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A DUAL-WAVELENGTH THERMAL INFRARED SCANNER AS A POTENTIAL
AIRBORNE GEOPHYSICAL EXPLORATION TOOL

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We are investigating a new airborne method for measuring surface temperatures that may be useful for identifying thermal anomalies of geologic origin. From Planck's equation we derive the valuable approximation that, for small temperature variations, the radiant emittance is proportional to the emissivity times the absolute temperature to the power of $(50/\text{wavelength in } \mu\text{m})$. From this, expressions are obtained for the emitted infrared (*ir*) radiation measured simultaneously in the 5 and 10 μm bands. Ratios of these expressions are shown to have the following useful properties at 288 K: (a) they are insensitive to surface emissivity variations for vegetated terrain, (b) they vary nearly as the 5th power of the surface temperature, and (c) they distinguish emissivity-related from temperature-related effects. We have made preliminary tests of this methodology at a field site in Scipio Center, New York. We have characterized the observed surface temperature variations, the significant effects of soil moisture, and separated out the purely emissivity-related features of vegetated terrain. Cluster analysis served to divide the *ir* data into groups that behave similarly as a function of the measured soil moisture. Two such distinct terrain groups were identified at the field site. The *ir* data were corrected for: (a) natural surface emissivity varia-

tions, (b) the intervening atmospheric path, and (c) the reflected sky radiation. The corrected surface temperature data were compared with calculated values computed from a model that simulates the surface temperature, using meteorological, hydrological, topographical, and thermal input parameters. The simulated mean surface temperatures, 291.9 K (group 1) and 291.8 K (group 2), differed only by, respectively, 0.0 K and 0.1 K from the measured mean surface temperatures. Our preliminary results suggest the potential for developing a new airborne geophysical method for isolating abnormal heat flows. Weak heat flows, about 10-20 times the terrestrial average, have the effect of raising the surface temperature about 0.1-0.2 K. These temperature anomalies would, with the methodology suggested, appear as a residual difference between the measured (corrected) surface temperature and the simulated surface temperature. Such surface temperature differences appear, from our research, to be measurable by airborne *ir* scanners when data over surface areas of 0.1 km^2 or larger are averaged. Accordingly, our research appears to support the conclusion that surface temperature enhancements of geophysical origin between 0.1 and 0.2 K can be identified using airborne infrared methods.

INTRODUCTION

It has long been recognized that the measurement of geophysical parameters of the earth by airborne means can provide rapid and relatively inexpensive preliminary surveys of geologic prospects. Aerial magnetic and scintillometer surveys are examples that come to mind most immedi-

ately. Although other airborne geophysical measurements have been made with greater or lesser success, the airborne measurement of surface temperature appears to be fraught with complicating factors. On the other hand, rapid wide area mapping of surface temperatures would be very useful. A number of geologic phenomena worthy of com-

Manuscript received by the Editor January 20, 1976; revised manuscript received May 5, 1976.

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Investigation are associated with the generation of anomalous heat; the oxidation of ore and increased heat flow owing to potential geothermal resources are examples. Accordingly, it seems desirable to increase our confidence in airborne geophysical methods to include the airborne measurement of minute surface temperature differences.

Measuring the time-dependent true surface temperature from an aircraft and extracting the geologic and comparatively very small steady-state temperature component of geologic origin is recognized as a complicated task. On the other hand, mapping surface temperatures by airborne survey methods is an established technique. The qualitative airborne infrared (*ir*) measurements made over geothermal anomalies by other workers have lent credence to the possibility of measuring small surface temperature differences from the air for geophysical exploration. This recently became feasible. Using an *ir* scanner operating in the 3–5 μm range, Dickinson (1973) has provided evidence of new geothermal areas in New Zealand and delineated the extent of hydrothermal activity. Measurements of 0.5–1.0 K above ambient (Dickinson, personal communication, 1974) made at depths of 15 cm in several areas outlined by the *ir* imagery confirmed Dickinson's interpretation. In general, surface temperature elevations of 1–3 K at Taupo, New Zealand and 1–2 K at Mt. Amiata, Italy (Pohn et al, 1973), outlined heat flow regions of conductive components as small as 24–45 mW/m² and convective components ranging to 100 HFU. It is uncertain as to what percentage of each component the scanners recorded during overflights. However, it is obvious from their data that few or none of the temperature anomalies recorded in the imagery were due solely to the effects of conductive heat flows, but rather to the combined effects of convective and conductive heat flows. The work of these researchers encourages us to investigate further the potential of airborne *ir* scanners as a geophysical tool and to place particular emphasis on the possibility of making quantitative measurements.

For quantitative measurements, however, the remotely sensed radiation must be corrected for the effects of surface emissivity, meteorology, topography, and surface material differences (Del Grande, 1975) including thermal inertia and albedo (Pohn et al, 1974), and particularly those due to soil moisture (LeSchack et al, 1975) prior

to obtaining any quantitative temperature measurement that could be useful for geophysical exploration. This paper discusses the theory of our approach to making quantitative temperature measurements and analyzes preliminary field data that support this theory. We found a number of references concerning the details of aerial infrared surveys of great value to our studies and they are included as general references.

DISCUSSION OF THEORY OF AIRBORNE TEMPERATURE MEASUREMENTS

As in the case of other geophysical exploration techniques, such as gravity or magnetic surveys, thermal surveys require the detection and isolation of a geologically induced field component that is only a small percentage of the overall field measured. For example, the cumulative extraneous effects that can mask temperature anomalies are often an order of magnitude greater than the sought-for anomalous component. Just as altitude, latitude, Bouguer, and terrain corrections must be made to gravity data, and diurnal and normal corrections need to be applied to magnetic data, large corrections are necessary before useful interpretations can be made from airborne surface temperature data. There are three basic steps in isolating anomalous temperatures of geologic origin: (a) Measure the true surface temperatures over the survey area at a given time, (b) calculate the normal surface temperatures that would be anticipated for the same location and time, and (c) subtract the simulated temperature field from the measured temperature field to obtain a residual or anomaly field.

Measuring the true temperature

The first step in measuring the true temperature over a given geographical area is to record the radiant temperature over this area using a quantitative airborne infrared line scanner capable of sensing emitted energy at two wavelengths, at least. After the data have been recorded, we must then correct the calibrated radiant temperature data for: (a) variations in natural surface emissivity, (b) absorption and reemission in the atmospheric column between the surface and the scanner, and (c) reflected sky radiation due to the nonblackness of the terrain.

Estimating emissivity from ir signal ratios

It was recently shown (Del Grande, 1975) that the earth's surface radiates signals which for small

temperature changes vary as the temperature T to the power $C_2/\lambda T_0$, where λ is the wavelength of the radiation, in μm , T_0 is the surface temperature in degrees Kelvin, and C_2 is a constant. At temperatures near 288 K, the spectral radiant emittance is proportional to the surface emissivity ϵ_λ times $T^{50/\lambda}$. This power law thermal model provides the physical rationale for ratioing narrow ir bands to produce signal ratios that are insensitive to the natural surface emissivity and have enhanced thermal responsivity. These signal ratios are used to obtain precise temperature measurements that are more easily identified with their geophysical origin. The derivation of this model, which is given in the Appendix, will be summarized below.

Planck's equation gives the hemispherical, spectral radiant emittance W_λ , measured in units of $(W \cdot m^{-2}/\mu\text{m})$:

$$W_\lambda = \epsilon_\lambda C_1 \lambda^{-5} [(\exp C_2/\lambda T) - 1]^{-1}, \quad (1)$$

where

W_λ = radiant emittance per unit wavelength at wavelength λ ,

ϵ_λ = surface emissivity at wavelength λ ,

C_1 = a constant, $3.7414 \times 10^8 (W/m^2) \cdot \mu\text{m}^4$,

C_2 = a constant, $1.4388 \times 10^4 \mu\text{m}(K)$,

λ = wavelength in μm , and

T = surface temperature in K.

