


LEVEL 39


LEVEL 42


FAR OFFSET SH $_{\mathrm{t}}$ SOURCE

## LEVEL 43



LEVEL 46


LEVEL 44


LEVEL 47


LEVEL 45


LEVEL 48


FAR OFFSET $\mathrm{SH}_{\mathrm{t}}$ SOURCE


LEVEL 52


LEVEL 50


LEVEL 51


## FAR OFFSET SV SOURCE

| LEVEL | WINDOW | PHI | THETA | F |
| :---: | :---: | :---: | :---: | :---: |
| 1 | 1 | 23.441 | -86.311 | 0.737 |
| 1 | 2 | 14.471 | -67.225 | 0.739 |
| 1 | 3 | 12.489 | -32.824 | 0.726 |
| 1 | 4 | 12.688 | -28.491 | 0.716 |
|  | 5 | 12.358 | -36.088 | 0.768 |
| 1 | 6 | 11.889 | -30.328 | 0.858 |
| 1 | 7 | 13.613 | -10.588 | 0.913 |
| 1 | 8 | 14.887 | -11.891 | 0.927 |
| 1 | 9 | 15.873 | -18.681 | 0.923 |
| 1 | 10 | 18.307 | -32.203 | 6.931 |
| 1 | 11 | 16.609 | -46.868 | 8. 988 |
| 1 | 12 | 16.569 | -39.132 | 0.986 |
| 1 | 13 | 16.411 | -36.952 | 0.983 |
| 1 | 14 | 15.753 | -46.881 | 0.969 |
| 1 | 15 | 16.184 | -69.746 | 0.962 |
| 1 | 16 | 18.580 | -85.462 | 0.969 |
| 1 | 17 | 18.885 | -83.869 | 0.941 |
| 1 | 18 | 18.331 | -76.493 | 0.694 |
| 2 | 1 | 5.714 | 62.379 | 0.602 |
| 2 | 2 | 6.948 | 63.843 | 0.766 |
| 2 | 3 | 5.938 | 88.886 | 0.862 |
| 2 | 4 | 4.364 | 56.699 | 0.949 |
| 2 | 5 | 3.500 | 56.813 | 0.963 |
| 2 | 6 | 4.102 | 58.390 | 0.983 |
| 2 | 7 | 3.974 | 49.800 | 0.990 |
| 2 | 8 | 2.172 | -2.472 | 0.977 |
| 2 | 9 | 4.521 | -74.968 | 0.971 |
| 2 | 10 | 0.178 | -88.962 | 0.989 |
| 2 | 11 | 10.750 | -92.507 | 0.991 |
| 2 | 12 | 11.642 | -92.119 | 6.998 |
| 2 | 13 | 11.536 | -92.564 | 0.985 |
| 3 | 1 | 4.893 | 19.307 | 0.610 |
| 3 | 2 | 2.883 | 19.379 | 0.704 |
| 3 | 3 | 1.353 | 36.148 | 0.815 |
| 3 | 4 | 1.203 | 68.458 | 0.887 |
| 3 | 6 | 1.908 | 75.262 | 0.936 |
| 3 | 6 | 2.636 | 81.641 | 0.961 |
| 3 | 7 | 2.301 | 99.433 | 0.964 |
| 3 | 8 | 1.291 | 161.261 | 0.957 |
| 3 | $\theta$ | 2.863 | -116.646 | 0.953 |
| 3 | 18 | 0. 238 | -94.487 | 0.056 |
| 4 | 1 | 5.856 | 164.691 | 0.056 |
| 4 | 2 | 3.120 | 117.797 | 0.921 |
| 4 | 3 | 2.386 | 126.860 | 0.983 |
| 4 | 4 | 2.306 | 122.108 | 6.982 |
| 4 | 5 | 1.822 | 163.989 | 8. 986 |
| 4 | 6 | 2.194 | -138.011 | 0.977 |
| 4 | 7 | 3.851 | -103.028 | 0.979 |
| 4 | 8 | 5.842 | -87.611 | 0.981 |
| 4 | 9 | 7.968 | -89.896 | 0. 984 |
| 5 | 1 | 13.248 | -108.096 | 0.760 |
| 6 | 2 | 19.280 | -186.547 | 0.851 |
| 6 | 3 | 23.285 | -144.161 | 0.824 |
| 6 | 4 | 28.843 | -136.565 | 0.948 |
| 6 | 5 | 29.806 | -138.811 | 0.967 |
| 6 | 6 | 29.668 | -128.626 | 0.964 |

## FAR OFFSET SV SOURCE

| LEVEL | WINDOW | PHI | THETA | $F$ |
| :---: | :---: | :---: | :---: | :---: |
| 6 | 7 | 30.382 | -125.882 | 0. 966 |
| 6 | 8 | 31.914 | -122.787 | 0.964 |
| 5 | 9 | 33.534 | -119.861 | 0.961 |
| 6 | 16 | 32.842 | -116.946 | 0.948 |
| 6 | 11 | 29.736 | -116.178 | 0.928 |
| 5 | 12 | 25.772 | -113.802 | 8.916 |
| 6 | 13 | 21.478 | -106.306 | 0.898 |
| 6 | 1 | 38.382 | -123.693 | 0.868 |
| - | 2 | 38.388 | -124.985 | 0.924 |
| 8 | 3 | 37.365 | -127.634 | 0.941 |
| 0 | 4 | 37.276 | -127.719 | 0.923 |
| 6 | 5 | 38.248 | -124.142 | 0.917 |
| 6 | 6 | 37.384 | -123.842 | 0.938 |
| 6 | 7 | 36.381 | -125.808 | 0.952 |
| 6 | 8 | 34.228 | -128.878 | 0.958 |
| 6 | 9 | 34.694 | -129.846 | 0.956 |
| 6 | 10 | 34.678 | -128.462 | 0.965 |
| 6 | 11 | 34.674 | -127:866 | 0.969 |
| 6 | 12 | 33.481 | -129.476 | 0.986 |
| 6 | 13 | 32.196 | -132.604 | 0.981 |
| 7 | 1 | 52.372 | -98.852 | 0.901 |
| 7 | 2 | 46.281 | -97.552 | 0.873 |
| 7 | 3 | 41.446 | -96.931 | 0.831 |
| 7 | 4 | 39.459 | -96.821 | 0.801 |
| 7 | 5 | 39.217 | -96.834 | 0.864 |
| 7 | 6 | 37.143 | -98.928 | 0.858 |
| 7 | 7 | 33.465 | -182.624 | 0.989 |
| 7 | 8 | 30.806 | -104.159 | 0.938 |
| 7 | 9 | 29.838 | -103.698 | 0.943 |
| 7 | 16 | 39.510 | -102.995 | 0.943 |
| 7 | 11 | 32.296 | -103.401 | 0.954 |
| 7 | 12 | 32.364 | -103.781 | 0.969 |
| 7 | 13 | 39.298 | -103.269 | 0.973 |
| 7 | 14 | 27.238 | -101.672 | 0.971 |
| 8 | 1 | 86.114 | -89.985 | 0. 785 |
| 8 | 2 | 87.045 | -99.199 | 0.769 |
| 8 | 3 | 86.333 | -89.976 | 0.770 |
| 8 | 4 | 81.112 | -87.973 | 0.786 |
| 8 | 5 | 86.446 | -82.886 | 0.729 |
| 8 | 6 | 47.279 | -83.618 | 0.727 |
| 8 | 7 | 54.804 | -98.113 | 0.775 |
| 8 | 8 | 68.924 | -96.888 | 0.923 |
| 8 | 9 | 02.436 | -93.291 | 0.947 |
| 8 | 10 | 02.844 | -92.371 | 0. 906 |
| 8 | 11 | 08.718 | -94.291 | 0.862 |
| 8 | 12 | 67.481 | -96.622 | 0.843 |
| 8 | 13 | 55.331 | -94.461 | 0.822 |
| $\bigcirc$ | 1 | 56.196 | 89.871 | 0.717 |
| 9 | 2 | 49.293 | 84.254 | 0.732 |
| 9 | 3 | 47.244 | 81.461 | 0.728 |
| 9 | 4 | 42.969 | 76.872 | 0.718 |
| 9 | 5 | 32.869 | 65.318 | 0.815 |
| 9 | 6 | 27.738 | 56.682 | 0.962 |
| 9 | 7 | 26.865 | 62.600 | 0.963 |
| 9 | 8 | 22.072 | 39.433 | 0.888 |