It is shown in the Appendix that (1) may be approximated as follows:

$$W_\lambda(T) = C_\lambda(T_0)(T)^{C_2/\lambda T_0} \propto \epsilon_\lambda T^{C_2/\lambda T_0}, \quad (2)$$

with $C_2 = 14388 \mu\text{m} \cdot K$ and a nominal value of $T_0 = 288 \text{ K}$, $C_2/T_0 = 49.96$, equation (2) can be closely approximated by:

$$W_\lambda = C_\lambda(T_0) T^{50/\lambda} \propto \epsilon_\lambda T^{50/\lambda}. \quad (3)$$

Accordingly, we have derived a convenient expression for calculating the temperature sensitivity of aerial infrared scanning systems. At a typical predawn temperature of 288 K, we see from equations (1) and (3) that the detectable radiation emitted over the wavelength band and centered at λ is proportional to $\epsilon_\lambda T^{50/\lambda}$.¹ Filters are used to define narrow bands. A few bands avoid both atmospheric absorption regions and wave-

¹ We note that the commonly used temperature power law is T^4 (resulting from Stefan-Boltzmann's equation). This describes the temperature response of the total radiant emittance [i.e., the integral of equation (1) over all wavelengths] as well as for a band centered near $12.5 \mu\text{m}$.

lengths associated with anion groups in common minerals where the terrain behaves as a spectral source. These bands are centered at 2.2, 3.5, 4.8, and $13.2 \mu\text{m}$. At these wavelengths, the signal responses vary, respectively, approximately as the 23rd, 14th, 13th, 10th, and 4th power of the absolute temperature.

The earth's emissivity is highly variable from one location to the next. However, the ratio of signals at two or more wavelengths can be used to obtain accurate surface-temperature measurements that depend very little upon emissivity variations. For two bands at wavelengths λ_1 and λ_2 , the temperature response of the signal ratio is:

$$\frac{W_{\lambda_1}}{W_{\lambda_2}} = \frac{\epsilon_{\lambda_1} T^{C_2/\lambda_1 T_0}}{\epsilon_{\lambda_2} T^{C_2/\lambda_2 T_0}}. \quad (4)$$

Variations in the emissivity ratio are smaller by a factor of ten or more than variations in the absolute emissivity for different natural terrains. At wavelengths where the radiation has the same spectral slope as a blackbody source, the emissivity ratio variations are very small. Thus, quantitative temperature measurements can be made using signal ratios that are calibrated against a standard blackbody source.

Where the surveyed terrain behaves in true graybody fashion, i.e., no matter what material is scanned in the survey, the ratio of $\epsilon_{\lambda_1}/\epsilon_{\lambda_2}$ is constant, then the ratio of two signals of differing wavelengths can be calibrated and the blackbody temperature obtained. In this work we let $\lambda = 5 \mu\text{m}$ and $10 \mu\text{m}$, respectively, because these are the common wavelengths that are recorded by present-day scanners and, as such, are the wavelengths in which our field data were obtained. However, energy radiated in the $10 \mu\text{m}$ band does not generally exhibit graybody behavior (Del Grande, 1975) so that for future surveys another wavelength, e.g., $13.2 \mu\text{m}$, would be more appropriate. On the other hand, as discussed below, where the terrain surface is covered by vegetation, radiation in the $10 \mu\text{m}$ band does appear to behave in a graybody fashion. This permits us to develop the following relationships which apply at about 288 K, a nominal field value for the surface blackbody temperature T_b from equation (4):

$$\frac{W_5}{W_{10}} = \frac{\epsilon_5 T_b^{10}}{\epsilon_{10} T_b^5}, \quad (5)$$

and

$$\left(\frac{W_5}{W_{10}}\right)^{1/5} = \left(\frac{\epsilon_5}{\epsilon_{10}}\right)^{1/5} T_b. \quad (6)$$

implies that by ratioing the 5 and 10 μm measurements and calibrating the resulting signal, a direct measurement of T_b , the blackbody temperature, can be obtained. For the specific problem of determining the emissivity necessary to correct measurements which was calibrated assuming a blackbody surface, the following relationships are obtained from equation (3):

$$T_{10} \propto W_{10}^{\lambda/50} \propto \epsilon_{10}^{1/5} T_b, \quad (7)$$

$$T_5 \propto W_5^{\lambda/50} \propto \epsilon_5^{1/10} T_b. \quad (8)$$

By raising equations (7) and (8) to the 10th power and dividing them, we derive the following expression for the surface emissivity ϵ_{10} :

$$\left(\frac{T_{10}}{T_5}\right)^{10} \propto \frac{\epsilon_{10}^2}{\epsilon_5} \propto \epsilon_{10}. \quad (9)$$

Equation (9) holds for vegetated surfaces which exhibit graybody behavior at 5 and 10 μm , i.e., the emissivity for these surfaces is constant. We have assumed that graybody behavior does pertain to the surfaces that we have measured since they are all vegetation covered, and from the data of Tanskanen and Tantraporn (1952) it can be shown that ϵ_{10}/ϵ_5 for 40 different vegetation samples is constant to better than one percent. This implies that ϵ_{10}/ϵ_5 is not necessarily unity, it is constant for a variety of different vegetation types, and, thus, can be determined by calibration measurements.

Mathematical and physical basis for atmospheric corrections

Atmospheric corrections for radiant temperature measurements must be made both for the absorption and emission in the intervening air (between the surface and the scanner) and for the reflected sky radiation due to the non-reflectivity of the terrain.

Tien (1974) has established the physical and mathematical basis for these corrections. Near the surface, the radiant emittance W_g is the sum of the surface radiation and the reflected sky radiation:

$$W_g = \epsilon_\lambda W_b + (1 - \epsilon_\lambda) W_s. \quad (10)$$

The subscripts g , λ , b , and s refer, respectively, to the calculated (graybody) temperature at the surface, the wavelength band, the directly measured (blackbody) surface temperature, and the

temperature associated with the sky and cloud cover. The emissivity is given by ϵ_λ and the reflectance coefficient by $1 - \epsilon_\lambda$.

At some distance h , above the terrain, the measured radiant emittance W_h consists of the transmitted surface radiance and the emitted radiance of the intervening column of air. Thus,

$$W_h = W_g (1 - A) + W_a A. \quad (11)$$

The subscript a refers to the temperature of the air column. Most previous studies approximate the temperature of the air column as the value measured for the air temperature outside the aircraft at the altitude of the scanner. These studies will be used to estimate a value for A , the absorption or emission of the air column appropriate for the conditions of our experiment to be discussed below.

Tien (1974) shows that by eliminating W_g , together with appropriate algebraic manipulations, equations (10) and (11) can be combined as follows:

$$W_b - W_h = \frac{1}{\epsilon_\lambda} \left(\frac{A}{1 - A_\lambda} \right) (W_h - W_a) + \left(\frac{1 - \epsilon_\lambda}{\epsilon_\lambda} \right) (W_h - W_a). \quad (12)$$

He further applies a two-term truncated Taylor series expansion of the $W(T)$ around $W_h(T_h)$ on both sides of equation (12):

$$W(T_h + \Delta T) \approx W_h(T) + \Delta T W'_h(T_h). \quad (13)$$

By canceling out terms of the form $W_h(T_h)$, and dividing both sides of the subsequent equation by $W'_h(T_h)$, he obtains:

$$T_b - T_h \approx \frac{1}{\epsilon_\lambda} \left(\frac{A}{1 - A_\lambda} \right) (T_h - T_a) + \left(\frac{1 - \epsilon_\lambda}{\epsilon_\lambda} \right) (T_h - T_s). \quad (14)$$

This approximation is valid when the temperature correction is small compared to the absolute surface temperature, as is typically the case. Equation (14) is in agreement with the experimental observations reported previously by Weiss (1971), Saunders (1970), and Shaw and Irbe (1972). Most of these measurements were made over ocean surfaces which had an emissivity of 0.986 near 10 μm . Equation (14) permits us, therefore, to correct directly for the blackbody temperatures at the earth's surface.

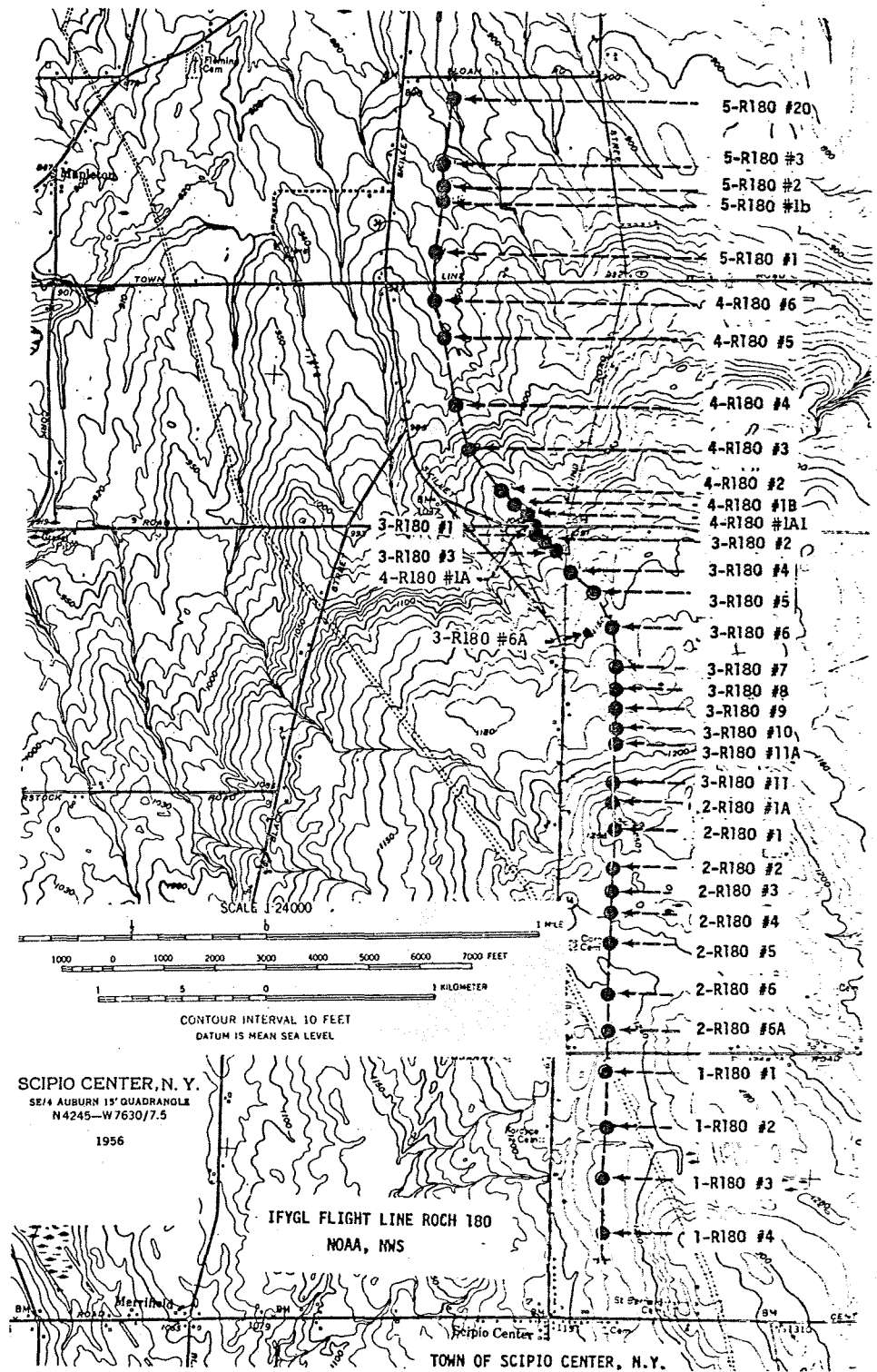


FIG. 1. Locations of soil sample sites (after Peck et al, 1972).



Fig. 2. Dicolor images of Scipio Center, N.Y. site (see Figure 1). $10\ \mu\text{m}$ image on left, $5\ \mu\text{m}$ image on right. The radiant temperatures range as follows: white = all above 59°F ; red = $58^\circ\text{--}59^\circ\text{F}$; yellow = $57^\circ\text{--}58^\circ\text{F}$; green = $56^\circ\text{--}57^\circ\text{F}$; cyan = $55^\circ\text{--}56^\circ\text{F}$; blue = $54^\circ\text{--}55^\circ\text{F}$; magenta = $53^\circ\text{--}54^\circ\text{F}$; black = all below 53°F .

theory, the above steps take the raw but uncorrected radiant temperature data obtained as directly from the scanner, make approximate corrections to them for variations in surface emissivity, and finally, enable corrections for the atmospheric column to be made. A map of such surface temperatures, although perhaps accurate for the period of time during which the data were recorded, is, however, transitory. Our next step, therefore, is to calculate, on a point-by-point basis over the same survey area, the "normal" surface temperatures (i.e., the temperatures that would be obtained under equilibrium conditions due to all environmental causes other than abnormal heat fluxes). We accomplish this with a digital equilibrium surface temperature simulator model.

Calculating the "normal" surface temperature

Outcalt (1972a, b, c, d) has demonstrated the application of an equilibrium surface temperature simulator to predict, with reasonable accuracy, the theoretical surface temperature for any given location on the terrain. The precision of the model is initially adjusted to a threshold of $1 \text{ mly} \cdot \text{min}^{-1}$ or 17 HFU^2 and this threshold could be further reduced. Outcalt (1972b) has shown that the model is sufficiently accurate to replicate surface temperature differences on the order of $\pm 0.2^\circ\text{C}$ associated with different types of terrain as measured by airborne *ir* surveys.

Outcalt (1972b) describes the model as follows:³

In brief, the operation of the general simulator is based upon the energy conservation equation which states that the four components of the energy budget (net radiation R , soil heat flux S , sensible heat flux H , and latent heat flux L) must have a zero sum across a surface:

$$R + S + H + L = 0 \quad (1)$$

In turn each of these terms is a complex function of the environmental variables which specify the radiation and thermal properties of the atmosphere and substrate media. At any instant in time these components may be represented as functions of a limited set of environmental variables and physical constants. These controlling variables are listed with their notation in Table 1.

The components of the energy budget equations can then be written in terms of these variables and the surface temperature (T) as

$$^3 \text{HFU} = 1 \mu \text{cal} \cdot \text{cm}^{-2} \cdot \text{S}^{-1} = 0.042 \text{ W} \cdot \text{m}^{-2}.$$

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$$R = f(\text{LAT}, \text{DEC}, D, R, \text{ALBEDO}, W, P, \text{TSKY}, T) \quad (2)$$

If the assumption is made that the soil temperature at the diurnal damping depth is approximately equal to the mean air temperature, then

$$S = f(\text{GEC}, \text{GD}, \text{TA}, T) \quad (3)$$

The turbulent transfer terms which are corrected for stability using the Richardson number may be expressed as

$$H = f(U, \text{ZO}, P, \text{TA}, T), \quad (4)$$

$$L = f(U, \text{ZO}, P, \text{RH}, \text{SRH}, \text{TA}, T) \quad (5)$$

Table 1. Environmental input variables

Station pressure (P)	Mean air Temperature (TA)
Latitude (LAT)	Mean Air Relative humidity (RH)
Solar declination (DEC)	Mean wind speed (U)
Dust particles cm^{-3} (D)	Soil thermal diffusivity (GD)
Orbital radius vector (R)	Soil volume
Surface Albedo (ALBEDO)	Heat capacity (GC)
Precipitable water (W)	Surface roughness length (ZO)
Sky radiant temperature (TSKY)	Soil surface wetness (SRH)

Note that in all of the above equations after specification of the input variables the surface temperature is the only unknown. The soil temperature profile is allowed to evolve by calculating a finite-difference solution from the preceding step. After the new soil thermal profile is calculated, the soil heat flux [equation (3)] is actually calculated from the uppermost soil temperature level in place of TA .

At each step through the diurnal cycle the solar radiation incident on a surface may be calculated for a clear day by means of a subroutine. Subroutines are also included to calculate specific humidity gradients, to fix the free air computation level, and to correct the thermal properties of the atmosphere for stability.

It is apparent that if a sequence of guesses as to the value of the surface temperature are entered into the equation the correct guess would bring the energy budget equation [equation (1)] to zero; that correct guess would be termed the equilibrium surface temperature. The equilibrium surface temperature is that temperature guess which produces a suitably small residual in the energy budget equation (e.g., $1 \text{ mly} \cdot \text{min}^{-1}$). Then all the components of the energy transfer regime (R , S , H , L) and the soil temperature vector are equally good

guesses, and the next iteration begins with a forward solution of the finite-difference form of the soil thermal diffusion equation.

The output of this model is the absolute temperature as a function of solar time for each survey point. It can be seen that if the theoretical point-by-point temperatures as computed from the above model are subtracted appropriately from the temperature data array that would be generated from the procedures previously described, the residual should be the thermal anomaly field sought.

ANALYSIS OF PRELIMINARY FIELD DATA

On 11 October 1972, soil moisture data were collected along a 7 km line near Scipio Center, New York, while concurrent 4.5–5.5 μm airborne *ir* data were recorded at an altitude of 2000 m above mean sea level (Peck et al., 1972). These data, along with simultaneous multiband photographs, were gathered by the National Oceanic and Atmospheric Administration in conjunction with the International Field Year for the Great Lakes program. These data, recorded over an area presumed to have no abnormal heat flow, were used to make a preliminary test of our theoretical approach to airborne surface temperature measurements discussed above. Adequate information derived from these measurements, from local National Weather Service records, and from the literature permit obtaining the arguments outlined in Table 1 needed to evaluate the surface temperature simulator.

Data collection

A Daedalus Enterprises Dual Channel Scanner, model DS 1230, was used to gather simultaneously the 4.5–5.5 μm and 8–12 μm data.⁴ The flight line (no. 1.1) was flown in a north-south direction over the study profile line. The profile line parallels Skillet Road and is offset 300 m (± 25 m) to the east of it (Figure 1). The airborne data were gathered at an altitude of 1700 m above mean terrain on 11 October 1972 at 1527 hours local solar time [1519 hours local mean time (EST)].