## FAR OFFSET SV SOURCE

| LEVEL | WINDOW | PHI | THETA | F |
| :---: | :---: | :---: | :---: | :---: |
| 10 | 1 | 62.481 | 94.794 | 0.833 |
| 10 | 2 | 68.747 | 85.171 | 0.780 |
| 10 | 3 | 43.171 | 71.728 | 0.789 |
| 10 | 4 | 39.996 | 58.924 | 0.854 |
| 10 | 5 | 36.888 | 49.688 | 0.938 |
| 10 | 6 | 33.369 | 43.811 | 0.963 |
| 10 | 7 | 32.825 | 37.816 | 8.967 |
| 10 | 8 | 33.181 | 30.585 | 0.971 |
| 11 | 1 | 54.736 | 135.000 | 0.008 |
| 11 | 2 | 54.736 | 135.006 | 0.008 |
| 11 | 3 | 54.736 | 136.008 | 0.600 |
| 11 | 4 | 54.736 | 136.880 | 0.600 |
| 11 | 5 | 54.736 | 136.060 | 0.008 |
| 11 | 6 | 54.736 | 135.606 | 0.060 |
| 11 | 7 | 54.736 | 136.008 | 0.060 |
| 11 | 8 | 64.736 | 136.600 | 0.080 |
| 12 | 1 | 54.736 | 136.600 | 0.008 |
| 12 | 2 | 64.736 | 136.008 | 0.008 |
| 12 | 3 | 54.736 | 136.060 | 0.068 |
| 12 | 4 | 54.736 | 136.600 | 0.008 |
| 12 | 5 | 64.736 | 136.006 | 0.096 |
| 12 | 6 | 64.736 | 136.008 | 0.080 |
| 12 | 7 | 54.736 | 136.080 | 0.000 |
| 12 | 8 | 54.736 | 135.006 | 0.080 |
| 12 | 9 | 54.736 | 135.008 | 0.000 |
| 12 | 10 | 54.736 | 135.008 | 0.800 |
| 12 | 11 | . 54.736 | 136.000 | - . 800 |
| 12 | 12 | 54.736 | 135.008 | 0.000 |
| 13 | 1 | 54.736 | 135.696 | 0.080 |
| 13 | 2 | 54.736 | 136.606 | 0.006 |
| 13 | 3 | 54.736 | 135.000 | 0.808 |
| 13 | 4 | 54.736 | 136.060 | 0.000 |
| 13 | 5 | 64.736 | 135.000 | 0.606 |
| 13 | 6 | 64.736 | 136.008 | 0.806 |
| 13 | 7 | 64.736 | 136.069 | 0.806 |
| 13 | 8 | 64.736 | 135.609 | 0.008 |
| 13 | 9 | 54.736 | 136.880 | 0.800 |
| 13 | 10 | 84.736 | 135.000 | 0.000 |
| 14 | 1 | 8.490 | -137.146 | 0.946 |
| 14 | 2 | 8.548 | -146.668 | 0.983 |
| 14 | 3 | 9.421 | -154.889 | 0.968 |
| 14 | 4 | 9.982 | -157.311 | 0.962 |
| 14 | 5 | 9.153 | -151.739 | 0.956 |
| 14 | 6 | 8.339 | -143.069 | 0.960 |
| 14 | 7 | 9.896 | -142.329 | 6.958 |
| 14 | 8 | 11.388 | -146.502 | 0.955 |
| 14 | 9 | 14.439 | -148.386 | 0.942 |
| 14 | 18 | 16.678 | -146.433 | 0.924 |
| 15 | 1 | 61.489 | -106. 468 | 0.879 |
| 15 | 2 | 67.861 | -180.436 | 0.913 |
| 15 | 3 | 60.537 | -99.596 | 0.931 |
| 16 | 4 | 69.725 | -106.961 | 0.848 |
| 16 | 5 | 86.941 | -104.168 | 0.710 |
| 15 | 6 | 81.312 | 72.344 | 0.713 |
| 16 | 7 | 79.472 | 79.364 | 0.750 |

## FAR OFFSET SV SOURCE

| LEvEl | WINDOW | PHI | theta | F |
| :---: | :---: | :---: | :---: | :---: |
| 15 | 8 | 79.459 | 88.787 | 0.784 |
| 15 | 9 | 74.452 | 68.984 | 0.848 |
| 15 | 10 | 67.732 | 89.333 | 0.882 |
| 16 | 1 | 16.465 | -54.036 | 0.939 |
| 16 | 2 | 15.444 | -89.798 | 0.937 |
| 18 | 3 | 15.618 | -79.971 | 0.928 |
| 18 | 4 | 15.828 | -81.825 | 0.894 |
| 16 | 5 | 16.088 | -77.829 | 0.876 |
| 16 | 6 | 16.987 | -76.372 | 0.874 |
| 18 | 7 | 18.807 | -81.598 | 0.872 |
| 16 | $\theta$ | 29.555 | -136.006 | 0.851 |
| 16 | 9 | 25.514 | -181.835 | 0.799 |
| 16 | 16 | 27.535 | -116.081 | 0.752 |
| 16 | 11 | 27.206 | -114.964 | 0.760 |
| 10 | 12 | 26.896 | -120.059 | 0.889 |
| 17 | 1 | 17.478 | -96.818 | 0.859 |
| 17 | 2 | 19.316 | -103.811 | 0.882 |
| 17 | 3 | 29.986 | -108.136 | 0.878 |
| 17 | 4 | 21.598 | -108.411 | 0.855 |
| 17 | 5 | 21.408 | -186.588 | 0.837 |
| 17 | 6 | 21.957 | -107.285 | 0.837 |
| 17 | 7 | 24.830 | -112.481 | 0.833 |
| 17 | 8 | 30.297 | -119.885 | 0.868 |
| 17 | 9 | 36.417 | -125.963 | 0.775 |
| 17 | 10 | 37.716 | -127.382 | 0.768 |
| 17 | 11 | 35.633 | -126.171 | 0.810 |
| 17 | 12 | 33.366 | -127.140 | ©. 884 |
| 18 | 1 | 15.266 | -166.156 | 0.924 |
| 18 | 2 | 10.626 | -168.497 | 0.937 |
| 18 | 3 | 16.527 | -108.136 | 0.946 |
| 18 | 4 | 16.388 | -106.199 | 0.937 |
| 18 | 5 | 26.711 | -182.360 | 0.929 |
| 18 | 6 | 16.163 | -181.782 | 0.928 |
| 18 | 7 | 15.329 | -182.480 | 0.905 |
| 18 | 8 | 16.971 | -104.169 | 0.872 |
| 18 | 9 | 20.194 | -167.669 | 0.834 |
| 18 | 10 | 22.644 | -110.987 | 0.832 |
| 18 | 11 | 24.069 | -114.520 | 0.858 |
| 19 | 1 | 19.085 | -04.898 | 6. 982 |
| 19 | 2 | 20.380 | -64.566 | 0.976 |
| 19 | 3 | 20.486 | -62.924 | 0.983 |
| 19 | 4 | 20.950 | -61.978 | 0.987 |
| 19 | 5 | 21.289 | -62.978 | 0.992 |
| 18 | 6 | 26.583 | -63.761 | 0.982 |
| 19 | 7 | 18.526 | -62.312 | 0.979 |
| 19 | 8 | 15.528 | -56.332 | 6.963 |
| 19 | $\bigcirc$ | 12.670 | -46.297 | 0.063 |
| 19 | 18 | 12.610 | -44.619 | 0.939 |
| 28 | 1 | 14.894 | -182.227 | 0.958 |
| 20 | 2 | 14.795 | -101.946 | 0.958 |
| 28 | 3 | 14.691 | -99.678 | 0.962 |
| 20 | 4 | 15.162 | -181. 189 | 0.974 |
| 20 | 5 | 14.787 | -104.783 | 0.996 |
| 20 | $\bigcirc$ | 12.872 | -104. 140 | 0.987 |
| 20 | 7 | 10.174 | -95.640 | 0.975 |

## FAR OFFSET SV SOURCE

| LEVEL | WINDOW | PHI | THETA | F |
| :---: | :---: | :---: | :---: | :---: |
| 20 | 8 | 8.149 | -77.193 | 0.971 |
| 28 | 9 | 7.896 | -61.647 | 6.973 |
| 21 | 1 | 11.728 | -66.421 | 8.968 |
| 21 | 2 | 11.767 | -86.828 | 0.978 |
| 21 | 3 | 11.352 | -70.649 | 0.978 |
| 21 | 4 | 10.121 | -79.206 | 0.981 |
| 21 | 5 | 8.713 | -84.836 | 0.993 |
| 21 | 6 | 7.864 | -79.498 | 6.992 |
| 21 | 7 | 7.465 | -67.677 | 0.986 |
| 21 | 8 | 7.844 | -65.264 | 0.982 |
| 21 | 9 | 7.713 | -69.489 | 0.981 |
| 21 | 10 | 7.592 | -46. 298 | 0.979 |
| 22 | 1 | 15.068 | -60.674 | 0.986 |
| 22 | 2 | 13.271 | -68.565 | 0.983 |
| 22 | 3 | 13.356 | -59.336 | 0.986 |
| 22 | 4 | 12.828 | -64.736 | 0.971 |
| 22 | 5 | 19.555 | -65.944 | 0.983 |
| 22 | 6 | 8.171 | -68.027 | 0.988 |
| 22 | 7 | 0.889 | -44.153 | 0.984 |
| 22 | 8 | 6.774 | -34.712 | 0.981 |
| 22 | 9 | 7.275 | -36.174 | 0.979 |
| 22 | 10 | 7.828 | -34.487 | 0.978 |
| 23 | 1 | 7.078 | -146.825 | 0.928 |
| 23 | 2 | 6.979 | -127.224 | 0.947 |
| 23 | 3 | 6.149 | -112.971 | 0.961 |
| 23 | 4 | 6.817 | -113.297 | 0.951 |
| 23 | 5 | 9.589 | -129.376 | 0.946 |
| 23 | 8 | 10.204 | -136.143 | 0.953 |
| 23 | 7 | 8.161 | -137.857 | 0.967 |
| 23 | 8 | 5.388 | -140.733 | 0.946 |
| 23 | 9 | 3.788 | -152.627 | 0.938 |
| 23 | 10 | 4.349 | -103.822 | 0.924 |
| 23 | 11 | 4.649 | -166.829 | 0.937 |
| 24 | 1 | 15.948 | -178.928 | 0.954 |
| 24 | 2 | 14.019 | -162.832 | 0.959 |
| 24 | 3 | 12.120 | -158.005 | 0.968 |
| 24 | 4 | 12.283 | -154.661 | 0.968 |
| 24 | 5 | 15.651 | -163.692 | 0.939 |
| 24 | 6 | 16.885 | -154.248 | 0.952 |
| 24 | 7 | 14.785 | -166.461 | 0.956 |
| 24 | 8 | 11.882 | -168.268 | 0.944 |
| 24 | 9 | 9.933 | -163.914 | 0.922 |
| 24 | 10 | 10.267 | -168.419 | 6.988 |
| 24. | 11 | 10.847 | -169.272 | 9.918 |
| 25 | 1 | 28.344 | -172.092 | 1.984 |
| 25 | 2 | 27.422 | -167.368 | 0. 968 |
| 25 | 3 | 28.441 | -165.568 | 6.989 |
| 25 | 4 | 27.418 | -164.269 | 0.973 |
| 25 | 5 | 29.620 | -162.175 | 6.974 |
| 25 | 6 | 29.264 | -160.669 | 0.979 |
| 25 | 7 | 26.747 | -160.022 | ¢. 977 |
| 26 | 8 | 23.689 | -162.522 | 0.968 |
| 25 | 9 | 21.182 | -166.096 | 0.955 |
| 26 | 10 | 29.824 | -168.694 | 8.944 |