The weather was sunny with high cirrus and scattered altocumulus clouds and 20-mile visibility. The ambient air temperature was 11°C at 1 m above the ground surface. The mean wind speed was 10 knots. Soil samples were taken from 10 cm beneath the surface by auger. The soil moisture

⁴ Mean values, i.e., 5 μm and 10 μm will be used for simplicity in the discussion that follows.

percent by weight in each core sample was determined later. A general description of the vegetation was made at each soil sampling site. Specific soil data were not collected, but general descriptions can be obtained from Soil Conservation Service maps.

The *ir* data were recorded simultaneously in the 5 and 10 μm wavelengths on magnetic tape. Calibrated digitized images were then prepared.⁵ The quantitative "Digicolor" format presents the radiant emittance data as six discrete levels, each level corresponding essentially to a 1°F (= 0.6°C) change of blackbody temperature from 53°F to 59°F (11.7°C–15.0°C). The calibration is established by two blackbody reference sources that are an integral part of the IR scanner used for this work. The blackbody reference sources are adjusted so as to straddle the radiant temperature values observed in the given airborne mission. For the Scipio Center flight line, reference values of 10°C and 20°C were used. The blackbody reference sources are calibrated in the laboratory before each mission by scanning, in the 10 μm range, a water bath⁶ 1 m from the scanner. The temperature of the water bath is varied from 5°C to 40°C in 2 degree steps. The radiant emittance data recorded by the scanner can, therefore, be divided into discrete levels corresponding to calibrated blackbody temperatures. This does not imply that these are the exact or directly measurable ground surface temperatures, but rather the radiant ground surface temperatures modified by (a) the surface emissivity, (b) the absorption and reemission of the intervening air column between the scanner and the ground, and (c) the sky radiation reflected from the ground. These effects must be corrected as discussed above.

Quantifying the calibrated Digicolor images

The calibrated radiant emittance data are presented as images on 70 mm color film (Figure 2). Each 1°F (0.6°C) level of radiant temperature from 53°F (11.7°C) to 59°F (15.0°C) is displayed in a different color. Everything lower than 53°F (11.7°C) is black and everything higher than 59°F (15.0°C) is white. The 70 mm Digicolor film strips for both the 5 μm and 10 μm wavelengths were optically enlarged to convenient dimensions and the soil profile line was overlain on the imagery. The radiant temperature values along the profile line were then recorded.

⁵ Digicolor Images by Daedalus Enterprises, Inc. presented on 70 mm color film.

⁶ The water is presumed to have an emissivity ϵ of 1.

Table 1. Emissivity (ϵ) of common vegetation (after Wolfe, 1965)

	3-5 μm	8-13 μm
Mountain laurel	= 0.90	= 0.92
White pine leaf (dry, top)	0.94	0.96
White pine (dry, top)	0.90	0.90
White pine (dry, bottom)	0.86	0.94
Deciduous maple leaf (dry, winter color—oak)	0.87	0.92
White pine (dry, top)	0.90	0.92
Coniferous twigs (jack pine)	0.96	0.97
Grass (meadow fescue (dry))	0.82	0.88
Northern red oak	0.90	0.96
North American jack pine	0.88	0.97
Colorado spruce	0.87	0.94

Radiant temperature data corrections

Emissivity corrections can be directly estimated when the emissivity ratio (in this case ϵ_{10}/ϵ_5) is constant. From equation (9) we obtain the following:

$$\left(\frac{T_{10}}{T_5}\right)^{10} = k\epsilon_{10} \propto \epsilon_{10}, \quad (15)$$

where k is a constant.

Then, using equation (15) we determine surface emissivity values at 18 locations in group 1 and 16 locations in group 2. The constant of proportionality k for these two groups was determined from published data. The mean value of ϵ_{10} for 11 vegetation types (Table 1) cited by Wolfe (1965) is equal to 0.935. We assumed ϵ_{10} for our data groups had the same mean value, and obtained the following equations from which the emissivity values were determined.

Group 1

$$\epsilon_{10} = .9593 \left(\frac{T_{10}}{T_5}\right)^{10} \quad (16)$$

Group 2

$$\epsilon_{10} = .9582 \left(\frac{T_{10}}{T_5}\right)^{10} \quad (17)$$

The emissivity values for each location along with

examination of Figure 1 shows that the data support our theory is taken over an area of 10 km². In a typical regional geophysical survey of an area of this size would generally contain only a few data points representing sections of the field around them. Since a cluster sampling technique divided our 38 data samples into two natural terrain groups (group 1 having live and dying vegetation, group 2 having live vegetation) (LeSchack et al, 1975), we elected to use the mean values of temperature associated with each group as an appropriate integrated temperature for comparison for a residual anomaly. Accordingly, our tabulations of observed and calculated temperatures will be listed following our grouping.

Table 2. Group 1—temperature data for dry vegetation-covered areas.

Location	Soil moisture, percent	T_b , K	T_c , K	ϵ_{10}	$\left(\frac{.3228}{\epsilon_{10}}\right) \times (T_{10} - 277.0)$	$\left(\frac{1 - \epsilon_{10}}{\epsilon_{10}}\right) \times (T_{10} - 260.7)$	T_5 , K	T_{10} , K
	38.1	292.5	290.2	.939	3.5	1.7	287.9	287.3
	36.6	292.7	293.0	.914	3.5	2.4	288.2	286.8
	34.9	293.0	291.4	.907	3.5	2.7	288.4	286.8
	32.6	292.5	293.1	.939	3.5	1.7	287.9	287.3
	31.8	293.2	290.4	.934	3.7	1.9	288.4	287.6
	31.8	292.5	289.8	.939	3.5	1.7	287.9	287.3
	31.8	292.3	292.9	.924	3.4	2.1	287.9	286.8
	31.7	293.0	291.2	.907	3.5	2.7	288.4	286.8
	31.5	291.0	289.9	.939	3.2	1.6	286.8	286.2
	28.8	292.5	290.0	.939	3.5	1.7	287.9	287.3
	28.6	291.0	292.9	.939	3.2	1.6	286.8	286.2
	28.5	291.5	293.5	.924	3.2	2.1	287.3	286.2
	27.8	291.0	292.5	.939	3.2	1.6	286.8	286.2
	25.4	291.0	289.9	.939	3.2	1.6	286.8	286.2
	24.4	291.5	292.7	.924	3.2	2.1	287.3	286.2
	23.6	291.8	293.5	.942	3.4	1.6	287.3	286.8
	20.5	290.7	292.4	.997	3.3	0.1	286.2	287.3
	16.4	290.2	294.3	.942	3.0	1.5	286.2	285.7
Mean value	29.2	291.88	291.87	.935	3.4	1.8	287.47	286.72
Standard deviation	5.5	.92	1.51	.019	0.2	0.6	0.71	0.53

Table 3. Group 2—temperature data for green vegetation-covered areas.

Location	Soil moisture, percent	T_b , K	T_c , K	ϵ_{10}	$\left(\frac{.3228}{\epsilon_{10}}\right) \times (T_{10} - 277.0)$	$\left(\frac{1 - \epsilon_{10}}{\epsilon_{10}}\right) \times (T_{10} - 260.7)$	T_s , K	T_{10} , K
41a1	46.3	293.0	290.0	.913	3.5	2.5	288.4	287.0
1-1	43.6	291.8	289.3	.942	3.4	1.6	287.3	286.8
4-1a	38.0	292.5	291.4	.939	3.5	1.7	286.5	286.2
1-2	34.2	290.7	289.9	.948	3.1	1.4	286.5	286.2
2-6	34.1	290.7	292.6	.948	3.1	1.4	286.5	286.2
1-4	33.4	290.8	291.8	.922	3.0	2.1	286.8	285.7
3-1	32.8	291.6	293.4	.922	3.2	2.2	287.3	286.2
3-8	29.2	292.5	290.8	.939	3.5	1.7	287.9	287.3
4-1b	29.2	291.5	293.2	.912	3.2	2.4	287.3	285.9
3-11a	28.9	291.6	291.8	.922	3.2	2.2	287.3	286.2
3-7	28.0	292.5	293.3	.929	3.5	2.0	287.9	287.0
3-11	28.0	290.3	290.7	.948	3.0	1.4	286.2	285.9
3-10	27.5	292.5	292.9	.929	3.5	2.0	287.9	287.0
2-5	26.4	289.4	291.6	.967	2.9	0.8	286.2	285.7
2-1a	26.4	291.3	291.0	.932	3.2	1.9	287.0	286.2
2-3	20.9	290.7	292.0	.948	3.1	1.4	286.5	286.2
Mean value	31.7	291.46	291.61	.935	3.2	1.8	287.18	286.42
Standard deviation	6.4	0.99	1.27	.014	0.2	0.4	0.67	0.53

the soil moisture, the radiant temperatures, the blackbody temperatures and the calculated "normal" temperatures are listed in Table 2 (for the group 1 data) and Table 3 (for the group 2 data).