## FAR OFFSET SV SOURCE



FAR OFFSET SV SOURCE

| LEVEL | WINDOW | PHI | THETA | $F$ |
| :---: | :---: | :---: | :---: | :---: |
| 31 | 3 | 24.128 | -124.321 | 0.964 |
| 31 | 4 | 25.310 | -126.409 | 0.980 |
| 31 | 5 | 24.488 | -127.610 | 0.962 |
| 31 | 0 | 22.814 | -128.719 | 0.951 |
| 31 | 7 | 28.924 | -135.586 | 0.933 |
| 31 | 8 | 32.541 | -140.716 | 0.94.7 |
| 31 | 9 | 34.916 | -144.148 | 0.978 |
| 31 | 16 | 35.289 | -147.332 | 0.980 |
| 31 | 11 | 34.182 | -158.282 | 0.98\% |
| 31 | 12 | 32.743 | -151.532 | d.95\% |
| 32 | 1 | 36.052 | -139.214 | 0.910 |
| 32 | 2 | 31.316 | -137.226 | 0.916 |
| 32 | 3 | 28.585 | -134.142 | 0.917 |
| 32 | 4 | 26.382 | -128.858 | 0.934 |
| 32 | 5 | 23.596 | -121.686 | g. 953 |
| 32 | - | 21.347 | -117.272 | 0.959 |
| 32 | 7 | 23.492 | -122.962 | 0.932 |
| 32 | 8 | 28.149 | -128.368 | 0.932 |
| 32 | 9 | 36. 608 | -129.579 | 0.962 |
| 32 | 10 | 29.588 | -139.681 | 0.968 |
| 32 | 11 | 28.462 | -133.670 | 0.946 |
| 32 | 12 | 27.673 | -137.603 | 0.937 |
| 32 | 13 | 27.242 | -139.752 | 0.937 |
| 33 | 1 | 39.452 | -153.801 | 0.961 |
| 33 | 2 | 36.698 | -151.106 | 0.932 |
| 33 | 3 | 33.399 | -148.399 | 0.917 |
| 33 | 4 | 31.698 | -144.407 | 0.915 |
| 33 | 5 | 29.561 | -138.400 | 0.919 |
| 33 | 6 | 25.860 | -132.266 | 0.936 |
| 33 | 7 | 24.203 | -135.971 | 0.942 |
| 33 | 8 | 28.627 | -146.160 | 0.928 |
| 33 | 8 | 33.311 | -150.691 | 0.948 |
| 33 | 10 | 36.361 | -152.253 | 0.976 |
| 33 | 11 | 36.562 | -163.161 | 0.984 |
| 33 | 12 | 36.608 | -153.687 | 0.974 |
| 33 | 13 | 34.627 | -163.796 | 0.966 |
| 34 | 1 | 31.854 | -130.373 | 0.986 |
| 34 | 2 | 29.416 | -130.067 | 0.946 |
| 34 | 3 | 26.764 | -129.069 | 0.983 |
| 34 | 4 | 24.637 | -128.092 | 0.975 |
| 34 | 5 | 22.893 | -126.296 | 0.976 |
| 34 | 6 | 28.989 | -123.334 | 0.972 |
| 34 | 7 | 19.557 | -126.183 | 0.973 |
| 34 | 8 | 28.794 | -137.782 | 0.973 |
| 34 | 8 | 25.886 | -150.845 | 0.971 |
| 34 | 10 | 29.233 | -157.472 | 6.979 |
| 34 | 11 | 31.646 | -160.579 | 0.979 |
| 34 | 12 | 32.119 | -161.798 | 0.978 |
| 34 | 13 | 32.017 | -162.396 | -. 981 |
| 36 | 1 | 32.896 | -147.653 | 0.978 |
| 36 | 2 | 31.872 | -148.993 | 0. 983 |
| 36 | 3 | 29.992 | -149.966 | 0.991 |
| 36 | 4 | 28.971 | -149.548 | 0.991 |
| 36 | 5 | 28.282 | -148.834 | 0.988 |
| 35 | 0 | 28.307 | -156.598 | 0.983 |
| 36 | 7 | 36.657 | -157.640 | 0.967 |

## FAR OFFSET SV SOURCE

| LEVEL | WINDOW | PHI | THETA | F |
| :---: | :---: | :---: | :---: | :---: |
| 35 | 8 | 32.886 | -185.389 | 0.965 |
| 35 | 9 | 34.687 | -189.140 | 0.972 |
| 35 | 10 | 34.553 | -189.138 | 0.973 |
| 35 | 11 | 33.590 | -167.525 | 0.987 |
| 36 | 12 | 32.486 | -186.396 | 0.980 |
| 36 | 13 | 31.142 | -186.684 | 0.958 |
| 36 | 1 | 37.959 | -148.415 | 0.968 |
| 36 | 2 | 36.778 | -148.314 | 8.986 |
| 36 | 3 | 35.641 | -147.899 | 8.978 |
| 36 | 4 | 34.072 | -146.788 | 0.988 |
| 36 | 5 | 32.486 | -145.396 | 0.992 |
| 36 | 8 | 32.143 | -145.467 | 0.989 |
| 36 | 7 | 33.616 | -149.233 | 0.984 |
| 36 | 8 | 34.958 | -155.069 | 0.985 |
| 38 | 9 | 36.336 | -180.069 | 0.985 |
| 38 | 10 | 35.221 | -162.984 | 0.983 |
| 36 | 11 | 35.387 | -183.845 | 0.984 |
| 36 | 12 | 35.865 | -163.911 | 0.987 |
| 36 | 13 | 35.967 | -164.461 | 0.988 |
| 37 | 1 | 45.388 | -157.794 | 0. 980 |
| 37 | 2 | 44.436 | -156.556 | - . 988 |
| 37 | 3 | 43.686 | -156.146 | 0.974 |
| 37 | 4 | 41.823 | -153.859 | 0.979 |
| 37 | 5 | 39.481 | -152.633 | 0.974 |
| 37 | 6 | 39.675 | -150.836 | 0.958 |
| 37 | 7 | 41.449 | -151.558 | 0.942 |
| 37 | 8 | 42.446 | -154.649 | ๑. 938 |
| 37 | 9 | 41.461 | -158.150 | - 0.940 |
| 37 | 10 | 39.874 | -160.463 | 0.939 |
| 37 | 11 | 39.024 | -160.689 | 0.944 |
| 37 | 12 | 38.498 | -159.984 | 0.956 |
| 37 | 13 | 37.289 | -169.378 | 0.969 |
| 38 | 1 | 52.324 | -158.965 | 8.975 |
| 38 | 2 | 51.903 | -158.684 | 0.981 |
| 38 | 3 | 51.340 | -158.253 | 0. 987 |
| 38 | 4 | 60.862 | -157.274 | 0.992 |
| 38 | 5 | 49.924 | -155.855 | 0.993 |
| 38 | 6 | 49.181 | -154.952 | 0.991 |
| 38 | 7 | 48.242 | -156.002 | 0.984 |
| 38 | 8 | 46.767 | -161.844 | 0.979 |
| 38 | 9 | 44.989 | -186.355 | 6.984 |
| 38 | 18 | 44.020 | -167.347 | 0.987 |
| 38 | 11 | 44.137 | -166.727 | 0.998 |
| 38 | 12 | 44.862 | -166.035 | 0.992 |
| 38 | 13 | 46.411 | -166.464 | 0.891 |
| 39 | 1 | 44.495 | -157.688 | 0.993 |
| 39 | 2 | 45.194 | -157.097 | 8.994 |
| 39 | 3 | 45.883 | -156.639 | 0.995 |
| 39 | 4 | 46.345 | -156.088 | 0.998 |
| 39 | 8 | 46.848 | -156.876 | 0.996 |
| 39 | 6 | 46.948 | -156.767 | 0.991 |
| 39 | 7 | 45.989 | -159.817 | 6.981 |
| 39 | 8 | 44.499 | -164.016 | 0.980 |
| 39 | 9 | 43.944 | -168.329 | 0.987 |
| 39 | 10 | 44.606 | -160.296 | 0. 988 |
| 39 | 11 | 48.416 | -166.229 | ¢. 988 |