Whereas the published data for ϵ_{10} ranged from 0.88 to 0.97 and had a standard deviation of 0.028, our data ranged from 0.907 to 0.997 with a standard deviation of 0.019 (for group 1) and from 0.912 to 0.967 with a standard deviation of 0.014 (for group 2). The emissivities that we have derived based on (a) the calibrated radiant temperature values, (b) the power law thermal model, and (c) the assumption of a mean value for ϵ_{10} equal to 0.935, appear to be in agreement with measurements made by Hodder (personal communication, 1974). They have a slightly lower average emissivity than the value 0.963, calculated from the data taken by Gates and Tantraporn (1952). Without field measurements of the emissivities at specific locations, the possibility exists that the mean emissivity value at 10 μm may be as much as 3 percent higher.

Corrections for the atmospheric column, equation (14), embodies the analytical technique that we use to obtain the true surface temperature at each point. Rewriting this equation gives us the difference between the surface blackbody temperature T_b for the vegetated terrain and the measured radiant temperature T_h , given by the calibrated radiant temperature T_{10} in our experiment:

$$T_b - T_{10} \approx \frac{1}{\epsilon_{10}} \left(\frac{A_{10}}{1 - A_{10}} \right) (T_{10} - T_a)$$

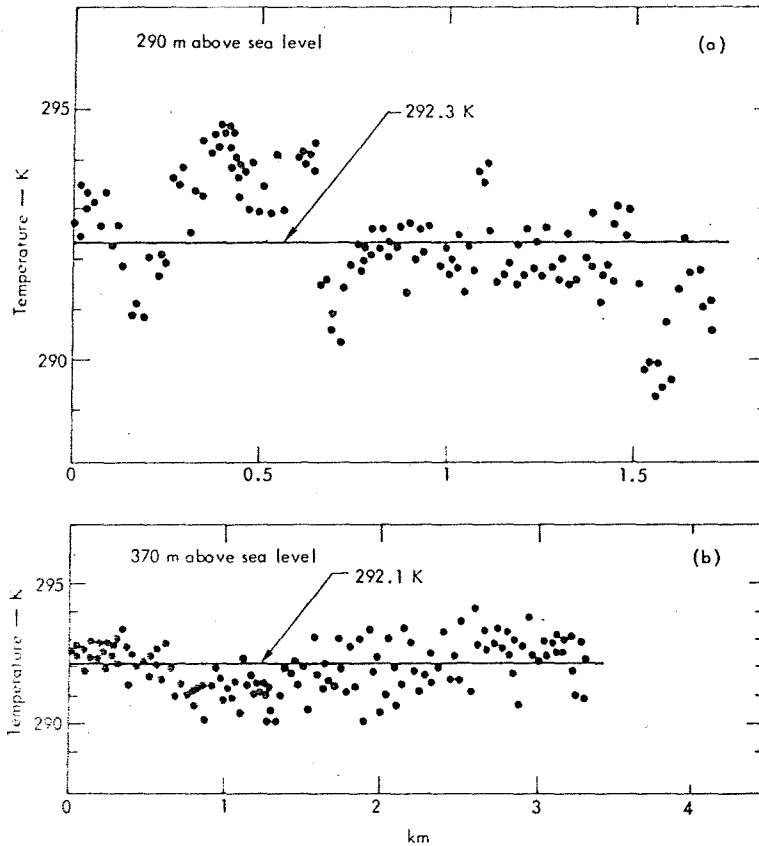
$$+ \left(\frac{1 - \epsilon_{10}}{\epsilon_{10}} \right) (T_{10} - T_a). \quad (18)$$

The first term on the right-hand side characterizes the temperature correction for the intervening atmospheric air path. According to the empirical results of Weiss (1971), the temperature decrease ($T_{10} - T_a$) for an air column of 1700 m is 9.73°C. Consequently we let $T_a = 277.0$ K which is 9.73°C lower than the mean value of T_{10} (for group 1). The absorption coefficient A_{10} was determined to be 0.244, based on the empirical results reported by Saunders (1970) corrected to meet the conditions of our experiment. The coefficient A_{10} , which depends on atmospheric absorption and reemission, is a function of the water vapor mixing ratio which we took as 3.5 gm/kg for a relative humidity of 48 percent. Given a 300 m path, Saunders (1970) computes the effective absorption A_{10} as 0.082 for this mixing ratio. We corrected (linearly) for a path of 1700 m, and reduced the coefficient to compensate for our experimental filter, which had a narrower detection band (about 8-12 μm compared to the 8-15 μm used by Saunders). The latter effect reduces the value of A_{10} by a factor of about 1.9 or less, based on the results of studies made by Weiss (1971). We believe the value taken for A_{10} , based on simplified assumptions for the longer column of air and for the different detection filter, could be as much as 20 percent higher than the value used in our calculations. Without field measurements to verify this atmospheric pa

T_s K	T_b K
288.4	287.0
287.3	286.8
286.5	286.5
286.5	286.5
286.5	286.5
286.8	286.5
287.3	286.5
287.9	286.5
287.3	286.5
287.3	286.5
287.9	286.5
286.2	286.5
287.9	286.5
286.2	286.5
287.0	286.5
286.5	286.5
287.18	286.5
0.67	0.13

$T_{10} - T_s$ (18)

hand side character
on for the intervening
ding to the empirical
temperature decrease
of 1700 m is 9.73°C
0.0 K which is 9.73°C
of T_{10} (for group 1)
 T_{10} was determined
tical results reported
d to meet the cond-
coefficient A_{10} , which
orption and recon-
r vapor mixing ratio
for a relative humid-
00 m path. Saunders
e absorption A_{10} in
e corrected (linearly)
duced the coefficient
imental filter, which
and (about 8-12 μ m
d by Saunders). The
of A_{10} by a factor of
he results of studies
ieve the value takes
assumptions for the
the different detec-
as 20 percent higher
lculations. Without
his atmospheric pr-



3. Corrected surface temperature measurements from ratios of dual-band infrared surveys at 5 and 10 μ m over vegetated terrain. The data for small areas of about 300 m² vary from the mean as much as 2 K at 1527 solar time. By smoothing the data, the average temperatures for large areas of 70,000 (curve a) and 160,000 (curve b) m² were found to differ by 0.2 ± 0.1 K. This difference mainly resulted from normal temperature drop with increasing altitude. Taken from Del Grande (1975).

meter, we must include this possibility in our subsequent discussion of errors. The second term on the right-hand side of equation (18) characterizes the temperature correction for the reflected sky radiation from a nonblack surface. The average sky temperature, contributing to this effect T_s , was taken at 260.7 K to represent the conditions during the survey.

Corrected surface temperature data

A summary of the data and temperature corrections is given in Tables 3 and 4. The location of the data positions, the percentage soil moisture, the corrected surface temperature, and the calculated surface temperature (discussed below) are given in the first four columns. Also included are the emissivity values ϵ_{10} calculated from equations (16) and (17) and the temperature corrections for

the atmospheric path and the reflected sky radiation from equation (18). The uncorrected radiant temperatures at 5 μ m and 10 μ m are given for purposes of comparison. The mean value and standard deviation are calculated for each of these parameters.

We note surprisingly good agreement between the correct mean experimental surface temperature T_b and the calculated mean surface temperature T_c , based on our model that is discussed below. For group 1 (dry vegetation), the experimental value for the mean is 0.01 K higher than the calculated value; for group 2 (green vegetation), the experimental value is 0.15 K lower than the calculated value.

The standard deviations calculated from the data are for each measurement somewhat less than the corresponding standard deviations based

on our model. There are two known effects which would contribute to this. First, the dispersion associated with partitioning the calibrated radiant temperature data into discrete temperatures, each separated by about 0.6 K, does not reflect the natural dispersion. Second, the surfaces with lower than average emissivities cool less efficiently since they radiate less efficiently. Hence, they are apt to be at a higher temperature. Conversely, surfaces with higher than average emissivities tend to be at a lower temperature. In each case the radiated signal, determined by a lower emissivity at a higher temperature or a higher emissivity at a lower temperature, has less dispersion than that of the surface temperature. The corrections which we made for the emissivities ϵ_{10} tend to increase the dispersion, but not to the extent calculated by our model.

It is also interesting to note how the reflected sky radiation and the intervening atmospheric radiation affect the radiant temperature measure-

ments. The reflected sky radiation increases detectable radiation, thus compensating for a nonblack surface with a larger effect for surfaces with lower emissivities, and thus masking the effect of emissivity variations. The atmospheric column introduces an error between the surface temperature and the measured temperature that increases approximately linearly with altitude based on the experimental studies of Weiss (1971).

Errors associated with the corrected temperature data

The effect of increasing the mean emissivity from 0.935 to 0.965 (by 3 percent) would decrease the temperature correction terms in equation (1) by about 1.0 K. If the higher mean value for ϵ_{10} is used together with an absorption coefficient A_{10} of 0.293 (20 percent higher than the value used), the two effects compensate for each other, and the resulting temperature correction is within 0.1 K of the values calculated in Tables 3 and 4. Error

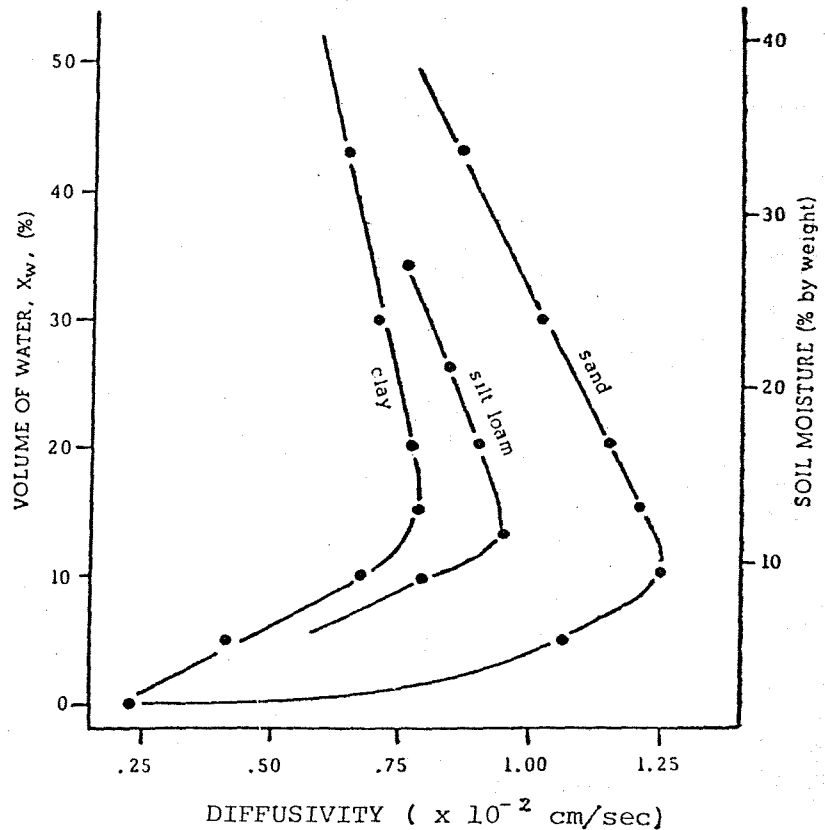


FIG. 4. Soil thermal diffusivity versus soil moisture.

with a ± 10 K uncertainty in the average temperature introduce a ± 0.7 K uncertainty in the correction for reflected sky radiation. A ± 2 K uncertainty in the temperature by the air column introduces a ± 0.7 K uncertainty for the associated correction term. Errors of a random nature associated with instrumentation range from 0.25 to 0.50 K. The overall experimental uncertainties associated with these effects, added in quadrature, are about 1 K. To summarize, uncertain atmospheric conditions could introduce systematic errors of ± 1 K. Natural temperature variations (as shown by standard deviations computed in Tables 3 and 4) are ± 1 K. However, if the mean temperature

were based on 100 vegetated terrain groups instead of 2 (each with about 20 temperature measurements), the mean temperature could be estimated with a precision of ± 0.1 K.

This approach was used to evaluate temper-

Table 4. Input meteorological modeling data to Outcalt model, October 11, 1972.*

Latitude	42.8°
Solar declination	-6.4°
Orbital radius vector	.9982
Dust	1 particle/cc
Station pressure	1018 m b
Precipitable water	9 mm
Mean diurnal temperature	9.5°C
Air humidity fraction	.48
Air wind velocity	447 cm/sec (10 mph)

* The last five values are 24-hour mean values.

Table 4. Continued.

Soil	Soil	Xw	Soil	Emis-	Points	Albedo	Surface	Ex-	Slope,		
diffusivity	heat	vol of	wet	sivity,		percent	roughness	posure**	percent		
(by vol)	capacity	H ₂ O,	fraction*	ϵ_{10}			Z ₀ cm				
cm ² /sec	cal cm ⁻³	percent	percent								
$\times 10^{-2}$	C ⁻¹										
1-R180-4	.70	.66	42	33.4	.94	1-R180-4	20	2	0	0	
3	.78	.53	29	23.6	.94	3	15	.9	0	0	
2	.67	.68	44	34.2	.94	2	20	4	0	0	
1	.63	.81	57	43.6	.94	1	15	7	0	0	
2-R180-6A	.75	.60	36	28.6	.94	2-R180-6A	15	.9	0	0	
6	.69	.68	44	34.1	.94	6	18	.4	0	0	
5	.77	.57	33	26.4	.94	5	22	2	0	0	
4	.90	.44	20	16.4	.93	4	20	.7	0	0	
3	.85	.50	26	20.9	.94	3	20	2	0	0	
2	.80	.62	31	24.4	.93	2	20	2	0	0	
1	.85	.50	26	20.5	.93	1	18	2	0	5	
1A	.77	.57	33	26.4	.93	1A	15	12	180(N)	6	
3-R180-11	.75	.60	36	28.0	.94	3-R180-11	22	3	135(NW)	1.5	
11A	.75	.60	36	28.0	.93	11A	18	7	180(N)	6	
10	.75	.59	35	27.5	.93	10	22	.8	135(NW)	1.5	
9	.76	.56	32	25.4	.94	9	18	12	135	1.5	
8	.72	.63	39	29.2	.94	8	20	4	135	1.5	
7	.75	.60	36	28.0	.93	7	18	.8	135	1.5	
6A	.83	.51	27	21.6	.95	in shade	6A	18	70	135	1.5
>6	.68	.69	45	35.3	.92	not used	>6	18	70	135	1.5
5	.70	.64	40	31.5	.94	5	12	9	135	1.5	
4	.70	.64	40	31.7	.92	4	15	9	135	4	
3	.65	.70	46	36.6	.92	3	18	.9	135	4	
2	.70	.64	40	31.8	.93	2	12	9	135	4	
1	.70	.65	41	32.8	.93	1	22	.4	135	4	
4-R180-1A	.65	.73	49	38.0	.94	4-R180-1A	20	1.1	135	2	
1A1	.75	.59	35	46.3	.92	1A1	20	7	135	2	
1B	.75	.61	37	29.2	.92	1B	22	1.0	135	2	
2	.75	.59	35	27.8	.94	2	18	1.1	135	2	
3	.75	.60	36	28.5	.93	3	18	.9	135	5	
4	.75	.60	36	28.8	.94	4	12	12	135	5	
5	.85	.50	26	20.1	.93	5	12	6	135	5	
6	.70	.64	40	31.8	.93	6	20	1.1	180(N)	1	
5-R180-1	.65	.73	49	38.1	.94	5-R180-1	12	7	180	1	
1B	.67	.69	45	34.9	.92	1B	13	9	180	1	
2	.70	.64	40	31.8	.94	2	13	9	180	1	
3	.70	.64	41	32.6	.94	3	12	7	180	1	

* Measured soil moisture by weight.

** Exposure 0° = south increases to 180° in both east and west directions, positive when directed west of south.

SOIL MOISTURE (% by weight)

10

30

20

10

ature variations associated with changing elevations (Del Grande, 1975) as shown in Figure 3. The data analyzed were based on the Scipio Center survey discussed in this paper. Del Grande discerned the temperature rise from 292.1 ± 0.1 K to 292.3 ± 0.1 K at the 290 m elevation, over that at the 370 m elevation, this 0.2 ± 0.1 K mean temperature difference, found at topographic elevations differing by 80 m, resulted mainly from the expected temperature drop with increasing altitude modified by vegetation and surface moisture differences. Whereas midafternoon temperatures for localized sites smaller than 300 m^2 varied as much as 1 or 2 K from the mean temperature averaged over a larger area (for example about $100,000 \text{ m}^2$), the standard errors of the mean temperatures for areas of $70,000$ and $160,000 \text{ m}^2$ were on the order of 0.1 K. A similar temperature difference of 0.2 ± 0.1 K would be produced by a conductive heat flow intensity of 30 ± 15 HFU. This is lower than the estimated heat flow threshold of 75 HFU believed realistic by some scientists (e.g., Watson, 1974). We believe as a result of the recent studies that it is feasible to identify conductive heat flows of 30 ± 15 HFU or lower.