## FAR OFFSET SV SOURCE

| LEVEL | WINDOW | PHI | THETA | $F$ |
| :---: | :---: | :---: | :---: | :---: |
| 39 | 12 | 46.087 | -104.441 | 0.981 |
| 40 | 1 | 48.874 | -188.087 | 6.993 |
| 40 | 2 | 49.258 | -167.586 | 0.996 |
| 40 | 3 | 49.380 | -188.881 | 0.997 |
| 46 | 4 | 49.538 | -168.182 | 0.998 |
| 40 | 5 | 49.953 | -165.833 | \%.998 |
| 40 | 6 | 68.348 | -165.991 | 0. 995 |
| 46 | 7 | 49.694 | -188.431 | 0.987 |
| 40 | 8 | 47.880 | -172.764 | 0.983 |
| 46 | 9 | 46.875 | -176.918 | 8.989 |
| 40 | 18 | 47.171 | -178.822 | 0.994 |
| 46 | 11 | 48.278 | -176.198 | 0.982 |
| 48 | 12 | 49.638 | -176.686 | 6. 989 |
| 46 | 13 | 49.327 | -174.824 | 8.984 |
| 41 | 1 | 48.907 | -100.317 | ¢. 982 |
| 41 | 2 | 49.723 | -159.419 | 6. 988 |
| 41 | 3 | 49.728 | -158.137 | 0.998 |
| 41 | 4 | 48.881 | -156.443 | 8.994 |
| 41 | 6 | 48.054 | -156.020 | 0.993 |
| 41 | 6 | 49.868 | -156.082 | 0.992 |
| 41 | 7 | 51.276 | -167.367 | 8. 989 |
| 41 | 8 | 51.587 | -101.587 | 6.984 |
| 41 | 9 | 51.068 | -168.211 | 6.986 |
| 41 | 10 | 51.824 | -169.174 | 0.981 |
| 41 | 11 | 61.732 | -169.447 | 8.988 |
| 41 | 12 | 62.449 | -167.972 | 8. 989 |
| 42 | 1 | 59.891 | - 171.882 | 6.986 |
| 42 | 2 | 69.449 | -176.941 | 6.989 |
| 42 | 3 | 58.491 | -169.387 | 6.986 |
| 42 | 4 | 57.738 | -188.486 | 0.972 |
| 42 | 6 | 58.811 | -189.686 | 0.948 |
| 42 | 8 | 61.995 | -172.464 | 0.938 |
| 42 | 7 | 63.981 | -173.824 | -. 948 |
| 42 | 8 | 83.782 | -174.877 | 0.982 |
| 42 | 9 | 82.998 | -174.435 | 0.964 |
| 42 | 18 | 82.658 | -174.938 | 6.968 |
| 42 | 11 | 62.251 | -174.832 | 0.961 |
| 42 | 12 | 61.438 | -173.898 | 6.986 |
| 43 | 1 | 56.886 | -98. 980 | 0.988 |
| 43 | 2 | 56.568 | -98.808 | B. 894 |
| 43 | 3 | 55.590 | -96. 696 | 6.996 |
| 43 | 4 | 68.443 | -98. 986 | 0.994 |
| 43 | 5 | 69.226 | -98.096 | 0.988 |
| 43 | 6 | 02.298 | -98. 288 | 6.978 |
| 43 | 7 | 03.311 | -90. 368 | 6.971 |
| 43 | 8 | 02.856 | -96. 886 | 0.988 |
| 43. | 9 | 61.863 | -98. 888 | ¢. 968 |
| 43 | $1 \%$ | 61.494 | -98. 386 | 6. 961 |
| 43 | 11 | 01.918 | -80.888 | 9. 986 |
| 44 | 1 | 58.897 | -189.781 | 6. 983 |
| 44 | 2 | 56.483 | -189.781 | 6.986 |
| 44 | 3 | 64.181 | -189.781 | 6.978 |
| 44 | 4 | 54.389 | -169.781 | 6.948 |
| 44 | 5 | 65.982 | -189.781 | 6. 922 |
| 44 | 6 | 50.824 | -100.781 | 6.911 |

## FAR OFFSET SV SOURCE

| LEVEL | WINOOW | PHI | THETA | F |
| :---: | :---: | :---: | :---: | :---: |
| 44 | 7 | 53.920 | -169.781 | 0.892 |
| 44 | 8 | 48.979 | -169.781 | 0.891 |
| 44 | 9 | 48.780 | -169.781 | 0.900 |
| 44 | 10 | 48.586 | -169.781 | 0.916 |
| 44 | 11 | 46.712 | -169.781 | 0.948 |
| 44 | 12 | 44.383 | -169.781 | 0.977 |
| 44 | 13 | 43.201 | -189.781 | 0.986 |
| 45 | 1 | 65.735 | -167.976 | 0.987 |
| 46 | 2 | 56.848 | -167.132 | 0.992 |
| 45 | 3 | 54.286 | -186.681 | 0.993 |
| 46 | 4 | 54.206 | -186.116 | 0.984 |
| 45 | 5 | 56.078 | -165.846 | 0.975 |
| 45 | 6 | 55.917 | -164.546 | 0.970 |
| 46 | 7 | 54.896 | -167.634 | 0.985 |
| 46 | 8 | 52.249 | -176.681 | 0.960 |
| 45 | 9 | 50.884 | 177.346 | 0.961 |
| 46 | 18 | 50.674 | 175.087 | 0.983 |
| 46 | 11 | 60.926 | 176.736 | 0.869 |
| 46 | 12 | 56.819 | 176.366 | 0.973 |
| 46 | 13 | 49.989 | 175.788 | 0.970 |
| 46 | 1 | 47.983 | -169.312 | 0.932 |
| 46 | 2 | 46.281 | -113.235 | 0.938 |
| 46 | 3 | 44.131 | -119.494 | 0.918 |
| 46 | 4 | 41.360 | -122.212 | 0.832 |
| 46 | 5 | 41.964 | -118.482 | 0.712 |
| 46 | 6 | 48.677 | -116.258 | 0.678 |
| 46 | 7 | 49.823 | -122.716 | 0.721 |
| 46 | 8 | 44.866 | -132.644 | 0.720 |
| 46 | 9 | 38.386 | -144.712 | 0.653 |
| 46 | 18 | 33.263 | -157.780 | 0.652 |
| 46 | 11 | 32.491 | -162.372 | 0. 489 |
| 46 | 12 | 36.267 | -158.884 | 0.562 |
| 47 | 1 | 55.458 | -177.541 | 0. 988 |
| 47 | 2 | 54.459 | -176.785 | 0.987 |
| 47 | 3 | 63.249 | -176.388 | 0.986 |
| 47 | 4 | 52.422 | -176.744 | 0.978 |
| 47 | 5 | 62.773 | -174.258 | 0.964 |
| 47 | 6 | 54.604 | -173.218 | 0.956 |
| 47 | 7 | 64.167 | -174.867 | 0. 968 |
| 47 | - | 53.112 | -178.649 | 0.961 |
| 47 | $\bigcirc$ | 52.369 | 178.212 | 0.982 |
| 47 | 10 | 62.341 | 177.602 | 0.964 |
| 47 | 11 | 52.438 | 177.616 | 0.971 |
| 48 | 1 | 68.737 | -169.386 | 0.971 |
| 48 | 2 | 67.468 | -168.836 | 6.967 |
| 48 | 3 | 55.788 | -167.887 | 0.963 |
| 48 | 4 | 53.822 | -167.814 | 0.958 |
| 48 | 5 | 52.258 | -166.152 | 0.949 |
| 48 | 6 | 52.439 | -102.388 | 0.946 |
| 48 | 7 | 53.297 | -162.066 | 0.947 |
| 48 | 8 | 53.680 | -164.476 | 0.958 |
| 48 | 9 | 52.571 | -167.296 | 0.983 |
| 40 | 10 | 62.503 | -169.252 | 0.963 |
| 46 | 11 | 52.711 | -189.977 | 8.966 |
| 40 | 1 | 55.613 | 165.183 | 0.997 |

## FAR OFFSET SV SOURCE

| LEVEL | WINOOW | PHI | THETA | F |
| :---: | :---: | :---: | :---: | :---: |
| 49 | 2 | 54.806 | 165.677 | 0.995 |
| 49 | 3 | 54.333 | 185.565 | 0.991 |
| 49 | 4 | 54.549 | 165.378 | 0.980 |
| 49 | 5 | 55.128 | 166.109 | 0.982 |
| 49 | 8 | 56.323 | 187.780 | 0.981 |
| 49 | 7 | 56.037 | 188.631 | 0.981 |
| 49 | 8 | 54.991 | 167.497 | 0.982 |
| 49 | 9 | 56.536 | 165.442 | 0.984 |
| 49 | 10 | 50.111 | 164.148 | 0.987 |
| 50 | 1 | 50.896 | -188.329 | 6.998 |
| 58 | 2 | 50.474 | -167.928 | 0.999 |
| 60 | 3 | 49.921 | -187.764 | 0.998 |
| 56 | 4 | 49.585 | -187.823 | 0.997 |
| 56 | 6 | 50.171 | -167.395 | 0.994 |
| 60 | 6 | 51.853 | -186.238 | 0.996 |
| 65 | 7 | 53.868 | -165.718 | 0.992 |
| 50 | 8 | 63.172 | -166.377 | 0.993 |
| 50 | 9 | 52.934 | -107.706 | 0.992 |
| 56 | 10 | 52.925 | -168.852 | 0.988 |
| 56 | 11 | 53.048 | -169.309 | 0.985 |