COMPUTATION OF "NORMAL" TEMPERATURES

Using the surface temperature simulation model outlined above and described in detail by Outcalt (1972a), we have computed the "normal" temperatures that would be theoretically expected at each of the sample locations shown in Figure 1, providing there is no anomalous heat flow.

Estimation of surface environmental parameters

The soil thermal diffusivity was estimated from Figure 4 which was constructed from published data (Baver, 1972; Van Wijk, 1966). The soil volumetric heat capacity (C) was estimated assuming a volumetric mineral fraction (X_m) of 0.47 and a volume fraction of organic matter (X_o) of 0.04, and a variable water content (X_w), assuming a 49 percent porosity by volume. The relationship used is described in Van Wijk (1966) by De Vries and presented here as:

$$C = .46 x_m + 0.6 x_o + x_w \text{ cal cm}^{-3}\text{C}^{-1} \quad (19)$$

The volume fraction was calculated from field data assuming a bulk density of 1.3 gm/cc.

The soil wet fraction or surface relative humidity fraction was used synonymously with soil moisture percent (by weight) that was determined

by laboratory analysis from samples collected in the field at the time of the experiment. The thermal radiation emissivity was determined by methods outlined above. Solar albedo was estimated for the field sites from published data (Chang, 1968; Kondratyev, 1969). Aerodynamic surface roughness was estimated using an empirical relationship developed by Kung and Lettau (see Chang, 1968). Slope and exposure were collected directly from the 7½ minute U.S.G.S. quadrangle of Scipio Center, N.Y. The solar declination and radius vector were obtained from List (1966). A 24-hour mean meteorological data were calculated from raw data observed at the first order NOAA/NWS station at Syracuse, N.Y.

Simulation results

The Outcalt model was evaluated using the input variables listed in Table 4. A preliminary value of emissivity ϵ'_{10} , estimated from the literature, was used in conjunction with the other listed variables to develop a raw, simulated graybody surface temperature for each site location. Since the model simulates surface temperatures on integral hours (solar time), we computed simulated temperatures for 1500 and 1600 hours and linearly interpolated surface temperature values for 1527 hours solar time (the calculated solar time corresponding to a flight time of 1519 hours local standard mean time for Scipio Center on that date). As our analysis progressed however, we felt that it would be valuable to convert the previously simulated graybody temperatures to blackbody temperatures so that they could be directly compared with the observed and corrected blackbody temperatures T_b , listed in Tables 2 and 3. To avoid having to recompute all the simulated temperatures so derived, each graybody temperature was multiplied by $(\epsilon'^{-1}_{10})^{1/4} - 1$ to obtain the surface blackbody temperatures. These simulated blackbody temperatures T_c are listed in Tables 2 and 3.

When these simulated surface temperatures T_c are compared with the observed and corrected temperatures T_b for the two groups, we can see that, given a modest number of samples, the model appears to simulate very well the mean observed temperatures for each group as shown in Tables 3 and 4.

Errors associated with preparing a residual anomaly map

We have discussed above the errors associated with the corrected surface temperature data.

There are also errors that must be considered that are associated with the results simulated by the Outcalt model. This is illustrated by the fact that an overall one-to-one correlation between observed and simulated temperature data for the two terrain groups is not observed. The lack of correlation may be due in part to the data corrections discussed above, to lack of agreement of the model, or more probably, to errors in the input variables that were used by the model. Values for many of the input variables had to be estimated from the literature, since these had not been measured at the time of overflight. In an operational situation, however, they can be measured without much difficulty, and if valid input data are available, the model has been shown to simulate surface temperatures measured during the night by an *ir* scanner (Outcalt 1972b). Unfortunately, lack of such input data measurements, there is no way of estimating the spread of errors associated with the temperatures simulated in the present work. We recognize that the closeness of the mean simulated values to the mean observed values, i.e., 0.01 and 0.15 K, may be fortuitous, but the significant error about these values can be estimated based on our analysis of errors. However, we think that it is meaningful that a Student's *t*-Test applied to data from both groups shows that the hypothesis that the observed and simulated temperatures statistically come from the same population, can be accepted at the 95 percent confidence level. This implies that, at least on a statistical basis, the "normal" temperature values computed by the Outcalt model can be compared with the observed values to derive a residual anomaly map. It seems likely that as we have more observed values and simulate more surface temperatures that match the geographic positions of the observed data, the significant error in each group of values will be reduced. It also appears from our limited data that as the number of values from both the observed and simulated populations increases, the means of these populations will converge. Whether this will prove to be the case as more data are recorded and analyzed, and more simulations are run, has yet to be determined. Accordingly, we think that the methodology discussed shows promise and considerably greater field and modeling efforts are warranted.

SUMMARY AND CONCLUSIONS

Our results suggest a way of utilizing *ir* data as

a geophysical tool instead of a medium amenable only to photointerpretation, as is most commonly done. To do this, numerous complicating factors must be removed, i.e., the effects of surface emissivity, soil composition and soil moisture, topography, hydrology, and meteorological conditions, which can completely mask the temperature effects sought. The techniques developed herein appear to be useful in making the needed corrections to raw *ir* data, thus permitting the isolation of the true temperature anomalies sought. The steps required to accomplish this are as follows:

- 1) Reduce the effect of emissivity. If the emissivity ratio is constant, as appeared to be the case where the terrain had vegetative cover, equation (6) suggests that the signal ratio $(W_5/W_{10})^{1/3} = (\epsilon_5/\epsilon_{10})^{1/3}T_b$ can be used to determine the surface temperature.
- 2) Once the effects of emissivity have been removed from the signal, the corrections for meteorological, topographical, and ground surface conditions can be applied by simulating the "normal" surface temperatures that would be expected at each point on this terrain at the time of data collection. The residual between the observed surface temperatures and the temperatures simulated by the model can then be attributed to the anomalous conditions sought.

As an example of this, we observed that for the 18 points belonging to group 1, the mean of the simulated temperatures were essentially equal to the mean of the associated 18 observed temperature values, and that for the 16 group 2 values the difference was 0.1°C. Moreover, the simulated and the observed temperatures appear to come from the same population. This suggests to us that with a sufficiently large number of data samples recorded by airborne means, and taken within a nominal area of perhaps 500 m on a side, the means of the simulated temperatures can be made to converge on the means of the observed (and corrected) temperature values such that a true surface temperature contour map could be constructed. The construction of such a map implies that data values will be sufficiently far apart so that the usual photograph-like surface detail will be deliberately averaged out to produce a temperature map representing a large area. If all corrections have been made properly such that the significant error is small, and there is no abnormal heat flow in the area owing to near-surface geo-

logic processes, then the residual between the observed and simulated data will approach zero. Where temperature residuals are not zero and meaningful geologic patterns can be interpreted from the residual anomaly maps, the results of our methodology suggest a potential new technique for geophysical exploration.

ACKNOWLEDGMENTS

We wish to thank Dr. Arthur H. Lachenbruch, U.S. Geological Survey, Menlo Park, Ca., Dr. John Lewis, McGill University, and Dr. Sam I. Outcalt of the University of Michigan for the many helpful discussions that we had with them during the course of the research. Additionally, we would like to thank Dr. Jules D. Friedman, U.S. Geological Survey, Denver, Colo., Donald R. Wiesnet of the National Oceanic and Atmospheric Administration, Suitland, Md., and Dr. Paul Ebert of the University of California Lawrence Livermore Laboratory for reviewing our manuscript. This work was supported in part by the Energy Research and Development Administration contract W-7405-Eng-48 and by the National Oceanic and Atmospheric Administration under contract 4-35308.

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APPENDIX

spectral radiant emittance W_λ , measured in watts/m²/unit wavelength, is what is detected by an airborne ir scanning system. This equation expresses W_λ as follows:

$$W_\lambda = \epsilon_\lambda C_1 \lambda^{-5} [\exp(C_2/\lambda T) - 1]^{-1} \quad \text{A-1}$$

where ϵ_λ is the surface emissivity at a given wavelength λ , 3.7414×10^8 W/m² · μm⁴, 1.4388×10^4 μm · K, λ is wavelength in μm, and T is surface temperature in K.

Since the scanner senses not the entire range of radiated energy but only that energy contained in a narrow wavelength band for which the system is designed, equation (A-1) should be expressed as

$$W = \int_{\lambda_1}^{\lambda_2} \epsilon_\lambda C_1 [\exp(C_2/\lambda T) - 1]^{-1} \lambda^{-5} d\lambda \quad \text{A-2}$$

To simplify the mathematics, we will make the following substitution:

$$y = C_2/\lambda T, \quad \text{A-3}$$

therefore, it follows that $\lambda = C_2/Ty$. By differentiating,

$$\lambda^{-5} dy = -C_2^{-4} T^4 y^3 dy \quad \text{A-4}$$

Equation (A-2) can now be rewritten

$$W = \int_{y_2}^{y_1} \frac{\epsilon_\lambda C_1}{\exp(y) - 1} (-C_2^{-4} T^4 y^3) dy \quad \text{A-5}$$

$$W = -\epsilon_\lambda C_1 C_2^{-4} T^4 \int_{y_2}^{y_1} \left(\frac{y^3}{\exp(y) - 1} \right) dy \quad \text{A-6}$$

Multiplying the numerator and denominator of the bracketed term by $(\exp - y)$, we have

$$\frac{y^3 \exp(-y)}{1 - \exp(-y)}$$

and

$$W = -\epsilon_\lambda C_1 C_2^{-4} T^4$$

¹ For integrations over small wavelength intervals, ϵ_λ is considered constant; λ applies to the mean wavelength.

$$\int_{y_1}^{y_2} \left(\frac{y^3 \exp(-y)}{1 - \exp(-y)} \right) dy \quad \text{A-7}$$

The term

$$\frac{\exp(-y)}{1 - \exp(-y)}$$

can be expanded in a binomial series, as

$$\begin{aligned} \exp^{-y} \sum_{m=0}^{\infty} \exp(-my) \\ = \exp(-y)[\exp(-0) + \exp(-y) \\ + \exp(-2y) + \dots \exp(-my)] \end{aligned}$$

or

$$\begin{aligned} [\exp(-y) + \exp(-2y) + \\ \exp(-3y) + \dots \exp(-m+1)y] \end{aligned}$$

which can be expressed as

$$\sum_{m=1}^{\infty} \exp(-my)$$

Equation (A-7) can now be rewritten using the above expansion and eliminating the negative sign by reversing the limits of integration as follows:

$$\begin{aligned} W = \epsilon_\lambda C_1 C_2^{-4} T^4 \\ \int_{y_2}^{y_1} y^3 \left(\sum_{m=1}^{\infty} \exp(-my) \right) dy \quad \text{A-8} \end{aligned}$$

Integrating equation (A-8) by parts, using the form

$$\begin{aligned} \int y^3 \exp(-my) dy = \frac{y^3 \exp(-my)}{-m} \\ - \frac{3}{-m} \int y^2 \exp(-my) dy, \end{aligned}$$

we obtain

$$\begin{aligned} W = \epsilon_\lambda C_1 C_2^{-4} T^4 \\ \left[\sum_{m=1}^{\infty} \exp(-my_2) \left(\frac{y_2^3}{m} + \frac{3y_2^2}{m^2} \right. \right. \\ \left. \left. + \frac{6y_2}{m^3} + \frac{6}{m^4} \right) - \sum_{m=1}^{\infty} \exp(-my_1) \left(\frac{y_1^3}{m} \right. \right. \\ \left. \left. + \frac{3y_1^2}{m^2} + \frac{6y_1}{m^3} + \frac{6}{m^4} \right) \right] \quad \text{A-9} \end{aligned}$$

² Standard Mathematical Tables, Chemical Rubber Company, 12th Ed., 1962, p. 309, Integral #354.

Equation (A-9) is a complete expansion of Planck's equation. However, it is unwieldy to evaluate for our emissivity ratio analysis.

Accordingly, we have derived an approximation of equation (A-9) that considerably simplifies the mathematics. Differentiating equation (A-1) with respect to T and holding λ constant, we obtain

$$dW_\lambda = \epsilon_\lambda C_1 \lambda^{-5} \frac{(C_2/\lambda T^2) \exp(C_2/\lambda T)}{(\exp(C_2/\lambda T) - 1)^2} dT. \quad (A-10)$$

Dividing equation (A-10) by W_λ , we obtain

$$\frac{dW_\lambda}{W_\lambda} = \frac{(C_2/\lambda T^2) \exp(C_2/\lambda T)}{[\exp(C_2/\lambda T) - 1]} dT, \quad (A-11)$$

or

$$\frac{dW_\lambda}{W_\lambda} = (C_2/\lambda T) \left[\frac{\exp(C_2/\lambda T)}{\exp(C_2/\lambda T) - 1} \right] \frac{dT}{T}. \quad (A-12)$$

The bracketed term, by multiplying and dividing by $\exp(-C_2/\lambda T)$, can be rewritten

$$\left[\frac{1}{1 - \exp(-C_2/\lambda T)} \right],$$

which is of the form $1/(1 - x)$, where $x = \exp(-C_2/\lambda T)$. It therefore can be expressed as a binomial expansion of the form

$$\frac{1}{1 - x} = 1 + x + x^2 + \dots + x^m = \sum_{m=0}^{\infty} x^m,$$

as long as $\exp(-C_2/\lambda T) < 1$. This will be the case for the wavelengths of interest in this work. With the nominal temperature of 15°C or 288 K (recorded at the time of the field study) and for data recorded in wavelengths less than 15 μm , the bracketed term is approximately equal to unity; this is because the second and higher order terms of the expansion are much less than 1.

Accordingly, Equation (A-12) can be expressed as

$$\frac{dW_\lambda}{W_\lambda} = \frac{C_2}{\lambda T} \frac{dT}{T}. \quad (A-13)$$

Since by definition, $dW_\lambda/W_\lambda \equiv d \ln W_\lambda$, and $dT/T \equiv d \ln T$, for small temperature excursions less than ± 5 K from T_0 , where $T = T_0(1 \pm \Delta T/T_0) \approx T_0$, integrating equation (A-13) becomes:

$$\ln W_\lambda \Big|_{W_\lambda(T_0)}^{W_\lambda(T)} = -\frac{C_2}{\lambda} \frac{1}{T} \Big|_{T_0}^T, \quad (A-14)$$

$$\begin{aligned} \ln \frac{W_\lambda(T)}{W_\lambda(T_0)} &= -\frac{C_2}{\lambda} \left(\frac{1}{T} - \frac{1}{T_0} \right) \\ &= -\frac{C_2}{\lambda T_0} \left(\frac{T_0}{T} - 1 \right), \end{aligned} \quad (A-15)$$

whence

$$\begin{aligned} W_\lambda(T) &= W_\lambda(T_0) \exp \left(\frac{T_0}{T} - 1 \right) \cdot \left(-\frac{C_2}{\lambda T_0} \right) \\ &= W_\lambda(T_0) \left[\exp \left(\frac{T_0}{T} - 1 \right) \right]^{-C_2/\lambda T_0} \end{aligned} \quad (A-16)$$

Since T is near T_0 , $(T_0/T - 1)$ is small, and may expand the exponential obtaining

$$\begin{aligned} W_\lambda(T) &= W_\lambda(T_0) \left[1 + \left(\frac{T_0}{T} - 1 \right) \right]^{-C_2/\lambda T_0}, \quad (A-17) \\ &= W_\lambda(T_0) \left(\frac{T_0}{T} \right)^{-C_2/\lambda T_0} \\ &= W_\lambda(T_0) \left[\frac{T}{T_0} \right]^{C_2/\lambda T_0}, \quad (A-18) \\ &= \frac{W_\lambda(T_0)}{T_0^{(C_2/\lambda T_0)}} \cdot T^{(C_2/\lambda T_0)}, \quad (A-19) \\ &= C_\lambda(T_0) T^{(C_2/\lambda T_0)} \propto \epsilon_\lambda T^{(C_2/\lambda T_0)}. \end{aligned} \quad (A-20)$$

For $C_2 = 14,388 \mu\text{m} \cdot \text{K}$ and a nominal value $T_0 = 288$ K, $C_2/T = 49.96$, equation (21) can closely be approximated by

$$W_\lambda \propto \epsilon_\lambda T^{50/\lambda}. \quad (A-21)$$

We have computed the difference for a change W_λ from $T = 288$ K to $T = 289$ K by evaluating equation (A-9), the complete expansion of Planck's equation, and Equation (A-22). Comparing the results, we found that for the wavelengths of interest to us (i.e., between 5 μm and 15 μm), equation (A-22) introduces an error of more than ± 0.04 percent. From the above, we have derived a convenient expression relating radiant emittance sensed by a typical ir scanner to the absolute surface temperature raised to the power of $50/\lambda$. This power law thermal mechanism becomes an extremely useful mechanism for enhancing either the effects of emissivity or of surface temperature.