FAR OFFSET SV SOURCE

LEVEL 1


LEVEL 4


LEVEL 2


LEVEL 3


LEVEL 6


FAR OFFSET SH $_{t}$ SOURCE


LEVEL 10


LEVEL 8


LEVEL 9


$$
\text { LEVEL } 12
$$



FAR OFFSET SV SOURCE


LEVEL 16


LEVEL 14


LEVEL 17


LEVEL 15


LEVEL 18


FAR OFFSET SV SOURCE


LEVEL 20


LEVEL 23


LEVEL 21


LEVEL 24


FAR OFFSET SV SOURCE


LEVEL 26


LEVEL 29


LEVEL 27


LEvel 30


## FAR OFFSET SV SOURCE



LEVEL 34


LEVEL 32


LEVEL 33


LEVEL 35


LEVEL 36


FAR OFFSET SV SOURCE


LEVEL 38


LEVEL 41


LEVEL 39


LEVEL 42


FAR OFFSET SV SOURCE

LEVEL 43


LEVEL 46



LEVEL 47


LEVEL 45


LEVEL 48


FAR OFFSET SV SOURCE


LEVEL 50


## APPENDIX 4

## FORTRAN 77 PROGRAM LISTINGS

HODOS.FOR VEL.FOR
ROTBOR.FOR

PROGRAM HODOS



GO TO 999



```
    YMAX=MAX(YMAX.ABS(TRACE2 J)))
    ZMAX=MAX(ZMAX,ABS(TRACE3(J)))
    CONTINUE
    GSCALE =MAX(XOMAX,YMAX,ZMAX)
    FF(GSCALE LT 10.0*-10)GSCALE=1.0
    WRITE(5, ')' SCALING FACTOR-',GSCALE
    DO 370 I=INPTS I NORMALIZE TO MAX =1.0
    TRACEl(J)= TRACEl(J)/GSCALE
    TRACE2(J)=TRACE2 (J)/GSCALE
    TRACE3(J)=TRACE3(J)/CSCALE
    CONTINUE
```




```
    WRITE(&")DO YOU WANT A NEW VIEW POINT FOR THE ID PLOTP (1= YES)*
    READ(S.")VUTEST
    PHI=-40
    THETA=20
    RADIUS=12
    IF(VUTEST EQ.1)THEN
        WRITE(6,*'ENTER VIEWPOINT - PHI,THETA,RADIUS'
        READ(5, ^PHD,THETA,RADIUS
    ENDPP
    ENCODE(28.389.TITLE2).LEVEL | CREATE PLOT TITLE
330 FORMATY'PARTICLE MOTION AT LEVEL''I3,'%')
    KOUNT2-KOUNT2+1
    JKOUNT2= \KOUNT2 +1
    IF( KOUNT2 GT 6)THEN
            JKOUNT2=1
            CALL ENDPLO)
        ENDF
            IP(SKOUNT2.EQ.1)THEN
        C CALL NOMINS
        CALL PTK4L
        CALL PVTE4O
        CALL TALARS
        CALL PAGE(8 5,110)
        ENDIP
            IF(JKOUNT2 EQ.1)CALL PHYSOR(0.1,70)
        F(JKOUNT2 EQ.2)CALL PHYSOR(0.1,3.5)
        JKOUNT2 EQ3)CALL PHYSOR(0.1.01
        JKOUNT2.EQ.4)CALL PHYSOR(4.1,7.0)
        JKOUNT2.EQ.5)CALL. PHYSOR(4.1,3.5
        FF(JKOUNT2 EQ.6)CALL PHYSOR(4.1.01)
    C CALL NOMIN& CALL PTK&1 PICK PLOTTING DEVICE
        CALL PAGE(8.5.8.5)
    CALL AREA2D(40.35)
    OALL VOLM3D(50.5.0.50)
    CALL MESSAC(%REF(TITLE2),27,0.5.3 25)
    CALL VUANGLPHITHETA,RADIUS)
    CALL X3NAME 'SV'.2)
    CALL YJNAME('SH'.2)
    CALL Z3NAME('RADINL',1)
```


## CALL GRAF3D(-1,1,1,-1,1,1,-1,1,1) <br> I DRAW AXIS

CALL BOX3D
IF(PLTST NE 2 0)THEN
CALL MARKER(15)
CALL CURV3D TRACE $2, T R A C E 3, T R A C E 1,3,1)$ ! MARK FIRST $\operatorname{I}$ POINTS
CALL MARKER(16)
CALL CURV3D(TRACE2,TRACE3,TRACEI,NPTS.1) ! PLOT S-D CURVE ENDIP

LF(PLTST NE. 3 O)THEN

```
    CALL MARKER(13)
    CALL CURVID TRACE2.TRACE3.TRACEA,3,1) I MARK FIRST s POINTS
    CALL CURV3D TRACE2.TRACEC,TRACE1.3,1)
    CALI CURV3DITRACEB TRACE3,TRACEIT3,
    CALL MARKER(4)
    CALL CURV3D(TRACE2.TRACE3,TRACEA,NPTS,1) ! PLOT &-D PROSECTIONS
    CALL CURV3D TRACE2,TRACEC,TRACEINPTS.1
    CALL CURV3D(TRACEB,TRACE3,TRACEI,NPTS.1) ,
```

ENDP
C CALL BGNMAT(NPTS,NPTS) : MAKE SD MATRIX

CALL BGNMATINPTS,NPTS
CALL ENDMAT(2MTR,0)
CALL ENDMAT(ZMTR,0) ! ! CALL ENDPL(O) CALL ENDGR(O) END SUBPLOT
continue
CALL DONEPL END


WRITE $2.9^{\circ}$

```
c
CCCC LOOP 20 COMPUTES THE AVERAGE VELOCJTIES
        IF(TEST2EQ.1)THEN I COMPUTE AVERAGE VELOCITIES
        DO 20 I= N,N
                        READ(1,5)DEPTH(t).PTT(l),SHTT(1),SVTTY(1)
                            DIST(I)=OFFSET (COS(ATAN(DEPTH{l)/OFFSET))) ! COMPUTE STRAIGHT
                                    ' LINE DISTANCE
                            OF(PTT(I)NE 0)THEN
                    VP(I)=DIST(I)/PTT(I) I P VELOCITIES AND RATIOS
                    SVRATIO(I)=SVTT(I)/PTTYI)
                    SHRATIO(I)=SHTT(I) /PTT(I)
                    SHPOIS(1)=0.5 (SHRATIO(I)*2-2.0)/(SHRATIO(1)**2-1 0)
                    SVPOIS(I)=0.5 (SVRATIO(I)**2-2.0)/(SVRATIO(I)*2-1.0)
    ENDTF
    IF(SVIT(I)NE.O)VSV(I)=DIST(I)/SVTT(I)
    IF(SHTT(I)NE O)VSH(1)=DIST(1)/SHTT(1)
    LF(TEST EQ.2)WRITE(2,11)DEPTH(I),VP(I),VSV(I),VSHII)
    L(TESTEQ 3)WRITE(2,30)DEPTH(I), SHRATIO(l),SIPOIS(l),
        * DONTINEE DEPTH(1),SVRATIO(I),SVPOIS(I)
        CONTINUE
        ENDTF
        IF(TEST2.EQ.2)THEN
C
CCCC LOOP so COMPUTES THE STRAIGHT RAYPATH DISTANCE
    DO 30 I=I.N
            READ(1,S)DEPTTH(I),PTT(I),SHTTTI),SVTT(I)
            DIST(I)=OFFSET /(COS(ATAN(DEPTH(I)/OFFSET)))
    30
C-Ccco
C
    DO }0|{(INT+1).
            CHECK=0
            CHECK2-0
            CHECK3=0
            CHECK4=0
            CHECKS=0
            VP(1)=0
            VSH(I)=0
            VSV(I)=0
            SHRATIO(I)=0
            SVRATIO(i)=0
            SHPOIS(1)=0
            SHPOIS(i)=0
C
CCCC COMPUTE P INTERVAL MEASUREMENTS
            IP(PTT(1).NE O AND.PTTYI-INT).NE.PTT(I))THEN
                    IF(PTT(I-1NT) EQ.O)THESN
                    DO 41 J=1,(1-INT)
                    INT2=| NT +J-\
                    IF(PTT(I-INT2) EQO) QO TO 11
                    LFCHECK EQ.IIGO TO 41
                    VP(1)=(DIST(I)-DIST(I-INT2))/(PTT(I)-PTT(I-INT2))
                    VPDEPTH{1)=(DEPTH{1)-DEPTH(1-INT2))/200+DEPTH(1-INT2)
                    CHECK=1
                CONTINUE
```




FORMAT(3(F1:0.F80))
FORMAT(F82.3F84)
FORMAT(A64)
FORMATIF10 0.3 F163)
FORMAT(4F180)

```
    39 FORMAT(2(F15.0.2F123)
```




```
    99 CONTINUE
    CALL NOMIN8
    CALL PVTE&O
    CALL PAGE(110,85)
    tP(TESTEQ.2)CALL YNAME''INTERVAL VELOCITY (FT/SEC)S'.100)
    [F(TESTEQJ)CALL YNAME('SH POISSON RATIOS'.100)
    C
    IN YNAMET PISH
    CALL XNAME( DEPTH (FEET)$',100)
    CALL AREA2D(8.0,60)
    [F(TESTEQ 3)CALL GRAF(0.,1000.,6000.00.0.05,5)
    TF(TEST EQ 2)CALL GRNF(0.,1000.,6000.00.3000 0,18000 0)
    CALL MARKCER(2)
    CF(TESTEQ.3)THEN
    CALL CURVE/SVRDEPTH.SVPOISISVP,1)
    CALL CURVE(SHRDEPTH.SHPOIS,ISHP,I)
    ENDIF
    C CALL CURVE(SHRDEPTH,SHRA TIO,SRSH,1)
    C CALL MARKER(4)
        F(TEST EQ 2)THEN
            CALL CURVE(VSHDEPTH.VSH.ISH.1)
            CALL CURVE(VPDEPTH,VP.IP.I)
        ENDF
    CALL ENDPLO)
    CALL DONEPL
    WRITEIG.9 \becauseDO YOU WANT ANOTFER PLOT OR MORE DATA! I=PLOT AGAN,
        * 2mMORE DATA, J=THE END
            READ(5, 9COPY
            IP(COPYEQ 1) CO TO &O
            [F(COPYEQ 2) OO TO 100
            CALL DONEPL
            END
            SUBROUTINE ZERO(VEL,VDEPTH,N,INT,ISTEP) ZERO
C
CCccc
CCCCC
C
    THIS SUBROUTINE REMOVES VALUES OF VEL EQUAL TO ZERO
    AND REORDERS VEL AND VDEPTH WITHOUT THOSE VALUES
    REAL VEL(100).VDEPTH(100)
    INTEGER N.INT.ISTEP
    ISTEP =0
    DO 10 Im (NT +1).N
                    LF(VELI) NE 0)THEN
                        ISTEP=ISTEP + I
                        VELISTEP)=VELI)
                                    VDEPTH(ISTEP)=VDEPTH(I)
                    ENDIP
    CONTINUE
    RETURN
    END
```

...2ERO


## PROGRAM ROTBOR

```
Originally written by Fred Eactwood, modifications by Tom Daley
```

", 0 ensen-
thie program compates the ratation of randamity
oriented thres component traces into a
c torthole coardinele ayotem

character filain 84 filoout 64 EIGEN 84

reel al lampa (1000), comph1(1000), temph2(1000)

```
open inpuc fle
```

    type 140. filen
    Cormay' input fleanaon=', a)
    accept 145, Gilem
    Pormata)
    
open output file
65p 240.6leout
orma4' output fileaame=', a)
ceapt 245, bleout
formá(a)

C WRITE(6, $9^{\circ}$ EIGEN VALUE FTLE NAME-
WRUTE(8, 9 EICE
READ 6,146 EIGEN
OPEN(UNIT = 4, FLE-EIOEN,STATUS = 'NEW')
read(2) atrece.
type 250,0trace
250 Pormak' hatars sey ntracem', 1
wrice(3.260) atrace
200 Popmad hoaders atrace ${ }^{\circ}$ ', I)
c colve syotem
do 1000 i=l,atrace/3 ! ! reod in data traces
reed (2) nastop
resed 2$)$
$($ traces $(j), j-1$, manap $)$
reed (2) namip
read 2) (uncab $(1), j=1,0 \operatorname{momp})$
reed (2) nsamp
reed (2) (uraceh2(1), $\mathbf{j = 1 , a m p}$ )
c
c calculase stre lay anse and eroce corvelationa
e

c normadise croos corrolation malriz
c

phi $=0.0 \quad$ I SET PHI $=0$ FOR BOREHOLE COORDINATES
type 800,i,phi,theta, ellip
800
 write( 3,900 ) i.phi, thete, ellip
$\mathbf{9 0 0}$
1000
continue
CLOSF2 2
CLOSD(3)
Close(4)
anp
and
subroutise polar(crom, phi,theta, ellip)
polar
pral 04 crom(3,3), xinven(3), cubid(3), croot(3),evect(3)
$\mathrm{pi}=3.1415986$
e
c
c

- -xinvar(3)
cubia 2 $=$ minver(3)
cubia(3) $=-$ Itivar(1)
call reubic/cubic, croot)
call mortarsel (eroot)
e
$\operatorname{crom}(1,1)=\operatorname{crom}(1,1)-\operatorname{croov}(1)$
$\operatorname{crom}(2,2)=\operatorname{crome}(2,2)=\operatorname{croov}(1)$
$\operatorname{crom}(3,3)=\operatorname{crom}(3,3)-\operatorname{crook}(1)$
c

```
    DO 30 JJ=1.3
    WRUTEX4.9 CROOT(Jd)
    OONTINUE
    call doovect(ercen, ovect)
    If (rveck(1).18.0.0) chee
        srech(1) - -बrect(1)
        ovec(2) =-\operatorname{arec}(2)
        svec(3) = -बvecu(3)
        andIf
```


## polar

    IF TRACES HAVE bEEN EDITED SET
    THETA =00
    PHI
    IF (EVECT(2)EQ 0)GO TO 8
    LF(EVECT(3).EQ. 0.) GO TO 10
    chets = 1800* atan2(evec<(3), evec<(2)) / pi
    phi =1800.acos(evecu(1)/ xmag(evect)) / pi
    ellip = 10-\operatorname{croot(2)/croou(1)}
    ratuen
    THETA - = 
    PHI =0
    ELLIP}=
    ELLPP}=
    end
    subroutine calcinv(matrix,invar)
    reale4 matrix(3,3),invar(3)
    Inver(1) = matrux(1, 1) + matrix(2, 2) + matrix(3, 3)
    inver(2)=0.0
    do 200 1= 1,3
                do }100\quadi=1,
                                    Iavar(2)=1avar(2)+05 (matrix(i, i) * matrix(j, 1) -
                                    matrxi(i, )) (matrix(i, 1))
        contlaue
    contlaue
    lave(3) = dot(matrix)
    renura
    fuseclon dete) det
    meal 4, <(3,3)
    tarm1 = (2, 2) * (3, 3) - (2, 3) : (3, 2)
    tarm2=\(2, 1) :(3, 3) = (3, 1) :2(2, 3)
    tarm2 =a(2, 1) (a, 3) =(3, 1) (2(2, 3)
    det - (1, 1) * lamal - (1, 2) - vom2 + © (t. 3) * varm3
    rasupt
    end
    mubroutise reubic (a,x) rcubic
    mala4 2(3),2(3)
    pime3.1415928
                                    calcinv
    do (2)=0.0
    I
    ```
```

pion 3.1415926

```
```

    \(p=(2)-(2(3) \bullet 2(3) / 3.0)\)
    ```
    \(p=(2)-(2(3) \bullet 2(3) / 3.0)\)
    \(9=(2.0 \cdot 2(3) \cdot 03 / 270)-(2(2) \cdot 2(3) / 3.0)+2(1)\)
    \(9=(2.0 \cdot 2(3) \cdot 03 / 270)-(2(2) \cdot 2(3) / 3.0)+2(1)\)
    dale \(=770^{-1} q^{\circ O 2}+40 \cdot p^{\circ O}\)
    dale \(=770^{-1} q^{\circ O 2}+40 \cdot p^{\circ O}\)
    If (delte gs. 0.0) then
    If (delte gs. 0.0) then
        type 100
        type 100
        formay this cubic hat two complex roots, and one real roor')
        formay this cubic hat two complex roots, and one real roor')
            phi \(=\operatorname{atan} 2(-a q r 4(-\) delta \(/ 270)\). a
            phi \(=\operatorname{atan} 2(-a q r 4(-\) delta \(/ 270)\). a
```

    olve
    ```
```

    olve
    ```

\section*{if(philt.0.0) \(p h=p h 1+p i\)}
```

$$
\begin{aligned}
& \operatorname{amp}=2.0 \cdot \operatorname{sqrt}(-p / 3.0) \\
& x(1)=\operatorname{amp} \cdot \cos (p h / 30)-2(3) / 30 \\
& x(2)=-\operatorname{amp} \cdot \cos (p 1-p h i) / 3.0)-2(3) / 3.0 \\
& x(3)=-\operatorname{amp} \cdot \cos (p i+p h i) / 30)-a(3) / 3.0
\end{aligned}
$$

endif
c

## petupn

```
end
cubroutine sorteigval (a) sorteigual
c. reslot \(x(3)\)
c
de \(200 \quad 1=1,2\)
do \(100 \mathrm{j}=1.2\)
in a \((j) \cdot \mathrm{lt} \cdot(j+1))\) call swapscalar(a(i) \(a(j+1))\)
100
200
continue
return
end
eubroutine swapecaler ( \(\mathrm{s}, \mathrm{b}\) )
c
nmp =*
\(\therefore=b\)
petura
end
subrousloe doerect (s,evect)
c
real a4 a(3,3), evect(3)
lacerer 4 sadoy(3)
c
c initidise indesce
inders 1)=1
\(i n d=3)^{2}=-2\)
\(i n d y\)
c
c aval columes and rowe if neoded\}
If (a(1,1) aq. 0.01 cell swap(1, 2, 6, iadax)
call calcoigonvect (s, croct, ifas)
If (fing ct. 0) thea
cell swep(2, 3, 3, iadex)
eall calcergenvecu(a evect, Itas)
eadif
c
c aneramble rigenesctor
c
eall uneerambláarect, index)
petura
and
on broutlee map (k, l, a, index)
swap
read an \(2(3,3)\)
Inceser as inderi3)
c
\(c\) cere colsmane
de \(2001=1.3\) \(\operatorname{lump}=2(i, k)\) \(2(1, k)=2(i, 1)\)
```

```
    200 continue a(i. ) = comp
                                    ..swap
    c
    c ewag rowe
do 300 i=3,3 (emp =a(k, 1)
    mpdate indezes
        i= Index(k)
        index(k) = Index(l)
        Madex(1) =1
        c {\begin{array}{l}{\mathrm{ end }}\\{\mathrm{ enbroutine calceigenvect (s. al, ifmac)}}\end{array})
        Prala4 a(3,3),a((3),&e(3),anavorm(3,3)
        imace=0
    sinitialise cinorrce
        do 100 1=1.3
            do 50 l=1.3
        conclaue
            conclaug
            mavora(1, 1) = 1.0 /a(1, 1)
            c do 500 a= 1,2
c calcwate co and of
                0N =000,0
                    *(1)}=0.
                        ak(i) =0.0
                        l}=2(i,a+1
                        do 300 imi,a
                        B=a(a+1, )
                                G=2(1,a+1)
                                ea(i) = ea(1)+ei - aunverse(i, 1)
                                a(d) =a(l) +a.auserse(i, i)
300
+\infty
        concloue
            continue eal = eal + O| maverma(. i) e ij
        a(a + i) = -10
        en(a+1)=-1.0
        If(a lt 2) : = a(a + l.a+1)
c
ce check for divide iy zero
    if (abds - eal) It. 0.005) then
        1428 =a
    ondif
```

        \(z=1.0 /(\mathrm{g}-\mathrm{ead})\)
        anversed \((i, j)=00\)
        do \(450: 1=1, a+1\)
        do \(425 i=1, a+1\)
        anversed \((i, j)=\operatorname{araverse}(i, i)+z \cdot a l(i) \cdot \operatorname{az}(j)\)
        continue
    cantinue
        continue
        return
        end
        aubroutlae unscramble (vect, index) unscramble
        realet vect(3).vtamp(3)
        integer at inderi 3 )
        do \(100 i=1.3\)
        vtemp(indox \((i))=\) vect \((i)\)
    continue
    do \(2001=1,3\)
        vecu(i) \(=\) vtamp \((i)\)
        contlaus
        peturn
        and
            Tunction umes (vect) Imag
            realat vece(3)
            ump \(=00\)
            do \(100 \cdot 1=1,3\)
            temp \(=\operatorname{tamp}+\operatorname{vect}(1) \cdot \operatorname{rect}(1)\)
            continue
            \(x\) cms \(=\) squ(temp)
            retura
            and
            aubrouthae docrom(tracel,trace2, amamp,crom) docross
            real at wecel (namp), trace2 (nmenp)
            cromemo. 0
            do 1000 i=1,4emp
            crosecroathtracel(i) Arace2 \({ }^{(i)}\)
            continue
            return
            and
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